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# The RAMSES code and related techniques

## 3. Gravity solvers

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# Outline

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- The Vlasov-Poisson equation
- The Particle-Mesh method in a nut shell
- Symplectic Integrators
- Mass assignment and force interpolation schemes
- FFT-based Poisson solver
- P3M, AP3M, TPM, and related high-resolution schemes
- AMR schemes
- Relaxation Poisson solvers
- Multigrid Poisson solvers and AMR specifics

# The Vlasov-Poisson equation

Collisionless limit of the Boltzmann equation:

$$\frac{df}{dt} = \frac{\partial}{\partial t} f(\mathbf{x}, \mathbf{p}, t) + \frac{\mathbf{p}}{ma^2} \frac{\partial}{\partial \mathbf{x}} f(\mathbf{x}, \mathbf{p}, t) - m \nabla_{\mathbf{x}} \cdot \Phi(\mathbf{x}) \frac{\partial}{\partial \mathbf{p}} f(\mathbf{x}, \mathbf{p}, t) = 0$$

Liouville theorem: number of particles is conserved in phase-space

Gravitational acceleration is given by the Poisson equation

$$\Delta \Phi(\mathbf{x}) = \frac{4\pi Gm}{a} \left( \int f(\mathbf{x}, \mathbf{p}, t) d^3 \mathbf{p} - \bar{n} \right),$$

3 solution strategies:

- pure fluid on a 6D grid
- pure N body using direct or Tree force computations
- mixture of the 2: the Particle-Mesh method

# The Particle-In-Cell method

N body integrator coupled to a grid-based Poisson solver

- Compute the mass density field on the grid from the particle distribution
- Solve for the Poisson equation on the grid
- Interpolate the force back to the particle position

$$\frac{d\mathbf{x}_p}{dt} = \mathbf{v}_p \quad \text{and} \quad \frac{d\mathbf{v}_p}{dt} = -\nabla_x \phi$$

Harlow, F.H., “A Machine Calculation Method for Hydrodynamics Problems”, Los Alamos Scientific Laboratory report LAMS-1956 (1955)

Hockney, R.W., Eastwood, J.W., “Computer Simulation Using Particles”, CRC Press (1985)

The PIC or PM (Particle-Mesh) scheme has been applied to:

- Hydrodynamics (compressible, incompressible, MHD)
- Plasma physics
- Self-gravitating systems

# Symplectic Time Integrator

Phase space:  $q = \mathbf{x}_p \quad p = \mathbf{v}_p \quad \dot{q} = p$

Hamiltonian:  $\mathcal{H}(q, p) = p^2/2 + \Phi(q) \quad \dot{p} = -\frac{\partial\Phi}{\partial q}$

The exact solution of an Hamiltonian system is area-preserving in phase-space (incompressible fluid in phase space).

The area is a Poincare invariant.

$$\mathbf{z} = (q, p) \quad \mathbf{f}(\mathbf{z}) = (p, -\partial\Phi/\partial q) \quad \dot{\mathbf{z}} = \mathbf{f}(\mathbf{z})$$

Define the orbits as:  $\mathbf{z}(t) = \mathcal{F}(\mathbf{z}_0)$

The area in phase-space is preserved if:  $\det \frac{\partial\mathcal{F}}{\partial\mathbf{z}} = 1$

# Classical First Order Time Integrators

- Explicit Euler:  $\mathbf{z}^{n+1} = \mathbf{z}^n + \Delta t \mathbf{f}(\mathbf{z}^n)$   $\det \frac{\partial \mathcal{F}}{\partial \mathbf{z}} = 1 + \Delta t^2 \frac{\partial^2 \Phi}{\partial q^2}$
- Implicit Euler:  $\mathbf{z}^{n+1} = \mathbf{z}^n + \Delta t \mathbf{f}(\mathbf{z}^{n+1})$   $\det \frac{\partial \mathcal{F}}{\partial \mathbf{z}} = \frac{1}{1 + \Delta t^2 \frac{\partial^2 \Phi}{\partial q^2}}$
- Symplectic Euler:  $\mathbf{z}^{n+1} = \mathbf{z}^n + \Delta t \mathbf{f}(q^n, p^{n+1})$   $\det \frac{\partial \mathcal{F}}{\partial \mathbf{z}} = 1$

# Second Order Symplectic Time Integrator

Kick-Drift-Kick algorithm (see also Leap Frog)

$$p^{n+1/2} = p^n - 0.5\Delta t \left( \frac{\partial \Phi}{\partial q} \right)^n$$

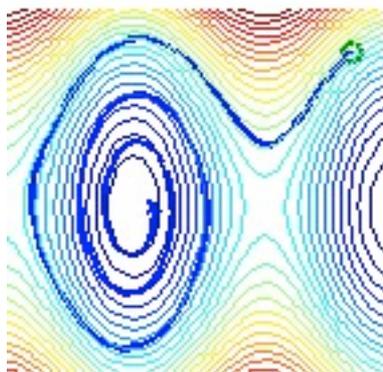
$$q^{n+1} = q^n + \Delta t p^{n+1/2}$$

$$p^{n+1} = p^{n+1/2} - 0.5\Delta t \left( \frac{\partial \Phi}{\partial q} \right)^{n+1}$$

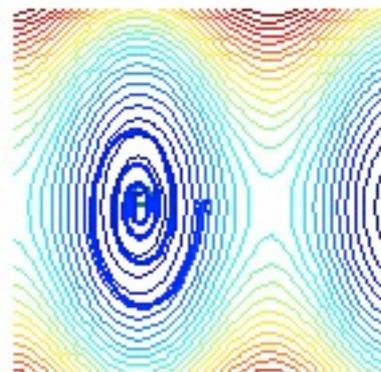
Exercise: check that the scheme is area-preserving.

# Why use a symplectic integrator ?

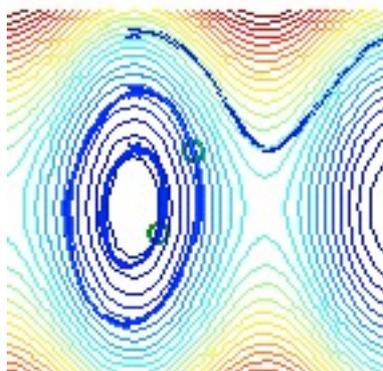
Explicit Euler



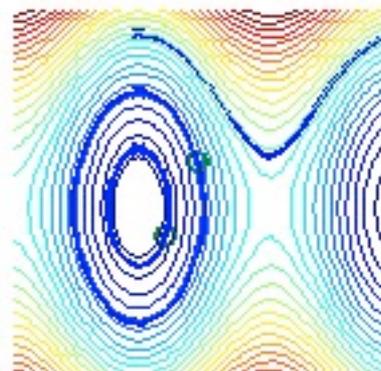
Implicit Euler



Symplectic Euler



Implicit Midpoint



# Charge assignment schemes

Assign to each particle a “shape”

Nearest Grid Point (NGP): 
$$S(x) = \frac{1}{\Delta x} \delta\left(\frac{x}{\Delta x}\right)$$

Cloud-In-Cell (CIC): 
$$S(x) = \frac{1}{\Delta x} \Pi\left(\frac{x}{\Delta x}\right)$$

Triangular Shape Cloud (TSC): 
$$S(x) = \frac{1}{\Delta x} \Delta\left(\frac{x}{\Delta x}\right)$$

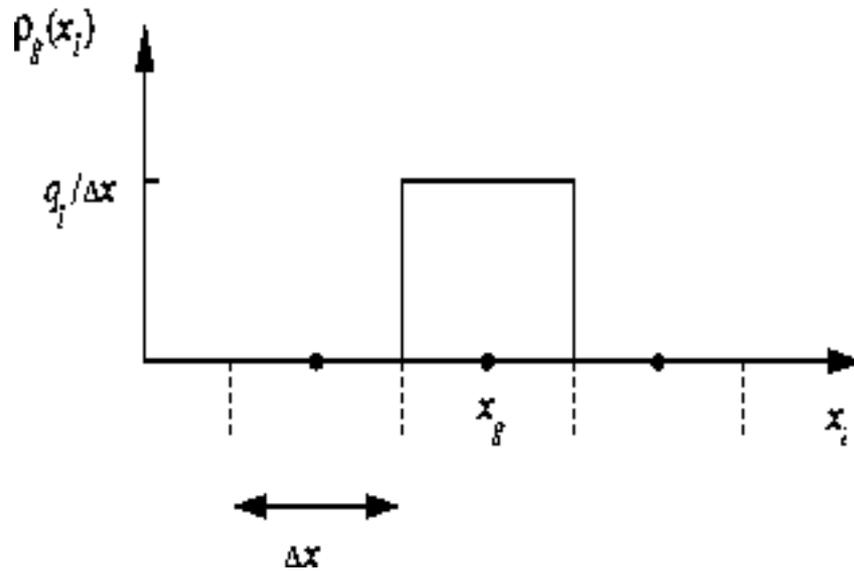
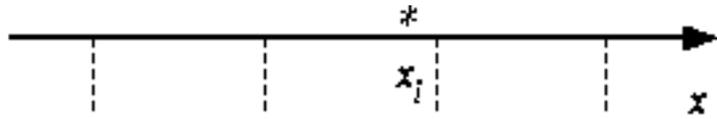
The contribution of each particle to the charge in the cell is:

$$W^\rho(x_p - x_i) = \int_{x_i - \Delta x/2}^{x_i + \Delta x/2} S(x_p - x) dx$$

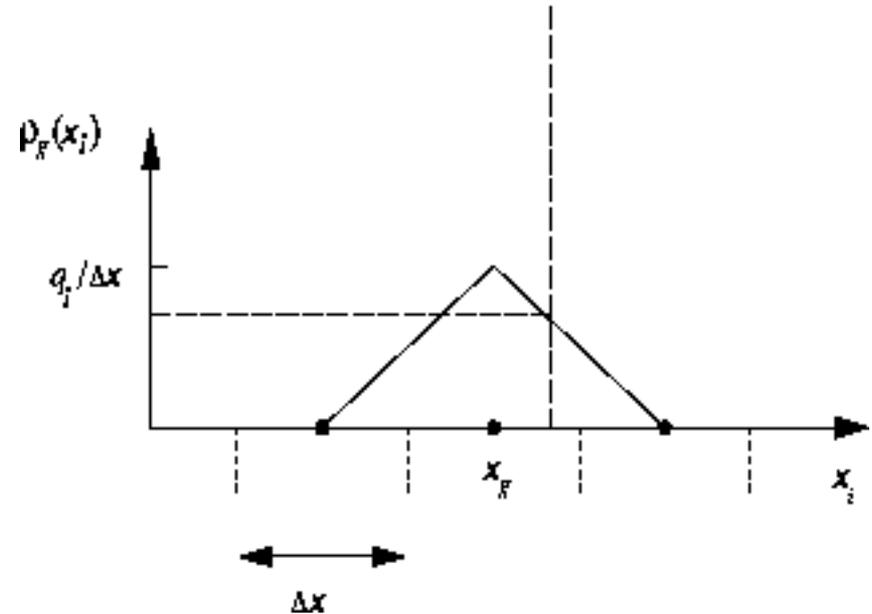
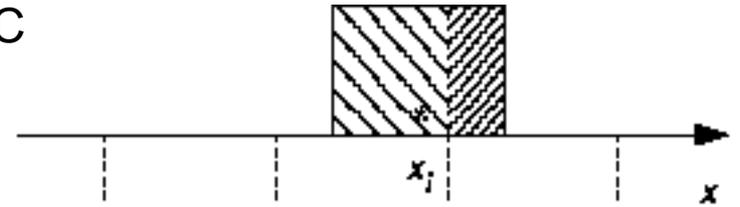
The total charge in the cell is: 
$$\rho_i = \frac{1}{\Delta x} \sum_{p=1}^{N_p} m_p W^\rho(x_p - x_i)$$

# A Hierarchy of Shapes

NGP



CIC



From NGP to CIC,  
From CIC to TSC,  
etc...

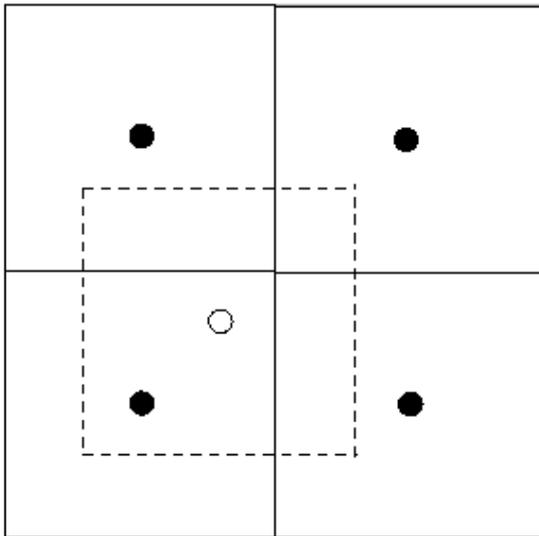
# Force interpolation

Use another interpolation scheme to get the mesh force at particle positions.

$$F(x_p) = m_p \sum W^F(x_p - x_i) F_i$$

Momentum conservation is enforced if:

- 2 interacting particles see equal but opposite forces ([exercise !](#))
- no self-forces



“Cloud-In-Cell” interpolation

Poisson equation

$$\Delta_{ij} \Phi_j = \rho_i$$

Gradient of the potential

$$F_i = -\nabla_{ij} \Phi_j$$

Self-force for particle p:

$$\partial F(x_p) = -m_p^2 \sum_i \sum_j W^F(x_p - x_i) \left( \nabla \cdot \Delta^{-1} \right)_{ij} W^\rho(x_p - x_j)$$

Self-force is zero if operator is antisymmetric and force and mass assignment schemes are equal.

# Force solver using Fourier analysis

Use of Fast Fourier Transform to solve for the Poisson equation

Poor's man Poisson solver:

$$\begin{aligned}\frac{\partial^2 \Phi}{\partial x^2} = \rho & & -k^2 \tilde{\Phi}(k) = \tilde{\rho}(k) & & \tilde{G}(k) = -\frac{1}{k^2} \\ \frac{\partial \Phi}{\partial x} = -F & & -ik \tilde{\Phi}(k) = \tilde{F}(k) & & \tilde{D}(k) = -ik\end{aligned}$$

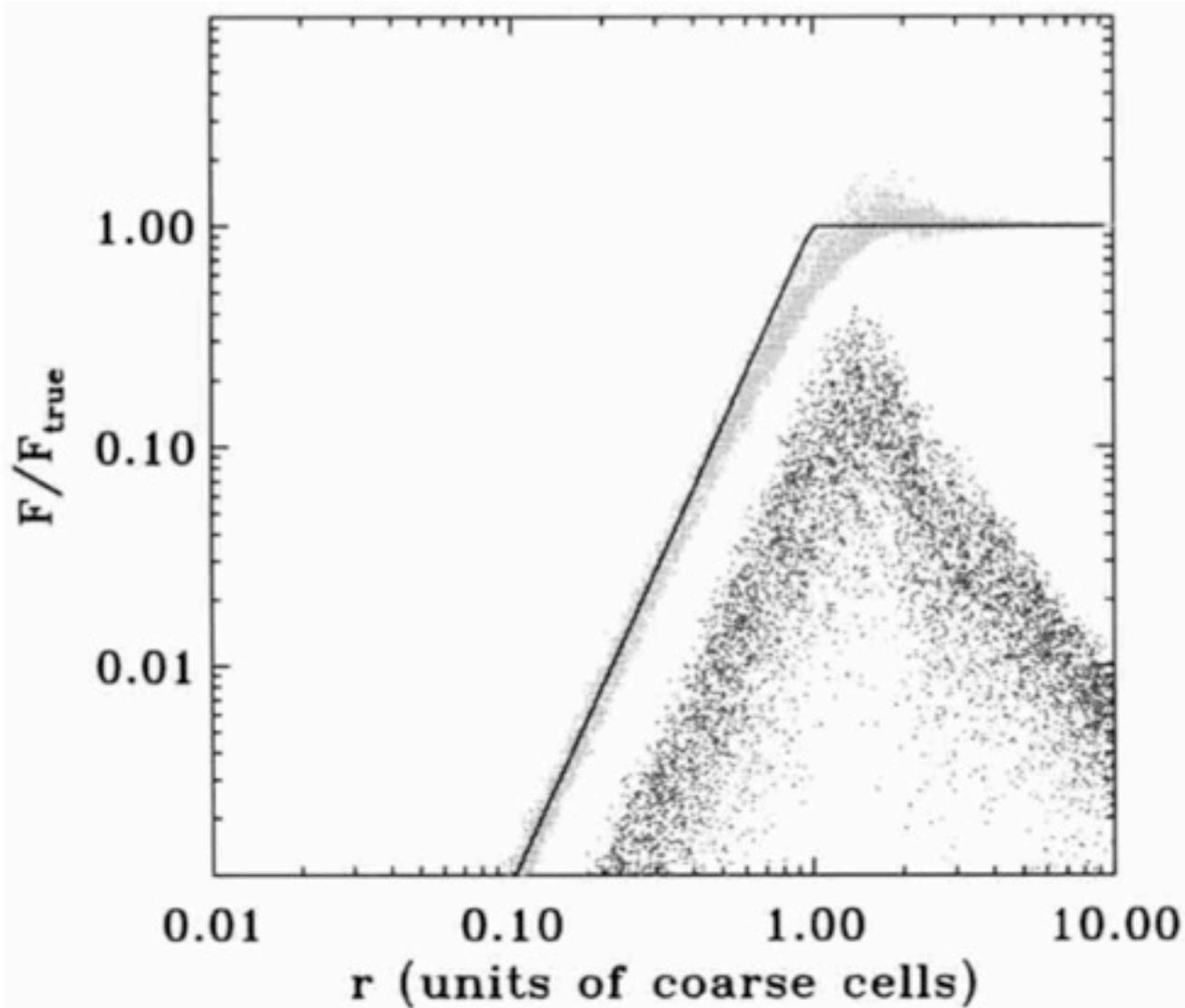
Using finite difference approximations:

$$\begin{aligned}\Phi_{i+1} - 2\Phi_i + \Phi_{i-1} = \rho_i \Delta_x^2 & & \tilde{G}(k) = -\frac{\Delta x^2 / 4}{\sin^2(\frac{k\Delta x}{2})} \\ -(\Phi_{i+1} - \Phi_{i-1}) = F_i \Delta_x & & \tilde{D}(k) = -i \frac{\sin(k\Delta x)}{\Delta x}\end{aligned}$$

Final force is given by:

$$\tilde{F}(k) = -\frac{m_p^2}{\Delta x^2} \tilde{W}^F(k) \tilde{D}(k) \tilde{G}(k) \tilde{W}^P(k) \tilde{n}(k)$$

# PM force accuracy



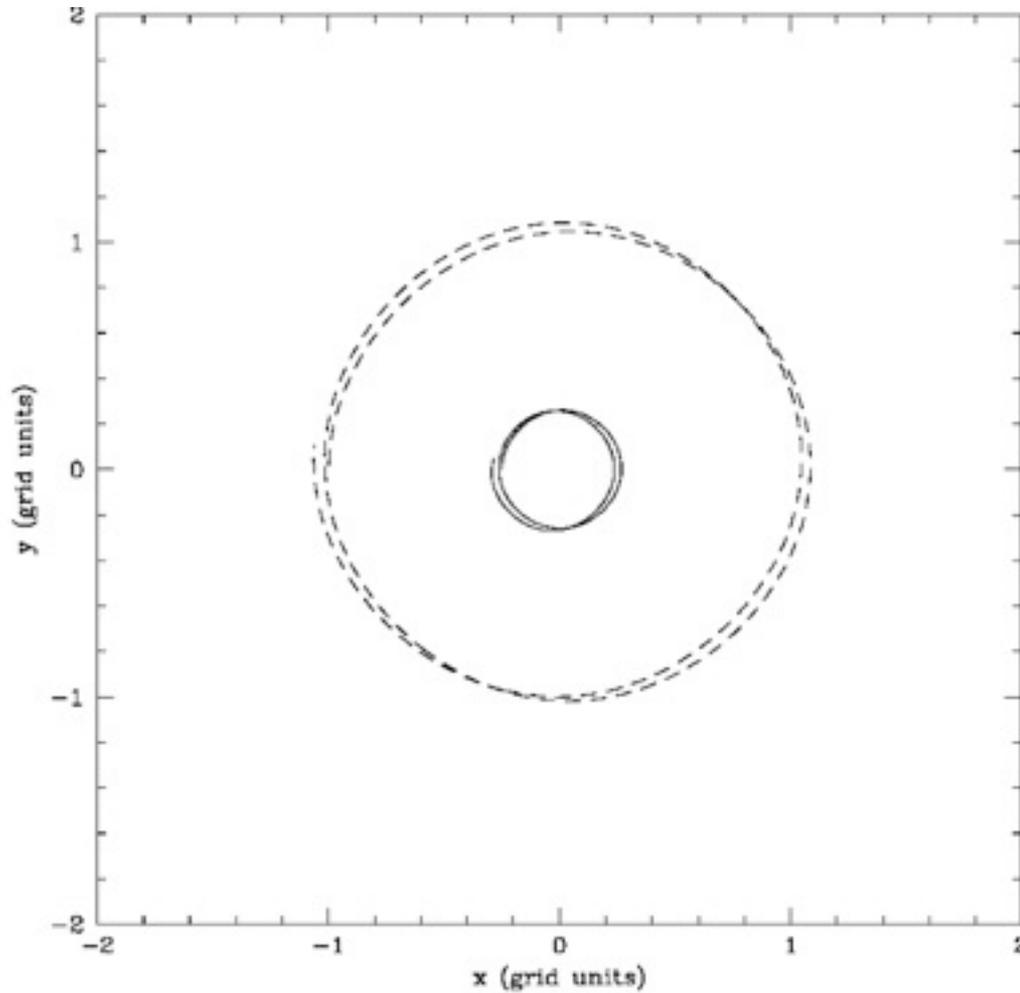
CIC

7-point  
Laplacian

2 points  
gradient

$\text{CIC}^{-1}$

# Force + integrator accuracy



examples of  
particle's  
trajectory

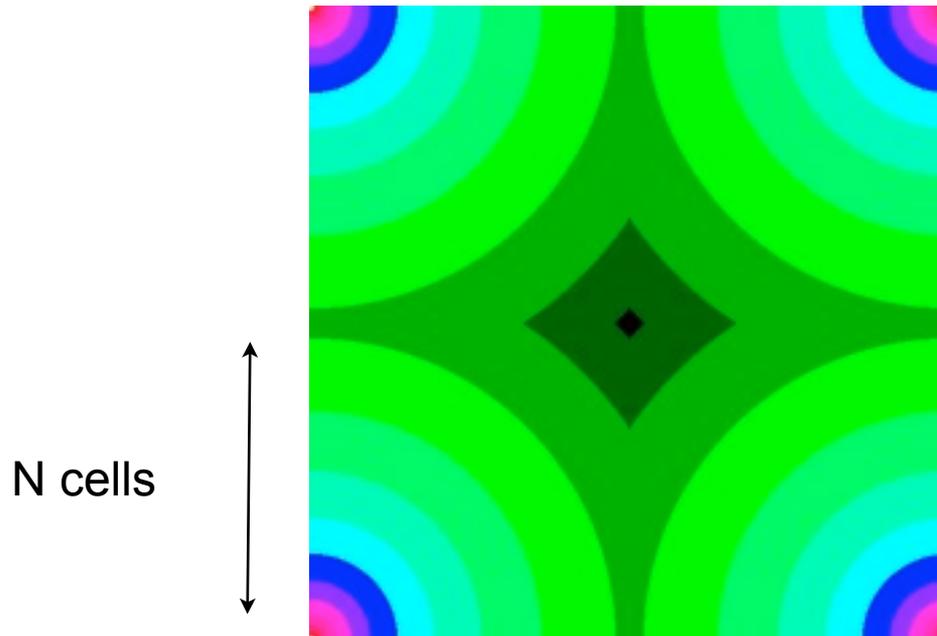
# What about isolated boundary conditions ?

Fourier analysis (and FFT) is more straightforwardly performed in a periodic system. How can one still use Fourier analysis in an isolated system ?

The “zero padding” trick (Hockney & Eatwood 81) for a  $N \times N$  grid.

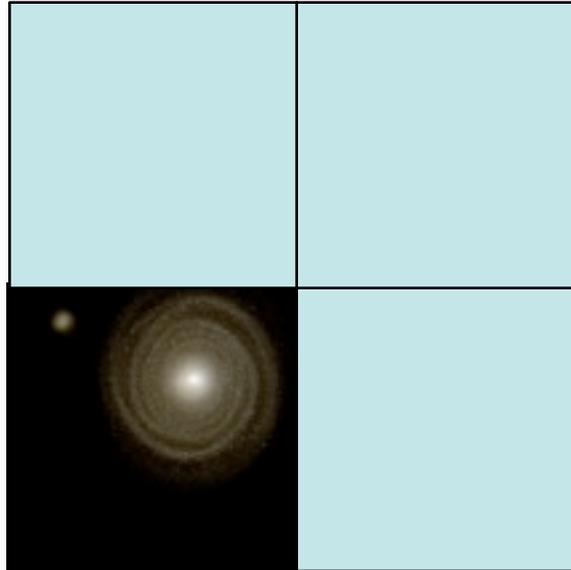
1- Compute the Green's function in real space, in the  $N \times N$  grid.

2- Duplicate and mirror the Green's function in the 3 “ghost” quadrants, so that the Green's function is now periodic in the  $2N \times 2N$  grid.



# What about isolated boundary conditions ?

3- Set the density field to zero in the 3 ghost quadrants.



4- Perform the convolution using Fourier analysis on the  $2N \times 2N$  grid.

Very simple and flexible ! It is quite expensive (8 times longer) !

Alternative method:

James, R.A., "The solution of Poisson's equation for isolated source distribution ", JCP, **25**, 71, (1977)

# Towards high-resolution: the P3M scheme

Decompose the total force in 2 components:

$$F_{hires}(r) = F_{PM}(r, a) + F_{PP}(r, \epsilon)$$

The PP force is a short-range interaction for  $r < r_c = 2-3a$ :

$$F_{PP}(r, \epsilon) = \frac{1}{r^2 + \epsilon^2} - F_{ref}(r, a)$$

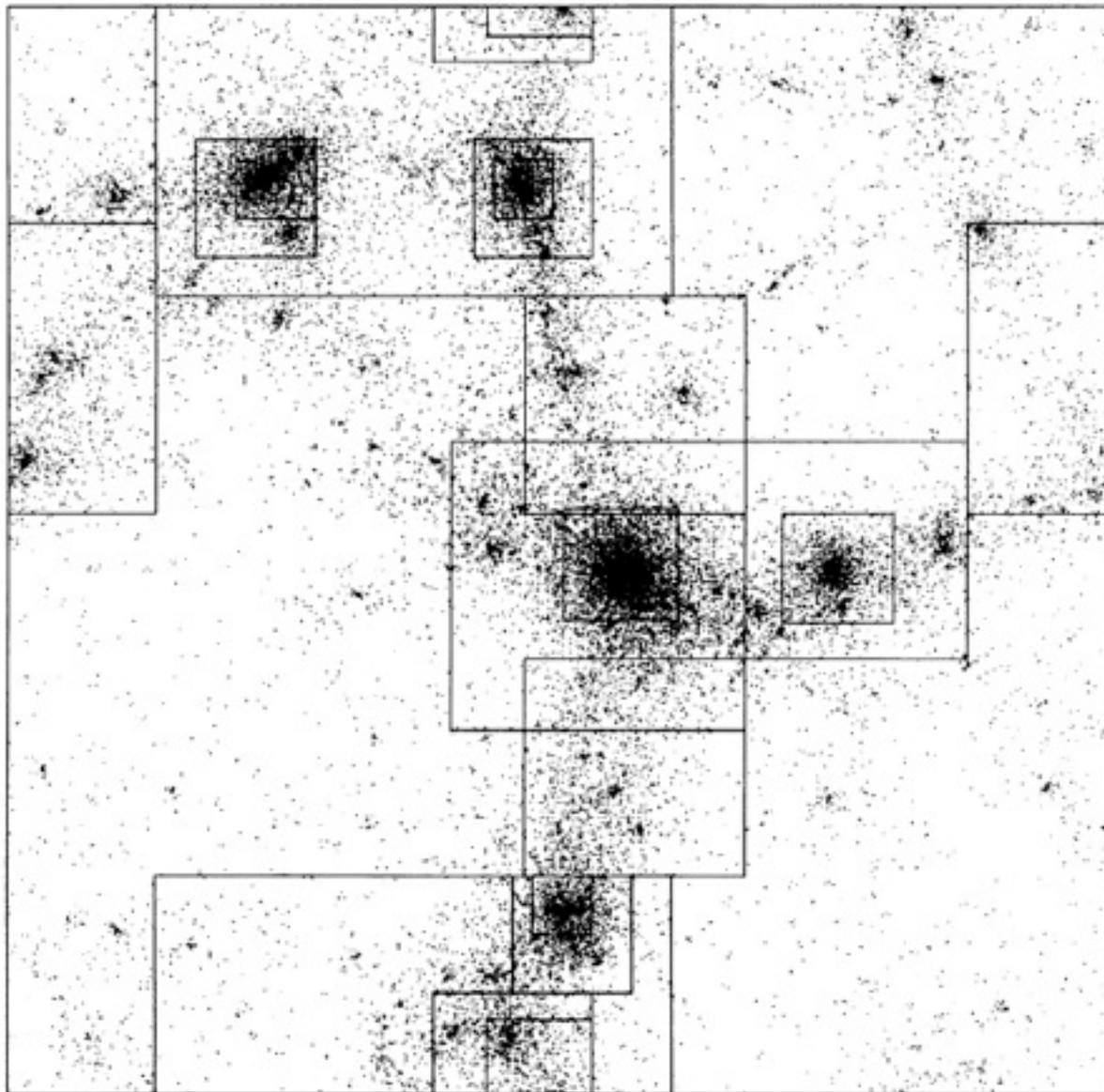
The PP force is computed with direct  $N^2$  summation, for all particles within the cut-off radius.

Problem: for highly clustered configuration, the PP force dominates the CPU time in the P3M scheme.

Solution: adaptive grids: the AP3M scheme.

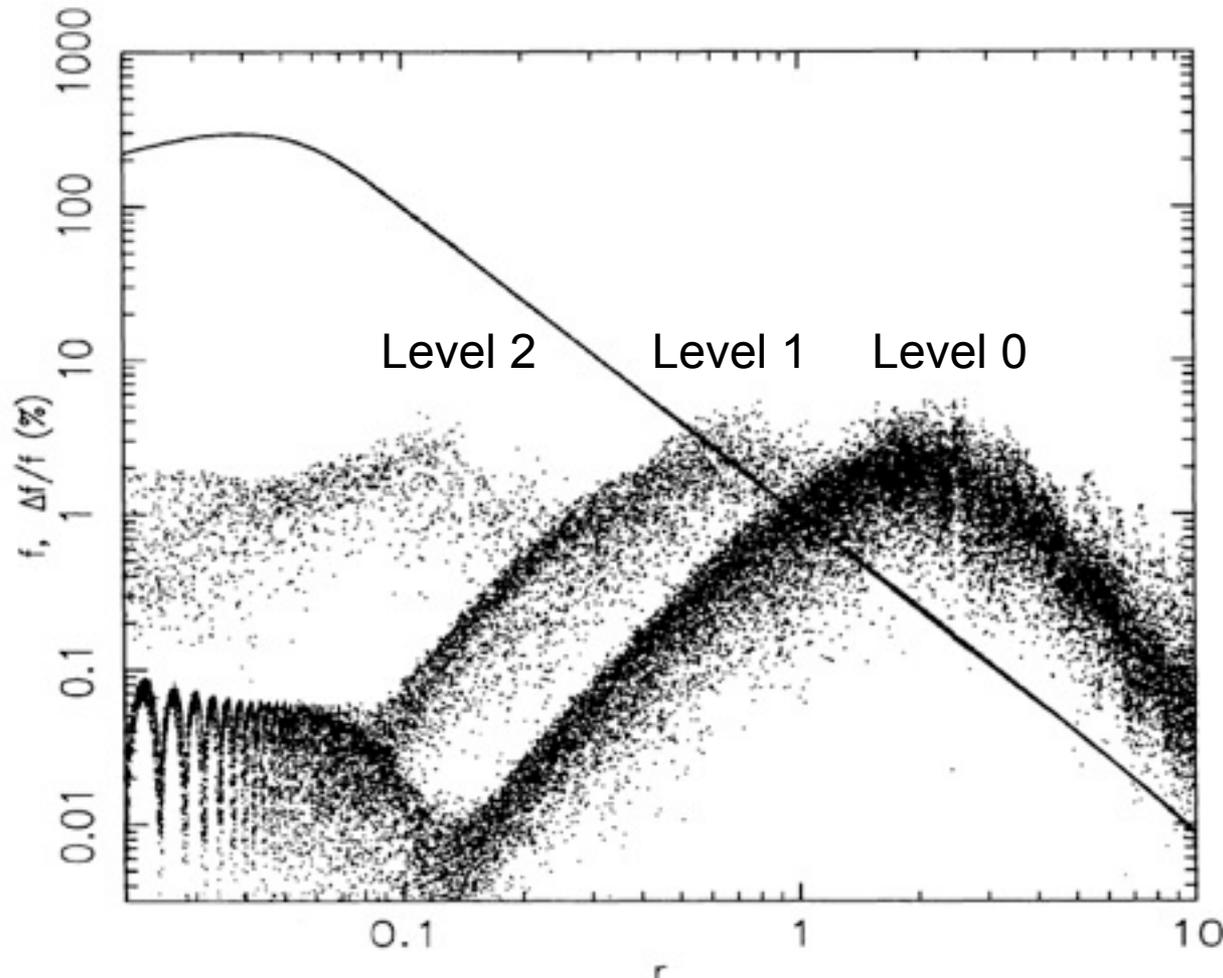
Couchman, H.M.P., “Mesh refined P3M: A fast adaptive N-body algorithm”, *ApJ*, **368**, 23, (1991)

# Adaptive grids: the AP3M scheme



# Adaptive grids: the AP3M scheme

Decompose the force recursively in smaller and smaller scale components. Only the coarse grid PM force is long-range. Requires force shaping at each level, and the use of an isolated boundary conditions Poisson solver.

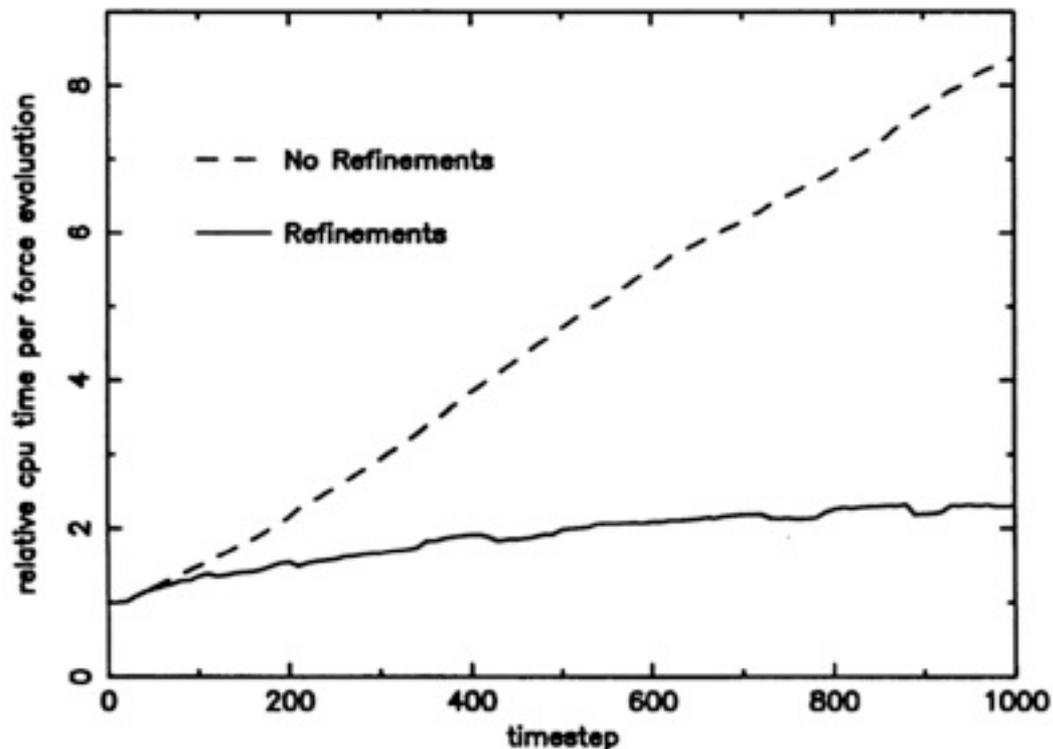


# High-resolution with constant force softening

Class of high-resolution N-body solvers based on the Particle In Cell method, with a constant (small scale) force softening:

- P3M: Hockney & Eastwood 81
- AP3M: Couchman 91
- TreePM: Bode & Ostriker 2003

Speed-up of the P3M scheme  
using adaptive grids  
(from Couchman et al. 95)



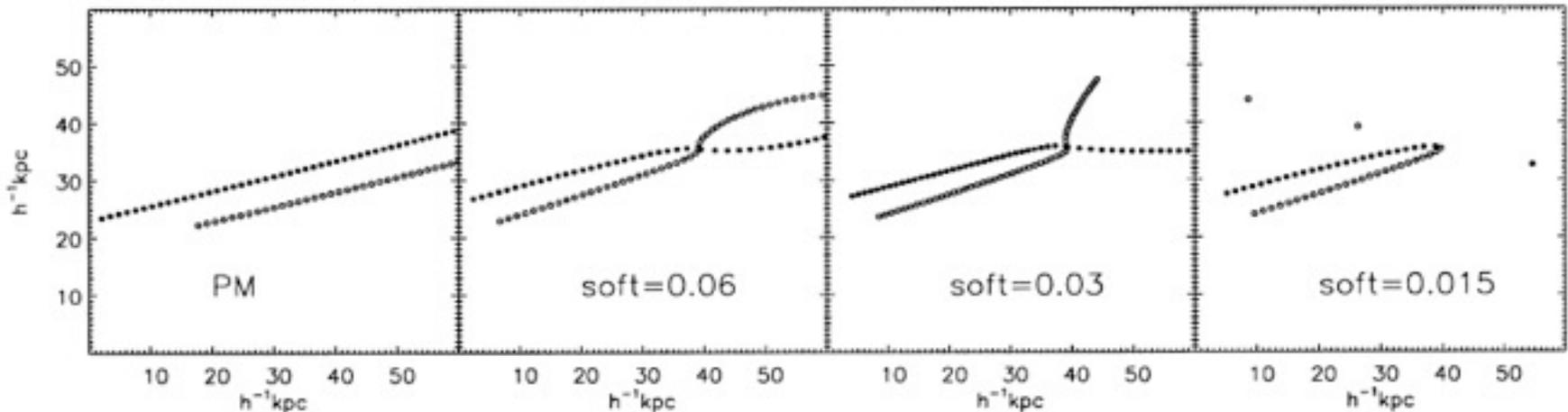
# High-resolution with constant force softening

Problem: how do we choose the force softening length ?

Usually, one has  $\epsilon = \frac{1}{10} \Delta x$  down to  $\frac{1}{100} \Delta x$   
where  $\Delta x$  is the mean inter-particle distance.

Pros: constant softening length ensures that the integration is symplectic, which results in good energy conservation

Cons: smoothing length is usually too small in low density regions and too large in high density regions.



Examples of 2-body scattering in low density region (from Knebe et al. 2000)

# High-resolution with constant force softening

Particle discreteness effects show up quite dramatically in Warm Dark Matter simulations (from Wang & White, MNRAS, 2007)

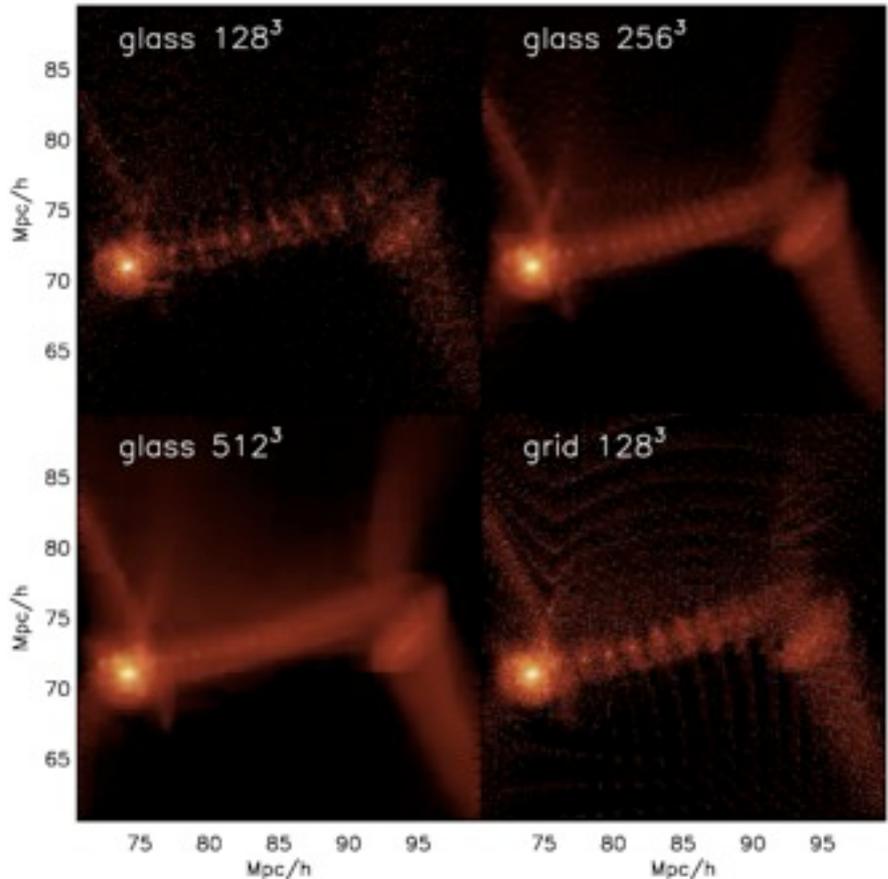
Very slow convergence  $N^{-1/3}$

These effects can be neglected if

$$\epsilon \approx 2d$$

$d$ : *local* inter-particle spacing

Adaptive force softening ?



Splinter, R.J., Melott, A.L., Shandarin, S.F., Suto, Y., “Fundamental Discreteness Limitations of Cosmological N-Body Clustering Simulations”, ApJ, **497**, 38, (1998)

Romeo, A.B., Agertz, O., Moore, B., Stadel, J., “Discreteness Effects in  $\Lambda$ CDM Simulations: A Wavelet-Statistical View”, ApJ, **686**, 1, (2008)

# PM with Adaptive Mesh Refinement

Idea: use the PM algorithm on a set of adaptively refined grid. Similar to the AP3M algorithm, without the short range PP part.

Method: each cell is recursively refined if the number of particle per cell exceed some threshold (around 10 particles).

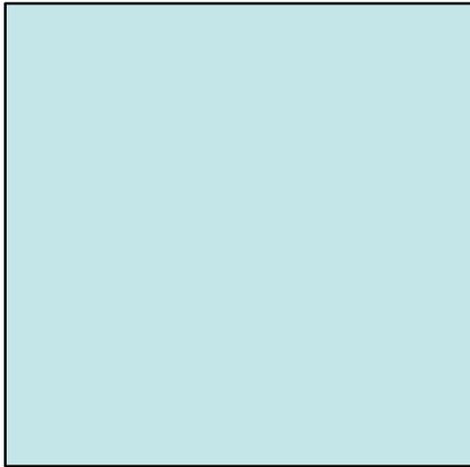
2 different methods for the Poisson solver:

- **the PANDORA code**: Villumsen, J.W., “A New Hierarchical PM Code for Very Large Scale Cosmological N-body Simulations”, ApJS, **71**, 407, (1989)
- **the ART code**: Kravtsov, A.V., Klypin, A.A., Khokhlov, A.M., “ART: a new high-resolution N-body code for cosmological simulations”, ApJS, **111**, 73, (1997)

A lot of different codes: ENZO (AP3M), ART, RAMSES (ART), PANDORA, MLAPM (ART), FLASH (unclear), CHARM (ART)...

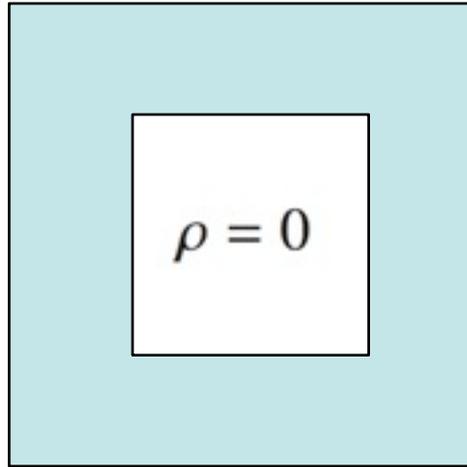
# Poisson solver with AMR: the PANDORA scheme

Isolated BC



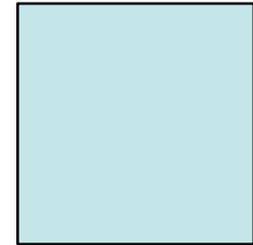
Step 1: compute the force on the coarse level due to *all* particles and interpolate to coarse-level particles

Isolated BC



Step 2: compute the force on the coarse level due to coarse-level particles and interpolate to fine-level particles

Isolated BC

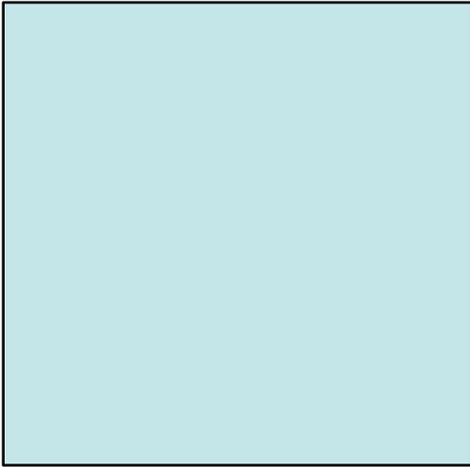


Step 3: compute the force on the fine level due to fine-level particles and interpolate to fine-level particles

**Only for rectangular grids !**

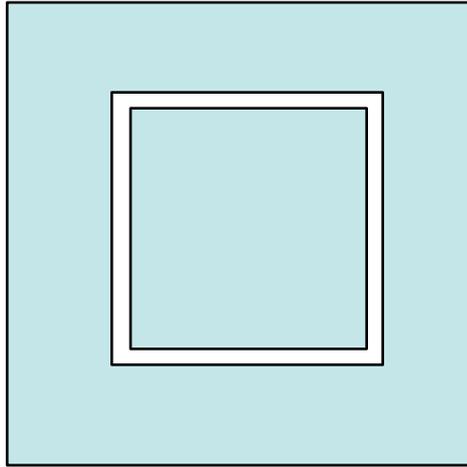
# Poisson solver with AMR: the ART scheme

Periodic BC

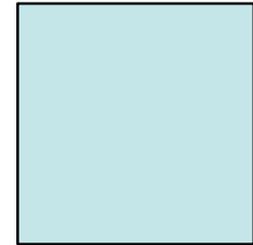


Step 1: compute the force on the coarse level due to *all* particles and interpolate to coarse-level particles

Dirichlet BC



Step 2: interpolate coarse-level potential at the coarse-fine boundary



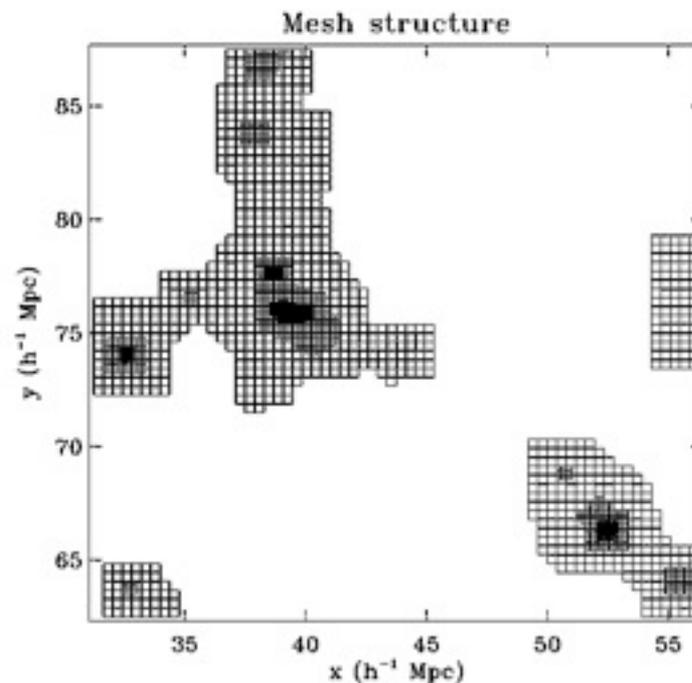
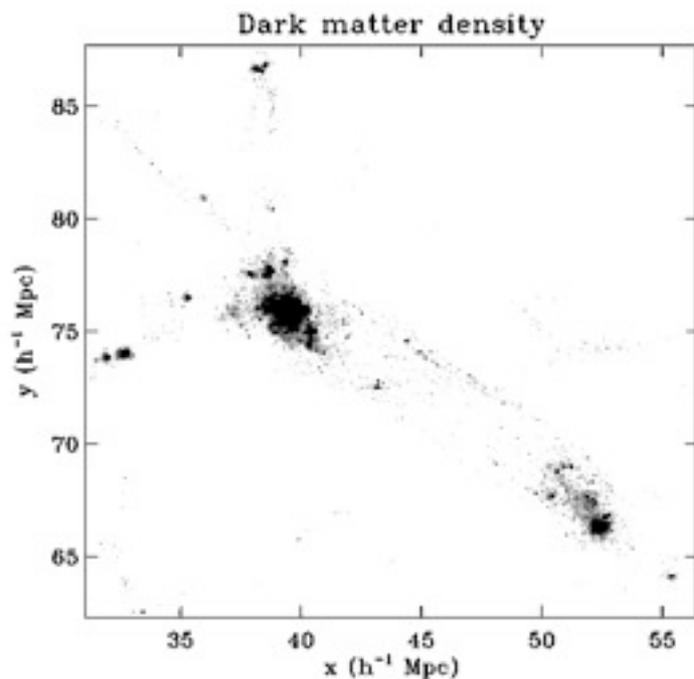
Step 3: compute the force on the fine level due to fine-level particles and interpolate to fine-level particles

**Valid for any grid geometry !**

# PM with Adaptive Mesh Refinement

At each grid level, the force softening is equal to the local grid size.

For pure dark matter simulations, using a quasi-Lagrangian strategy, the particle shot noise is kept roughly constant.



# ART: a one-way interface Poisson solver

Fine-level particles are temporarily passed to the coarse-level list to compute  $\rho_c$ .

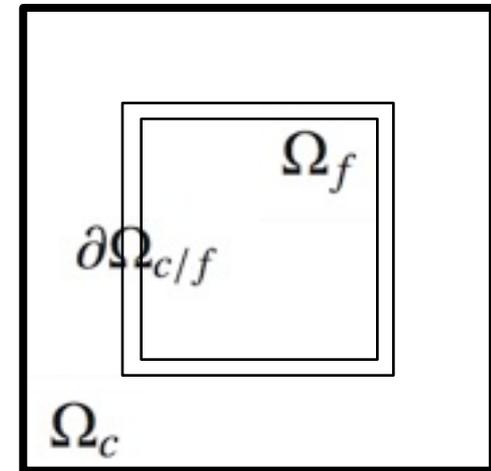
Fine-to-coarse information is mediated by particles only.

Solve  $\Delta\Phi_c = \rho_c$  on  $\Omega_c$

Interpolate BC on  $\partial\Omega_{c/f}$

Solve  $\Delta\Phi_f = \rho_f$  on  $\Omega_f$

Coarse-to-fine information is propagated by the potential through Dirichlet BCs.



A variant proposed by Miniati & Colella, JCP, 2007: in each coarse-level cell, the center-of-mass is computed and its contribution to  $\rho_c$  is added with CIC interpolation. Useful for self-gravity and gas dynamics.

# A two-way interface Poisson solver for AMR

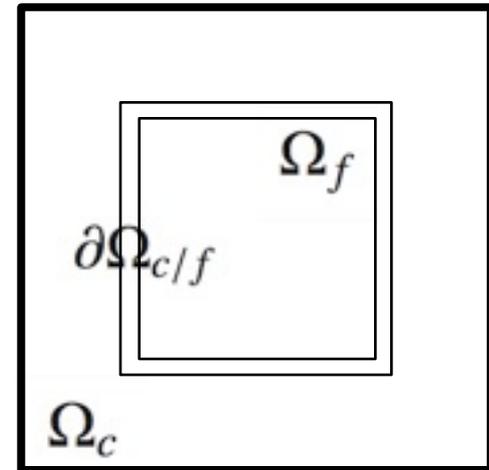
Proposed by Colella and co-workers for incompressible fluid dynamics and gravity solvers for compressible fluids.

Solve  $\Delta\Phi_c = \rho_c$  on  $\Omega_c - \Omega_f$

Impose Dirichlet and Neuman potential matching at  $\partial\Omega_{c/f}$

Solve  $\Delta\Phi_f = \rho_f$  on  $\Omega_f$

Iterate until global convergence



Applied to the Particle Mesh scheme: Miniati, F., Colella, P., “Block structured adaptive mesh and time refinement for hybrid, hyperbolic + N-body systems”, JCP, **227**, 400, (2007)

In case of adaptive time steps, revert to one-way coupling.

# Relaxation solvers for the Poisson equation

Solve the linear system  $\Delta_{ij}\Phi_j = \rho_i$  with arbitrary mesh geometry.

Simplest scheme: the Jacobi or Gauss-Seidel method (in 2D).

$$\phi_{i,j}^{n+1} = \frac{1}{4} (\phi_{i+1,j}^n + \phi_{i-1,j}^n + \phi_{i,j+1}^n + \phi_{i,j-1}^n) - \frac{1}{4}\rho_{i,j}$$

Converge very slowly for long wavelength and large grids.

Very sensitive to the initial guess.

Faster convergence is obtained for Gauss-Seidel “over-relaxation” method with red-black ordering.

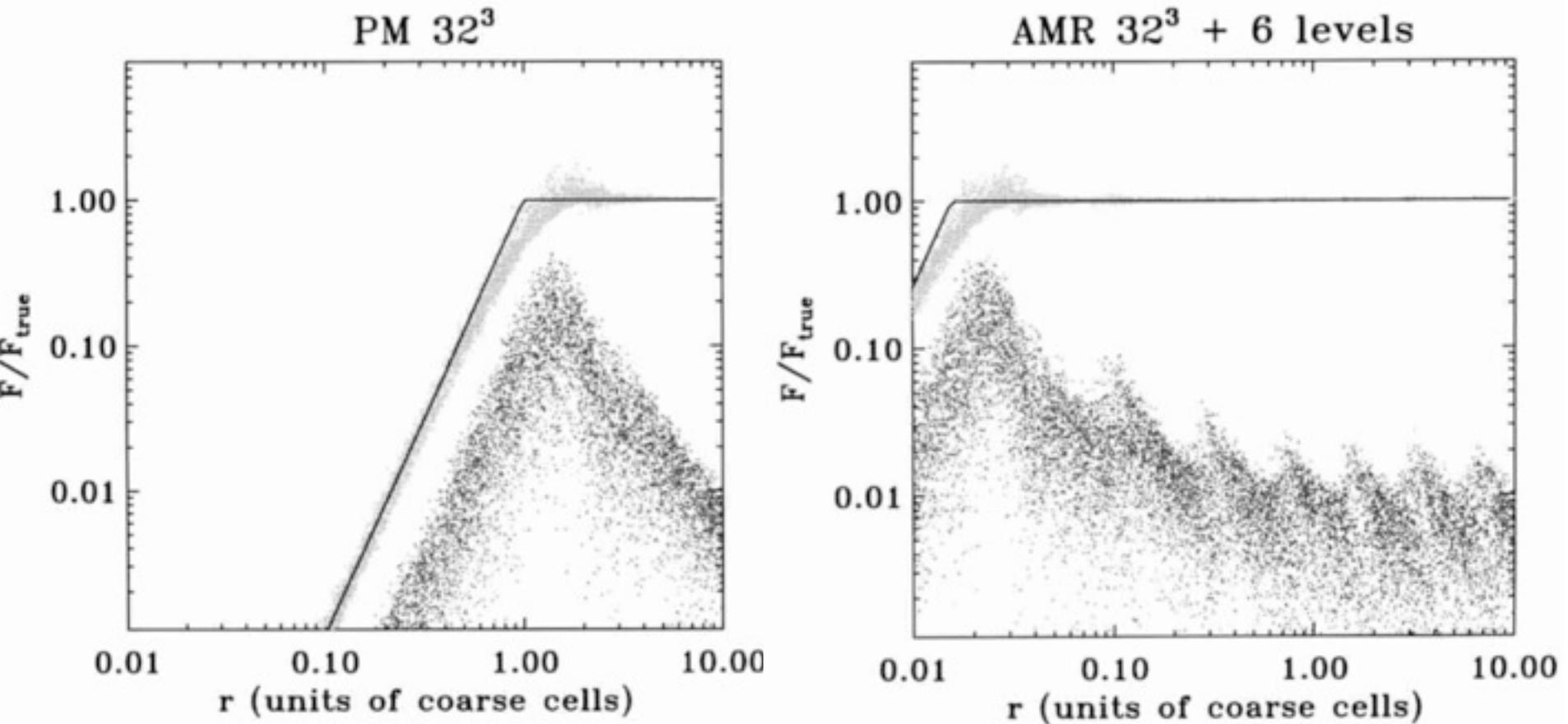
$$\phi_{i,j}^{n+1} = \omega\phi_{i,j}^n + (1 - \omega)\phi_{i,j}^{n+1} \quad \text{with } 1 < \omega < 2$$

Fastest convergence for  $\omega \simeq \frac{2}{1 + \alpha \frac{\pi}{N}}$

Similar performance with the Conjugate Gradient method. For a NxN grid: exact convergence in  $N^2$  iterations,

In practice, order N iterations are necessary to reach the level of truncation errors `epsilon=1e-3 to 1e-4`

# PM-AMR force accuracy



# Multigrid solver for the Poisson equation

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Use coarse-grid sampling to speed-up convergence at large scale.

Proposed by Brandt (1973) to solve elliptic problems.

Use smoothing properties of Jacobi and Gauss-Seidel scheme to reduce high-frequency modes in the error.

Use coarsening to reduce low-frequency modes at a faster rate.

Reduce the cost of relaxation solvers from  $N^2$  to  $N \ln N$  or even  $N$ .

Briggs, W.L., “A Multigrid Tutorial”, SIAM Monograph, (2000)

# Two-grid scheme

On the fine grid, define the residual  
and the error

$$r_\ell^n = \Delta_\ell \Phi_\ell^n - \rho_\ell$$
$$e_\ell^n = \Phi_\ell^n - \Phi_\ell^\infty$$

$$\Delta_\ell e_\ell^n = r_\ell^n$$

1- Perform a few J/GS iterations (smoothing).

2- Restrict the residual to the coarse grid:

$$r_\ell^n \longrightarrow r_{\ell-1}^0$$

3- Solve for the coarse grid system:

$$\Delta_{\ell-1} e_{\ell-1} = r_{\ell-1}$$

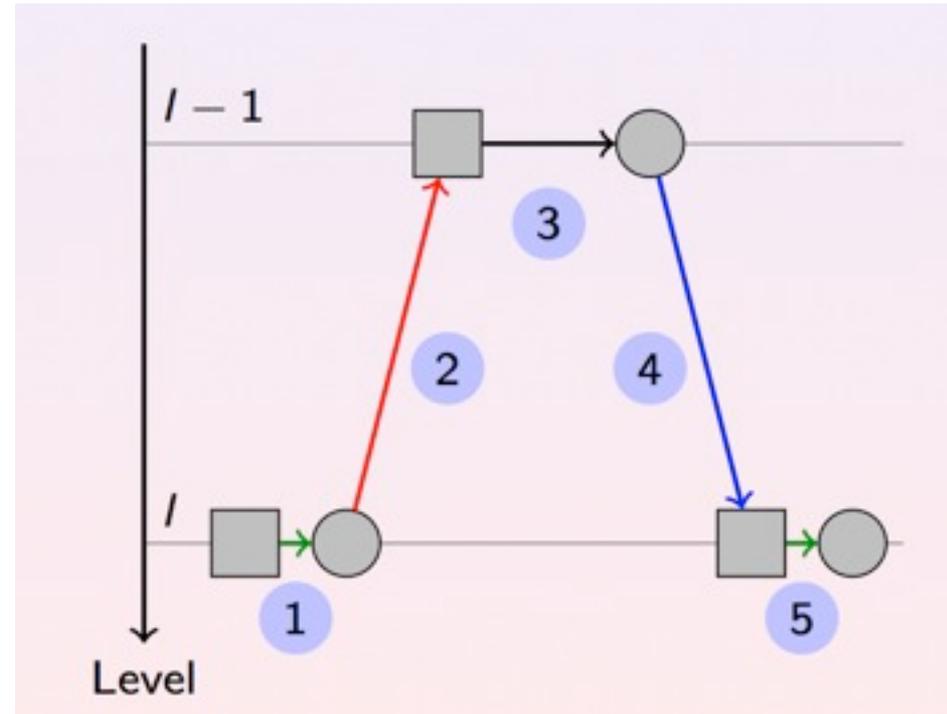
4- Prolong back the error to the fine grid:

$$e_{\ell-1}^\infty \longrightarrow e_\ell^{n+1}$$

5- Correct the fine grid solution:

$$\Phi_\ell^{n+1} = \Phi_\ell^n + e_\ell^{n+1}$$

and perform a few J/GS iteration.

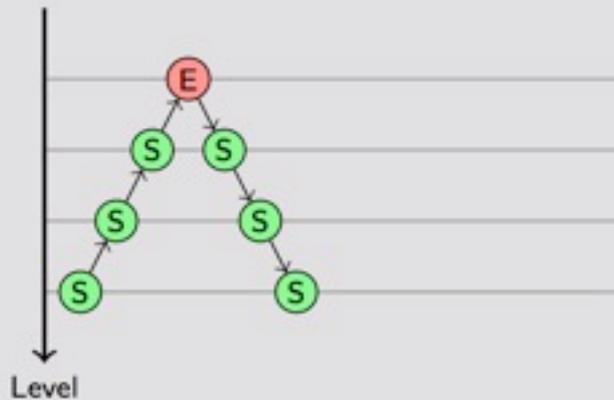


# Multigrid scheme

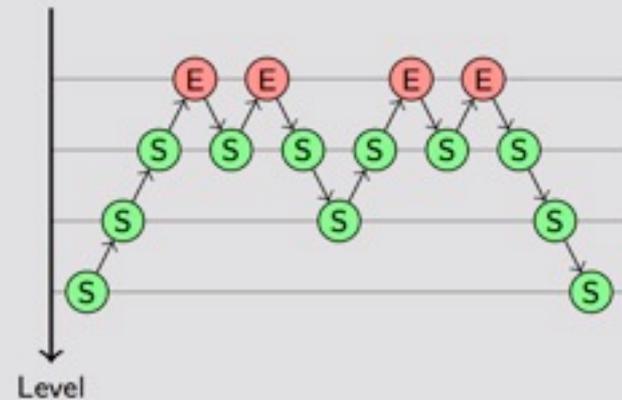
Recursively apply the 2-grid scheme. Solve for the exact solution only at the coarsest level.

Iterate one or twice before going to the finer level.

## Common multigrid schedules



1 iteration per level  
("V-cycle")



2 iterations per level  
("W-cycle")

Converge in very few iterations, independently of grid size.

Quasi-insensitive to the quality of the initial guess.

# Restriction and prolongation operators

**Restriction:** smooth or average the solution. Propagate the information from small to large scales. Use Cloud-In-Cell on the coarse grid.

**Prolongation:** interpolate the solution. Warning: this creates new information. Use Cloud-In-Cell from the coarse grid.

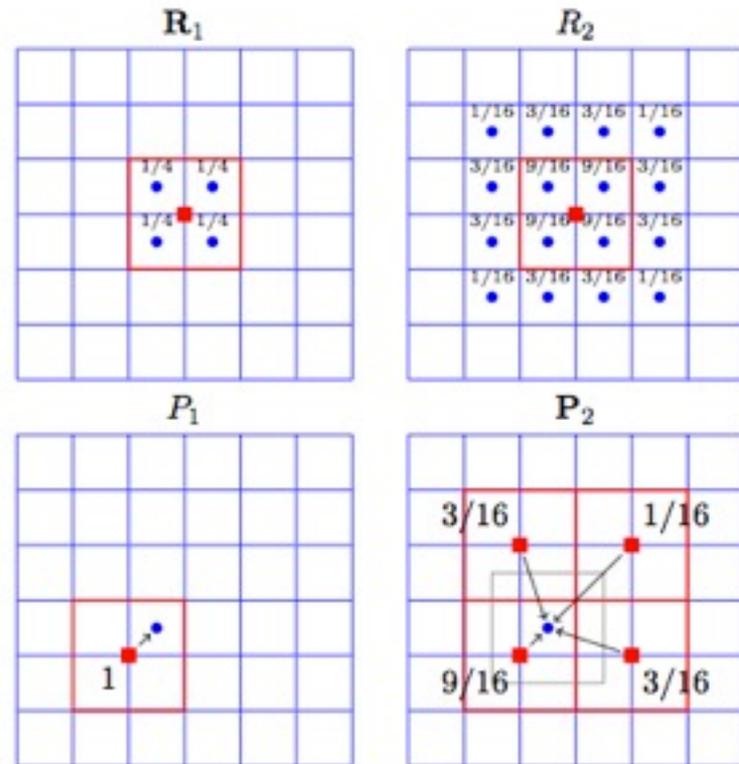
## Rules:

-  $\text{order}(R) + \text{order}(P) > \text{order}(\Delta)$

-  $R^T P = I$

-  $\Delta_{\ell-1} = R \Delta_{\ell} P$

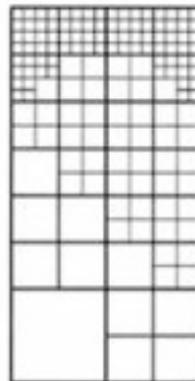
In RAMSES, we use  $R_1$  and  $P_2$



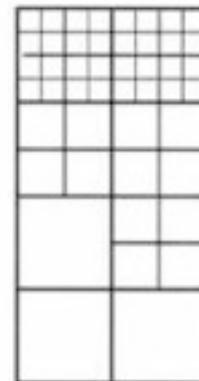
# Multigrid for AMR

**Two-way coupling:** one needs to define an ensemble of AMR grids, each AMR grid corresponds to a level in the multigrid hierarchy.

Berger, M., "An adaptive multigrid method for the euler equations", Springer Berlin



Level 3



Level 2

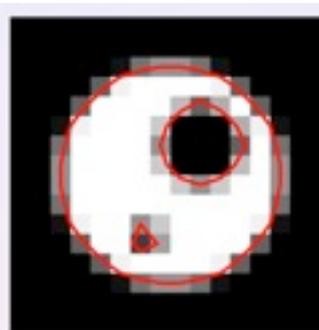


Level 1

**One-way coupling:** for each AMR level, one needs to design a multigrid scheme for arbitrary-shaped boundary conditions.



Level 3



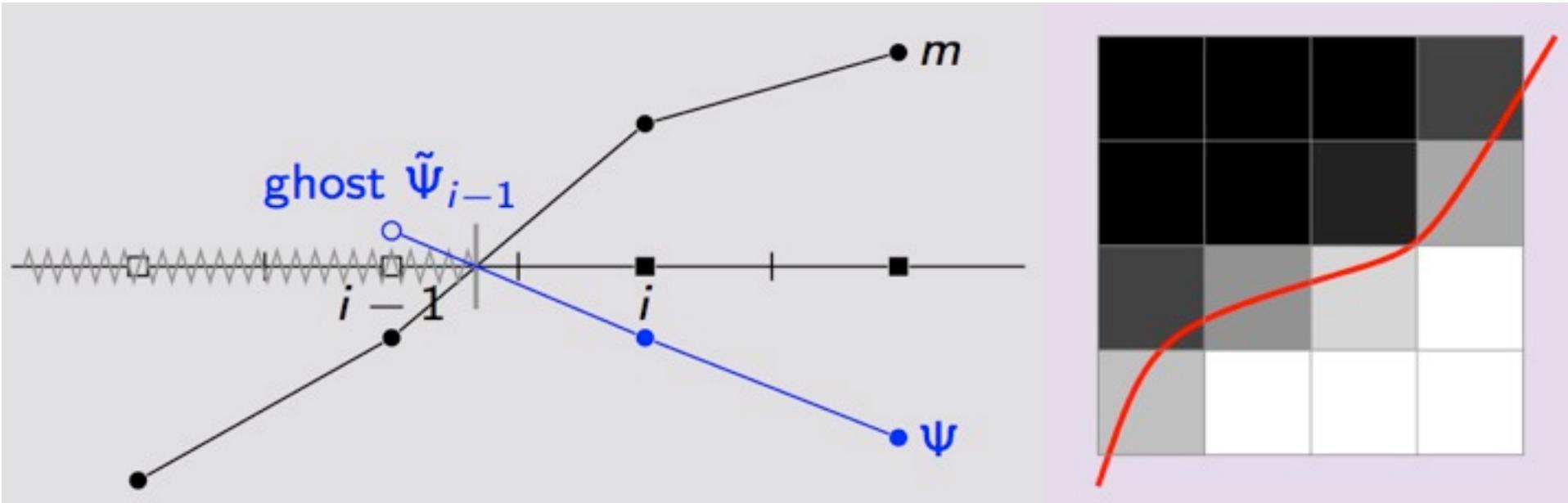
Level 2



Level 1

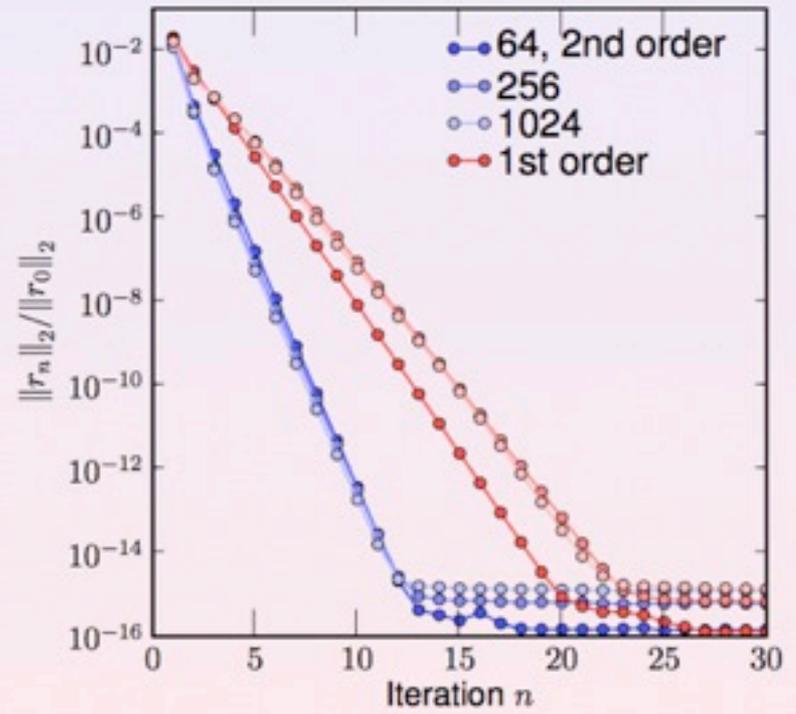
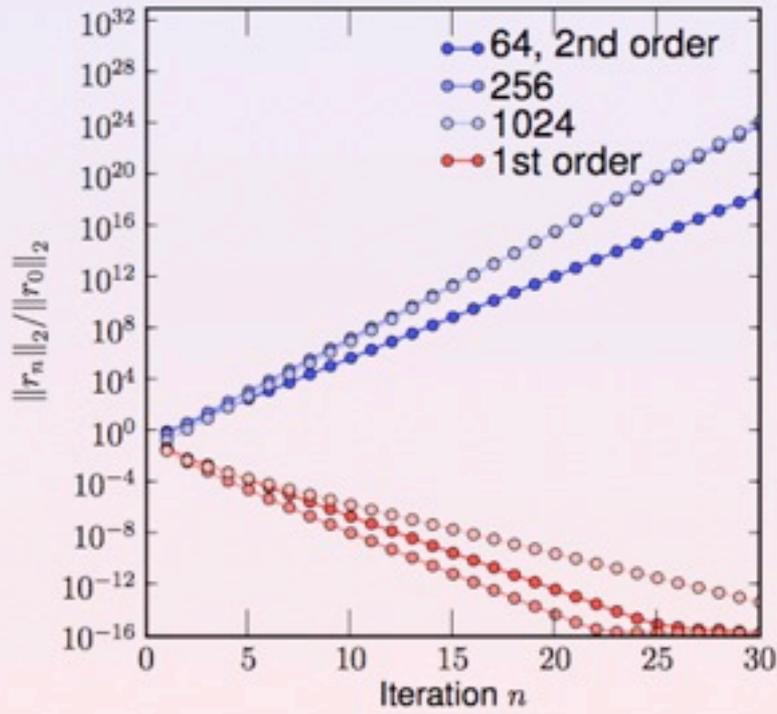
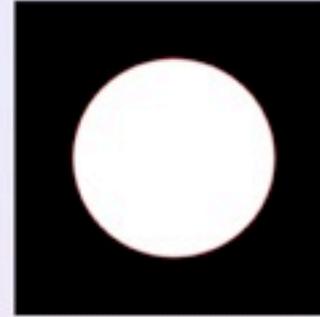
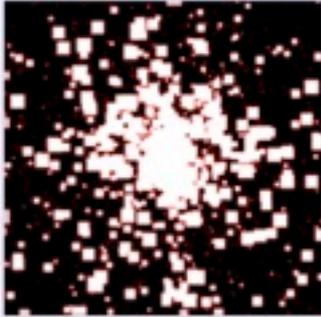
# Boundary-capturing technique

On each level, the boundary is defined as the zero-level set of a domain-fitted function (distance to the interface or volume fraction).

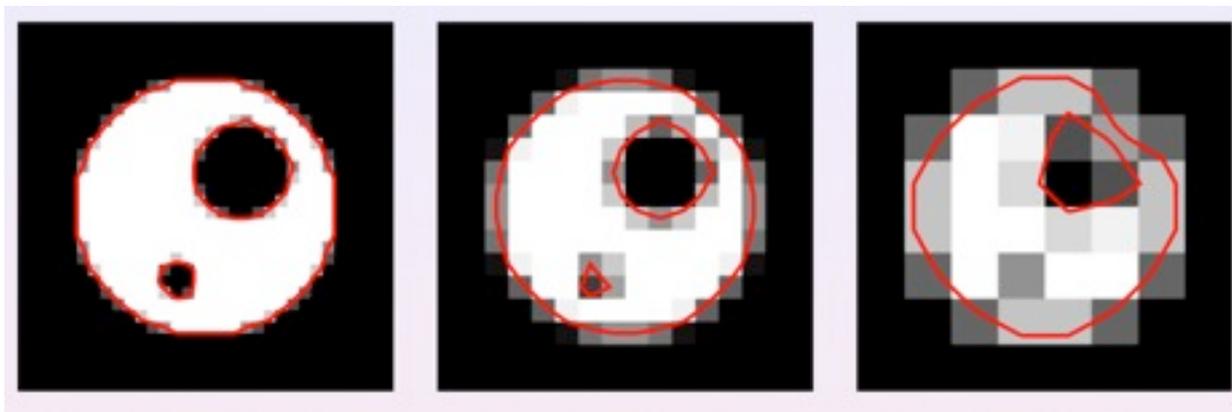


Boundary condition is enforced by linear extrapolation: second order boundary reconstruction as in Gibou (2002).

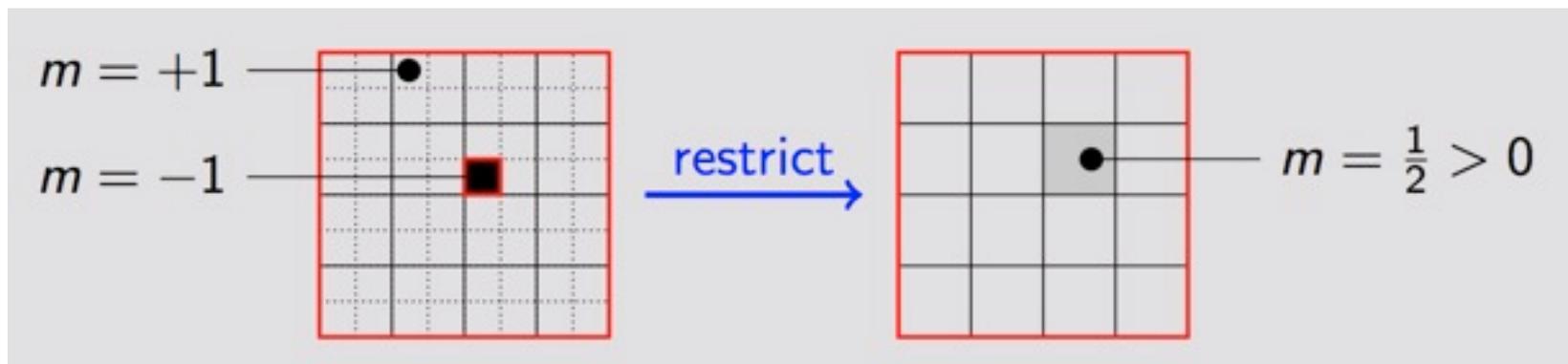
# Performance of multigrid on arbitrary domains



# Small island-problem on complex boundaries



In some cases (holes), the boundary disappears at coarse levels. Poisson solution on coarse levels is different from the one at fine levels. Multigrid diverges ! (Brandt 1995).



A simple solution: Switch to first order boundary-capturing. Slower convergence, partially compensated by W-cycles (Guillet 2009)

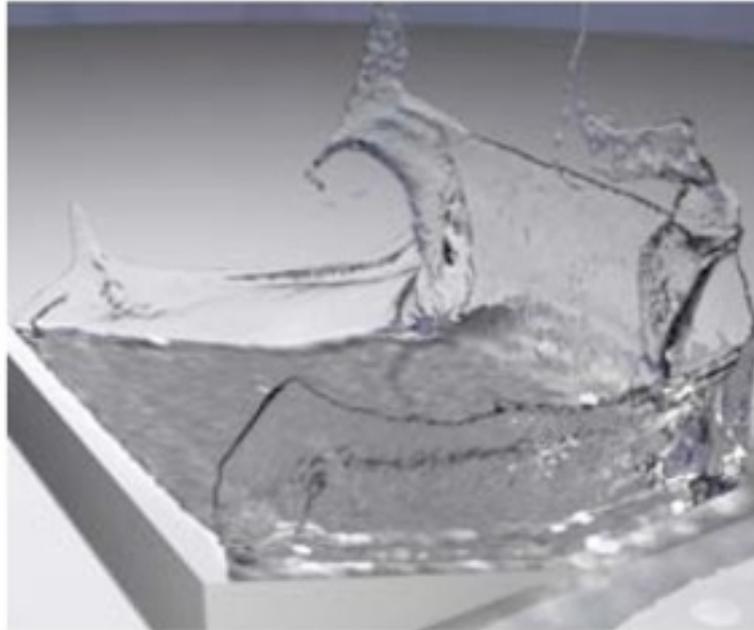
# Conclusion

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- Using force shaping and small scale PP corrections, the Particle Mesh method can reach very high resolution and accuracy. The AP3M method is the ultimate N-body solver for this family. Not well for arbitrary mesh geometry.
- Question: constant force softening versus adaptive force softening ?
- PM-AMR has proven to be a powerful alternative.
- Two distinct approaches for AMR: one-way versus two-way coupling.
- PM-AMR reaches a similar accuracy than other codes (P3M or Tree), although force discontinuities can be an issue.
- PM-AMR is a very efficient method when used in conjunction with AMR-based multigrid solvers.
- RAMSES gravity solver implements most of these schemes.

## A side note...

PIC scheme is used also for compressible and incompressible fluid flows simulations (10-20 particle per cell are needed together with adaptive grids)



Harlow, F.H., “A Machine Calculation Method for Hydrodynamics Problems”, Los Alamos Scientific Laboratory report LAMS-1956 (1955)

Brackbill, J.U., Ruppel, H.M., “FLIP: A Method for Adaptively Zoned, Particle-in-Cell Calculations of Fluid Flows in Two Dimensions”, JCP, **65**, 314 (1986)