Ton ABEL, OLIVER HAHRI, RALT KAEHLER KIPAL/STANTORD Outline: - Dark Mafter Sheef? - COSHOLOGICAL N-BOBY SITTULATIONS Q⁽ - Thase Space - Projecting to configuration space - How use fal is this? Ø - Collisionless fluids - Now woys to simulate IN & collisionless Plasmas.

The Dark Matter Sheet - Con M Dar Matter is commonly hypothicized to originate within seconds ofthe the BIG BANG If it will noving relationstically loday, salaries and other structures would not exist. We speak of COLD DARK MANY IR. Working HYPOTHESIS: - Weakby interacting massive particle (say = 100 GeV). - Very cold. Even kerpentides would only have - is speeds loday. - Almost pareally uniformly distributed initially. - Negligible entert along velocity directions in phase space. У X

The Dark Matter Sheet ? l, vi d # OF DARK MATTER PARTILLES IN THE MILLEY WAY : Non = 1067 (100GeV) >> # OF STARS IN THE UNIVERSE > # OF PARTICLES THAT FIT ON A LOHPOTER USING ALL THE COTPUTERS IN THE WORLD : & 1017 portides SOLVE VLASON - POISSON SYSTEM INSTEAD. 2+ V. V.f. + Q. V.f = 0 J: distribution function IN PHASE CPACE ₹-- 5¢ O: potential FOR PHASE SPALE ELEVIENT TO CONTAIN 10' PARTICLES @ MEAN DENSITY IT HAS TO BE LARGER THAN 500 m . 1/3 (<u>Mor</u> /3 1006ev) 1 ~ 500 m (Mor 1/3 1006ev)

The Dark Matter Sheet? The Dark Matter Sheet? Those space volume is conserved. Mounth My Slow N' (AV) = const SIX States Spatial volume volume in velocity space tedshilt when Ealoobel: lev @z~1000 1006ev @z~1014 Matter density dropped by a factor ~ 1042 since they Tiny initial peculier velocities => distribution function $f(\vec{x}, \vec{v}) = f_0 \delta(\vec{v})$ Tiny initial pecultor velocities 01 is single valued at every \$ DIRAL DIRAL











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3 dimensional manifold in 6D Phase Space - / artural tessellation todes unit where & splits it anto the Six equal size feria hadra. - mass per vervation = 1/6 of IM particle mass. $V = \frac{1\vec{a} \cdot (\vec{b} \times \vec{t})}{6}$ ** $= \sum S = \frac{M_P}{6V} = \frac{M_P}{\left[\vec{p} \cdot (\vec{b} \times \vec{c})\right]}$ N-tody particle - Number the edges of the cube - think of lattice - Looping Over The initial cartesian (LAGIRANGIAN) Lattice journates the GN totahedta. 2





Fig. 9. A direct comparison between our tetrahedral cell-projection approach (left) and a standard SPH adaptive kernel smoothing method. Artifacts due to the poor density estimates in low-density regions are obvious for the SPH method, whereas the tetrahedral approach achieves an overall high image quality, on small and large structures. Application to same N-body data Adaptive Rornel smoothim Shows a number of unphysical clumps Now approach gives exact dassities Recovers all causfics ABEL, HAHN & LAENZER 2011 KAENIN, HAHN & ABEL 2012















Figure 15. Density fields at fixed force resolution with increasing mass resolution for the two methods. Shown is the maximum density projection of a $13 \times 13 \times 40h^{-1}$ Mpc region in the 128^3 , 256^3 and 512^3 particle simulations for TCM (top row) and standard PM (bottom row). In all cases a 512^3 mesh was used to compute the forces. CIC density estimates were obtained on a mesh with two times the number of cells compared to the number of particles per linear dimension, i.e. 256^3 for the 128^3 particle runs for example. Compared to the PM simulations, all the beads-on-a-string artificial fragments are gone in TCM.





Conclusions:

- thinking of the dark matter distribution as a fluid being traced by markers yields a wonderful new way to study the results of N-body simulations.
- visualization techniques much superior to previous approaches:
 - no artificial clumping
 - much better contrast
 - truthful to caustics
- Powerful new techniques to solve the continuum limit.

Refinement in phase space effectively yields a multi resolution technique able to track fine grain phase space structure for dark matter.

New approaches also useful for studying collision-less plasmas.

Serious drawback: TCM has to resolve the sheet which however can grow exponentially in halos.... Refinement will help but not overcome this limitation.