

# Light Profile Decompositions from M31 to Virgo

with **Mike McDonald (Maryland)**

**Yucong Zhu (Harvard)**

Larry Widrow (Queen's)

Raja Guhathakurta (UCSC)

Stéphane Courteau



Queen's  
UNIVERSITY

# Light Profile Decompositions from M31 to Virgo

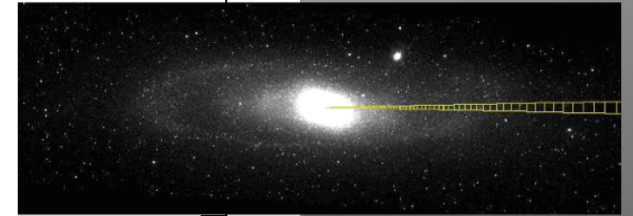
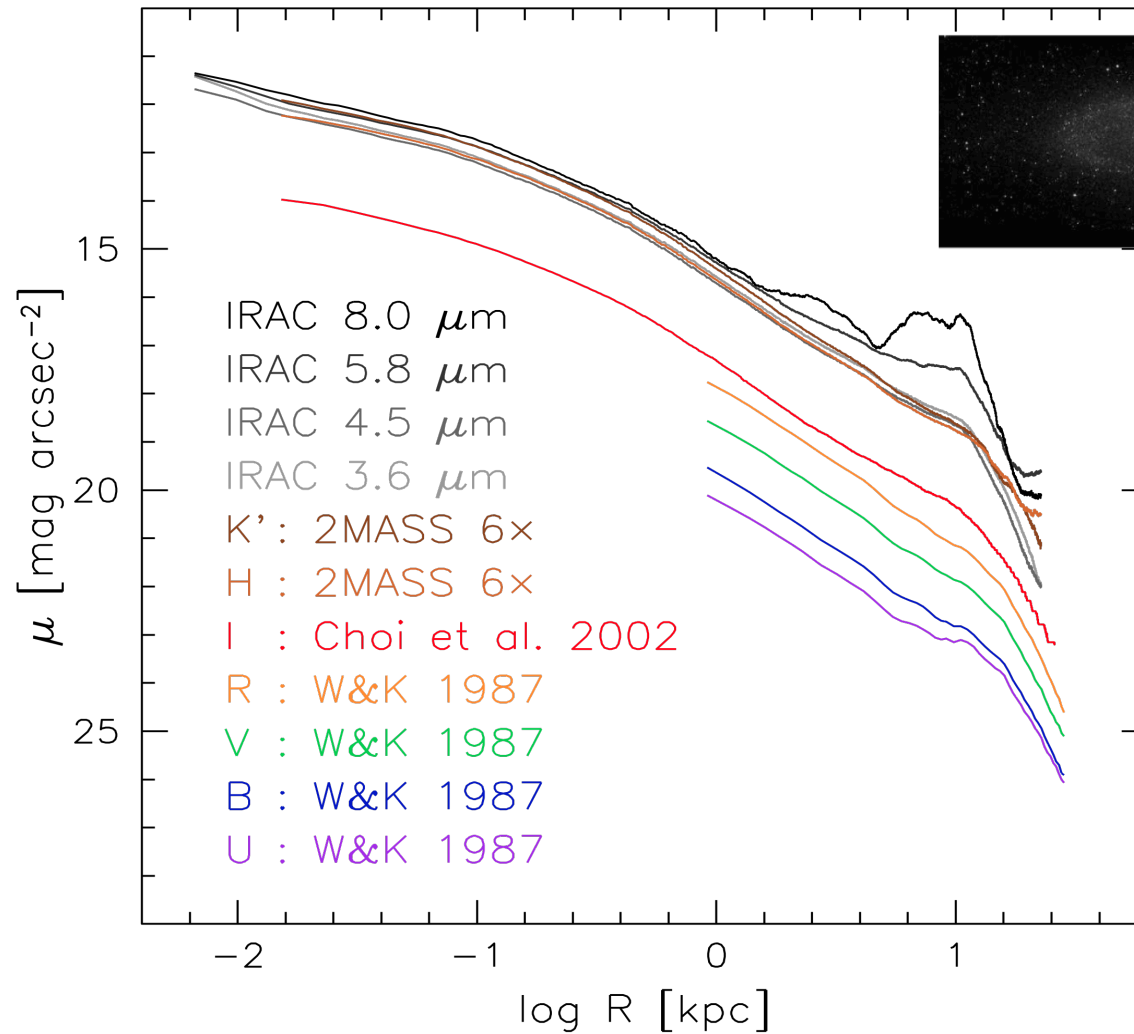
- Testing decomposition techniques on M31
- Testing 2D decomposition models with 2-component models of galaxies
- B/D decompositions of Virgo galaxies

# Model Decompositions

## techniques:

- 1D NLLS (Levenberg-Marquardt)
- Bayesian / Markov-chain Monte Carlo (MCMC)
- 2D NLLS (GALFIT), 2D MCMC (GIM2D)
- *bin vs unbin*
- *minimization counts vs mag*
- *features (spiral arm) clipping/modeling*
- *cuts vs azimuthally-averaged SB profiles*

# Available light profiles for M31



IRAC  
3.6 $\mu\text{m}$   
image

# 1D NLLS vs MCMC?

Courteau+2010



TABLE 2  
BULGE/DISK DECOMPOSITIONS FOR M31

Data	Method	Cut	$n$	$R_e$ (kpc)	$\mu_e$ (mag arcsec <sup>-2</sup> )	$\epsilon_{\text{bulge}}$	$R_d$ (kpc)	$\mu_0$ (mag arcsec <sup>-2</sup> )	$\epsilon_{\text{disk}}$	
A.	IRAC	MCMC	Maj	$2.46 \pm 0.09$	$1.09 \pm 0.07$	$16.2 \pm 0.10$	--	$5.09 \pm 0.17$	$16.83 \pm 0.07$	--
B.	IRAC	NLLS	Maj	$2.3 \pm 0.3$	$1.0 \pm 0.2$	$16.1 \pm 0.3$	--	$5.4 \pm 0.3$	$16.8 \pm 0.1$	--
C.	IRAC	MCMC	Min	$2.01 \pm 0.11$	$0.53 \pm 0.05$	$15.47 \pm 0.13$	--	$1.29 \pm 0.09$	$16.49 \pm 0.25$	--
D.	IRAC	NLLS	Min	$1.9 \pm 0.6$	$0.5 \pm 0.3$	$15.4 \pm 0.7$	--	$1.3 \pm 0.6$	$16.4 \pm .15$	--
E.	IRAC	MCMC	MinMaj	$2.18 \pm 0.06$	$0.82 \pm 0.04$	$15.77 \pm 0.07$	$0.21 \pm 0.01$	$4.71 \pm 0.14$	$16.62 \pm 0.05$	$0.74 \pm 0.01$
F.	IRAC	NLLS	AZAV	$2.4 \pm 0.2$	$1.10 \pm 0.10$	$16.1 \pm 0.10$	--	$5.8 \pm 0.1$	$16.79 \pm 0.02$	$0.74 \pm 0.02$
F*.	IRAC	MCMC	AZAV	$1.66 \pm 0.03$	$0.68 \pm 0.01$	$15.34 \pm 0.03$	--	$4.75 \pm 0.01$	$16.41 \pm 0.01$	--
G.	IRAC	NLLS	AZAVmsk	$2.2 \pm 0.3$	$1.00 \pm 0.10$	$16.0 \pm 0.20$	--	$4.9 \pm 0.1$	$16.70 \pm 0.10$	$0.74 \pm 0.02$
H.	Choi02	MCMC	Maj	$2.06 \pm 0.06$	$0.91 \pm 0.04$	$18.12 \pm 0.08$	--	$5.69 \pm 0.09$	$18.97 \pm 0.03$	--
I.	Choi02	NLLS	Maj	$2.2 \pm 0.3$	$1.12 \pm 0.10$	$17.3 \pm 0.2$	--	$6.4 \pm 0.1$	$18.14 \pm 0.04$	--
J.	Choi02	MCMC	Min	$1.85 \pm 0.07$	$0.51 \pm 0.02$	$17.62 \pm 0.08$	--	$1.73 \pm 0.05$	$18.99 \pm 0.08$	--
K.	Choi02	NLLS	Min	$1.9 \pm 0.2$	$0.53 \pm 0.03$	$17.7 \pm 0.1$	--	$1.68 \pm 0.04$	$18.8 \pm 0.1$	--
L.	Choi02	MCMC	MinMaj	$1.83 \pm 0.04$	$0.74 \pm 0.02$	$17.73 \pm 0.05$	$0.28 \pm 0.01$	$5.47 \pm 0.08$	$18.91 \pm 0.03$	$0.70 \pm 0.01$
M.	Choi02	NLLS	AZAV	$2.00 \pm 0.4$	$1.00 \pm 0.30$	$18.2 \pm 0.4$	--	$5.80 \pm 0.10$	$19.0 \pm 0.3$	$0.71 \pm 0.02$



# 2D (GALFIT) power...

**Bulge:**

$$\epsilon_{bulge} (2D) \sim 0.37$$

$$\epsilon_{bulge} (1D) \sim 0.21$$

**Disk**

$$\epsilon_{disk} (2D) \sim 0.73$$

$$\epsilon_{disk} (1D) \sim 0.71$$

$$\Delta(PA) \approx 20^\circ$$

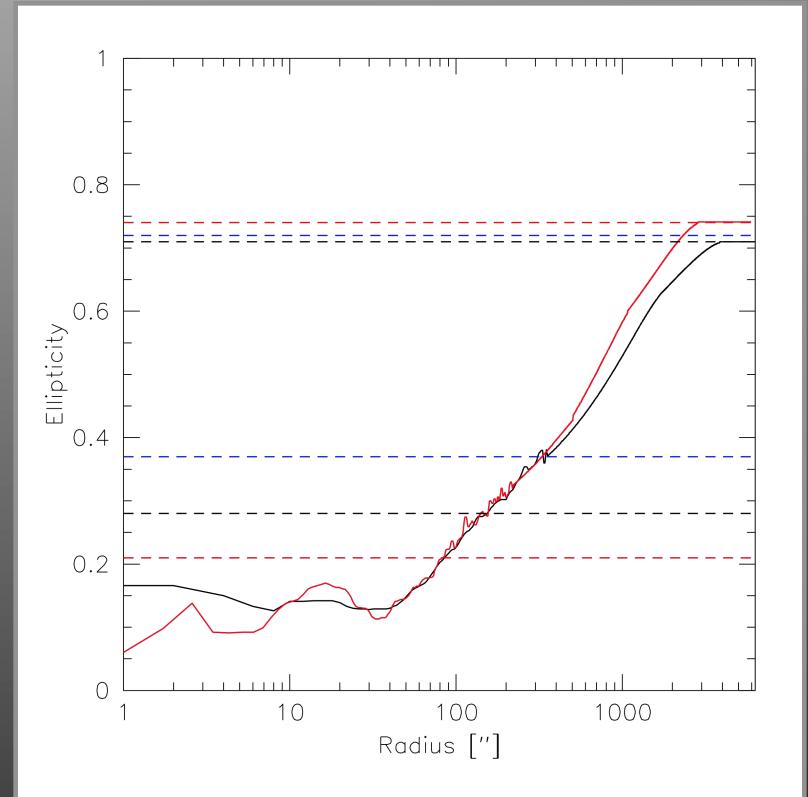


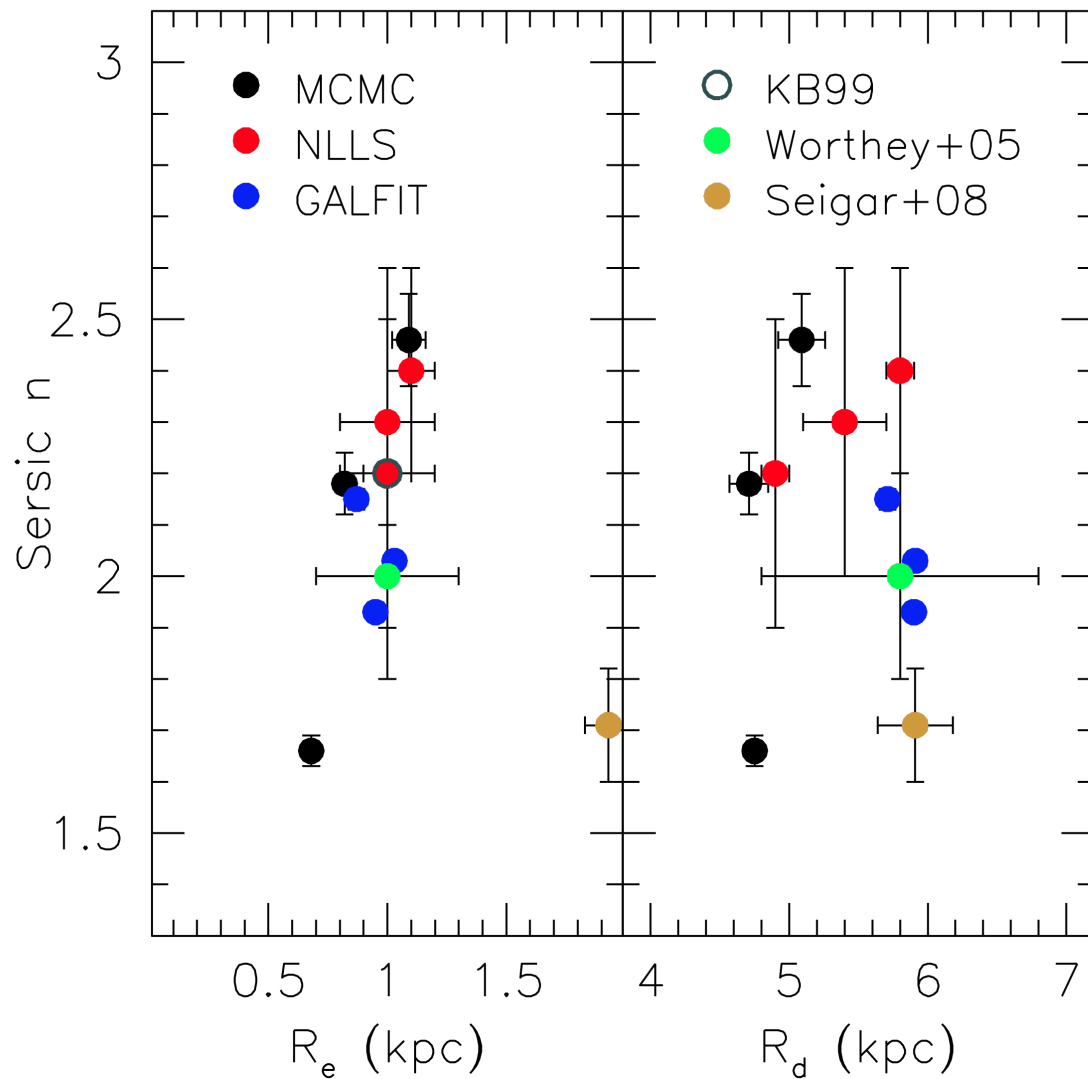
TABLE 2  
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A.	IRAC	MCMC	Maj	$2.46 \pm 0.09$	$1.09 \pm 0.07$	$16.2 \pm 0.10$	--	$5.09 \pm 0.17$	$16.83 \pm 0.07$	--
B.	IRAC	NLLS	Maj	$2.3 \pm 0.3$	$1.0 \pm 0.2$	$16.1 \pm 0.3$	--	$5.4 \pm 0.3$	$16.8 \pm 0.1$	--
C.	IRAC	MCMC	Min	$2.01 \pm 0.11$	$0.53 \pm 0.05$	$15.47 \pm 0.13$	--	$1.29 \pm 0.09$	$16.49 \pm 0.25$	--
D.	IRAC	NLLS	Min	$1.9 \pm 0.6$	$0.5 \pm 0.3$	$15.4 \pm 0.7$	--	$1.3 \pm 0.6$	$16.4 \pm .15$	--
E.	IRAC	MCMC	MinMaj	$2.18 \pm 0.06$	$0.82 \pm 0.04$	$15.77 \pm 0.07$	$0.21 \pm 0.01$	$4.71 \pm 0.14$	$16.62 \pm 0.05$	$0.74 \pm 0.01$
F.	IRAC	NLLS	AZAV	$2.4 \pm 0.2$	$1.10 \pm 0.10$	$16.1 \pm 0.10$	--	$5.8 \pm 0.1$	$16.79 \pm 0.02$	$0.74 \pm 0.02$
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I.	Choi02	NLLS	Maj	$2.2 \pm 0.3$	$1.12 \pm 0.10$	$17.3 \pm 0.2$	--	$6.4 \pm 0.1$	$18.14 \pm 0.04$	--
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K.	Choi02	NLLS	Min	$1.9 \pm 0.2$	$0.53 \pm 0.03$	$17.7 \pm 0.1$	--	$1.68 \pm 0.04$	$18.8 \pm 0.1$	--
L.	Choi02	MCMC	MinMaj	$1.83 \pm 0.04$	$0.74 \pm 0.02$	$17.73 \pm 0.05$	$0.28 \pm 0.01$	$5.47 \pm 0.08$	$18.91 \pm 0.03$	$0.70 \pm 0.01$
M.	Choi02	NLLS	AZAV	$2.00 \pm 0.4$	$1.00 \pm 0.30$	$18.2 \pm 0.4$	--	$5.80 \pm 0.10$	$19.0 \pm 0.3$	$0.71 \pm 0.02$

TABLE 3  
GALFIT BULGE/DISK DECOMPOSITIONS OF THE M31 IRAC 3.6  $\mu$  IMAGE

Method	$n$	$R_e$ (kpc)	$\mu_e$ (mag arcsec <sup>-2</sup> )	$\epsilon_{\text{bulge}}$	$R_d$ (kpc)	$\mu_0$ (mag arcsec <sup>-2</sup> )	$\epsilon_{\text{disk}}$
N. Original Image	$1.93 \pm 0.00$	$0.95 \pm$	$16.06 \pm$	$0.37 \pm$	$5.90 \pm$	$17.06 \pm$	$0.72 \pm$
O. Masked Image	$2.03 \pm 0.00$	$1.03 \pm$	$16.18 \pm$	$0.37 \pm$	$5.91 \pm$	$17.14 \pm$	$0.72 \pm$
P. Forced $\epsilon_{\text{bulge}}, \epsilon_{\text{disk}}$	$2.15 \pm 0.00$	$0.87 \pm$	$16.87 \pm$	$0.21 \pm$	$5.71 \pm$	$16.77 \pm$	$0.74 \pm$

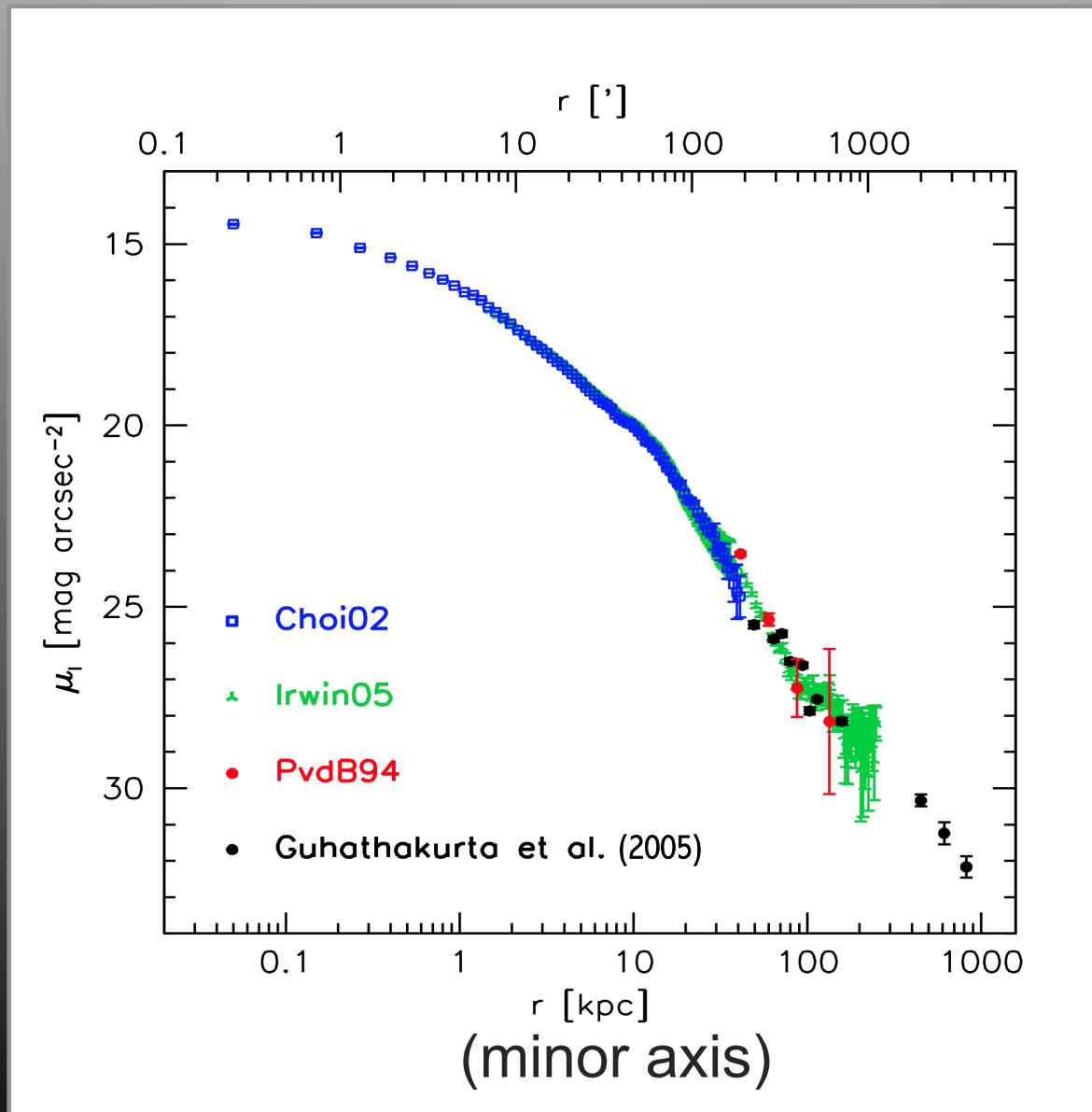
# Parameter range



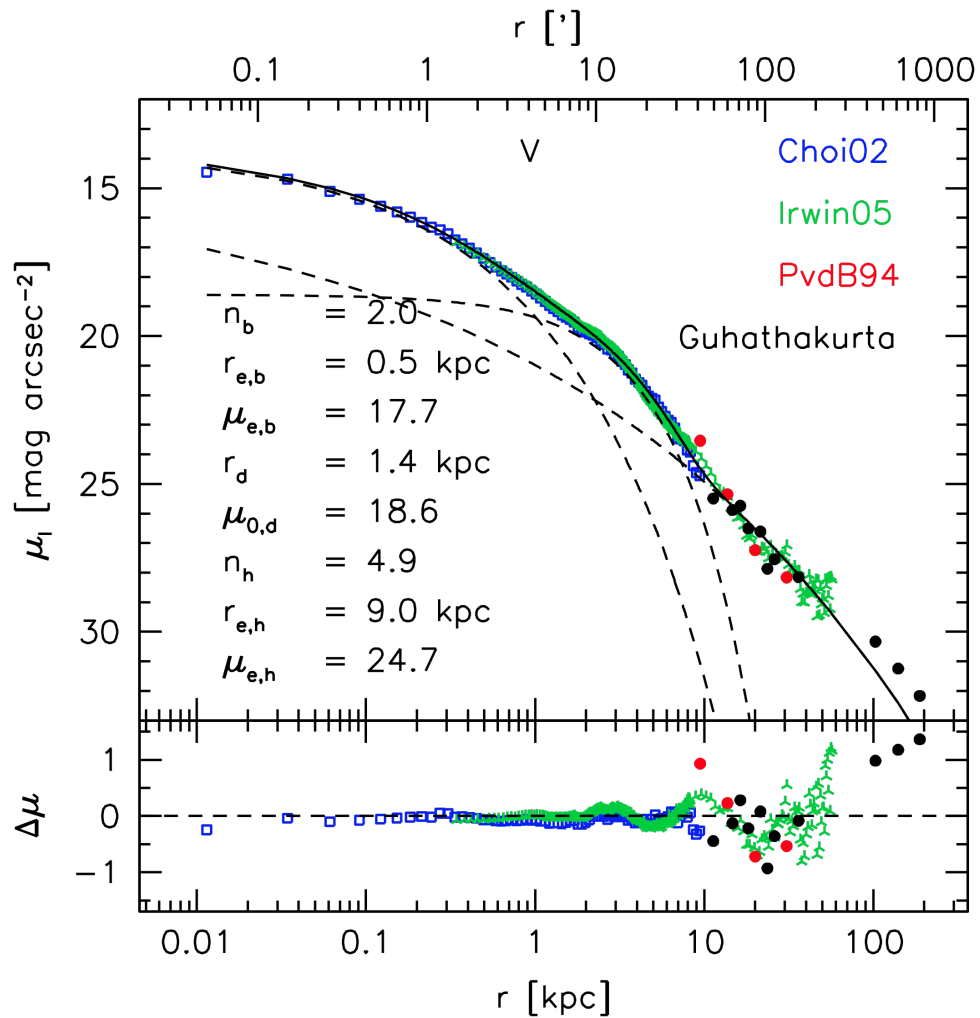
$\sim 20\%$  parameter uncertainty



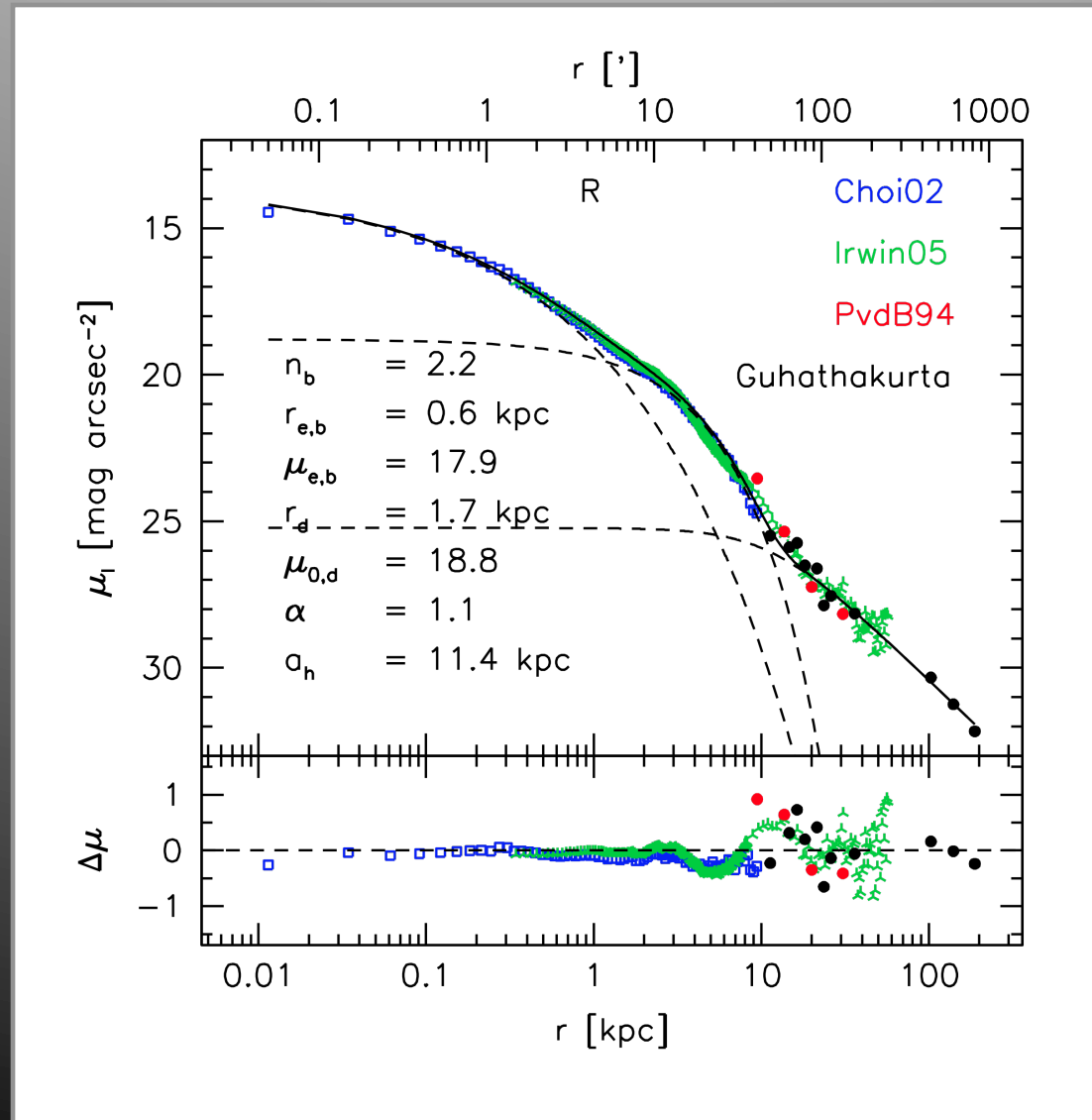
# Halo Star Counts



# Bulge / Disk / Sersic halo



# Bulge / Disk / Power law halo



# M31 B/D/H Decomps

TABLE 4  
I-BAND COMPOSITE; BULGE/DISK/POWER-LAW FAINT

Method	Cut	$n$	$R_{e,b}$ (kpc)	$\mu_{e,b}$ (mag arcsec <sup>-2</sup> )	$R_d$ (kpc)	$\mu_0$ (mag arcsec <sup>-2</sup> )	$\alpha$	$\mu_*$ (mag arcsec <sup>-2</sup> )	$a_h$ (kpc)	
Q.	NLLS	Min/Bin	2.20 <sup>+20</sup> <sub>-20</sub>	0.60 <sup>+10</sup> <sub>-10</sub>	17.90 <sup>+30</sup> <sub>-30</sub>	1.70 <sup>+30</sup> <sub>-30</sub>	18.80 <sup>+30</sup> <sub>-30</sub>	1.10 <sup>+10</sup> <sub>-10</sub>	27.70 <sup>+10</sup> <sub>-10</sub>	11.40 <sup>+0.10</sup> <sub>-0.10</sub>
R.	MCMC	Min/Bin	2.67 <sup>+11</sup> <sub>-11</sub>	1.77 <sup>+05</sup> <sub>-05</sub>	19.53 <sup>+07</sup> <sub>-07</sub>	--	--	1.78 <sup>+29</sup> <sub>-30</sub>	27.90 <sup>+08</sup> <sub>-08</sub>	55.65 <sup>+12.3</sup> <sub>-12.6</sub>
S.	MCMC	Min/Bin	2.72 <sup>+11</sup> <sub>-11</sub>	1.75 <sup>+05</sup> <sub>-05</sub>	19.52 <sup>+06</sup> <sub>-07</sub>	--	--	1.77 <sup>+31</sup> <sub>-31</sub>	28.05 <sup>+08</sup> <sub>-08</sub>	55.47 <sup>+12.7</sup> <sub>-12.9</sub>
T.	MCMC	Min/Unbin	1.77 <sup>+22</sup> <sub>-22</sub>	0.45 <sup>+07</sup> <sub>-08</sub>	17.50 <sup>+24</sup> <sub>-27</sub>	1.35 <sup>+18</sup> <sub>-22</sub>	18.48 <sup>+26</sup> <sub>-31</sub>	1.11 <sup>+25</sup> <sub>-19</sub>	27.90 <sup>+24</sup> <sub>-30</sub>	3.05 <sup>+90</sup> <sub>-80</sub>
U.	MCMC	Min/UnBin	1.89 <sup>+10</sup> <sub>-10</sub>	0.51 <sup>+03</sup> <sub>-03</sub>	17.66 <sup>+11</sup> <sub>-11</sub>	1.54 <sup>+05</sup> <sub>-05</sub>	18.72 <sup>+09</sup> <sub>-09</sub>	1.14 <sup>+09</sup> <sub>-09</sub>	28.32 <sup>+09</sup> <sub>-09</sub>	6.60 <sup>+1.22</sup> <sub>-1.20</sub>
V.	MCMC	MinMaj/UnBin	1.85 <sup>+05</sup> <sub>-05</sub>	0.70 <sup>+02</sup> <sub>-02</sub>	17.73 <sup>+05</sup> <sub>-05</sub>	4.89 <sup>+06</sup> <sub>-06</sub>	18.76 <sup>+02</sup> <sub>-02</sub>	1.17 <sup>+06</sup> <sub>-06</sub>	28.37 <sup>+08</sup> <sub>-08</sub>	6.48 <sup>+1.10</sup> <sub>-1.11</sub>

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Courteau et al. 2010

TABLE 5  
I-BAND COMPOSITE; BULGE/DISK/SÉRSIC FAINT

Method	Cut	$n$	$R_{e,b}$ (kpc)	$\mu_{e,b}$ (mag arcsec <sup>-2</sup> )	$R_d$ (kpc)	$\mu_0$ (mag arcsec <sup>-2</sup> )	$n_f$	$\mu_{e,f}$ (mag arcsec <sup>-2</sup> )	$R_{e,f}$ (kpc)	
X.	NLLS	Min <sup>†</sup> /Bin/Mag	2.00 <sup>+30</sup> <sub>-30</sub>	0.50 <sup>+10</sup> <sub>-10</sub>	17.70 <sup>+20</sup> <sub>-20</sub>	5.00 <sup>+70</sup> <sub>-70</sub>	18.60 <sup>+30</sup> <sub>-30</sub>	4.90 <sup>+0.80</sup> <sub>-0.80</sub>	24.70 <sup>+2.0</sup> <sub>-2.0</sub>	9.00 <sup>+13.0</sup> <sub>-13.0</sub>
Z.	MCMC	MinMaj/Unbin/Cts	1.79 <sup>+06</sup> <sub>-06</sub>	0.69 <sup>+02</sup> <sub>-02</sub>	17.71 <sup>+05</sup> <sub>-05</sub>	4.90 <sup>+06</sup> <sub>-06</sub>	18.76 <sup>+02</sup> <sub>-02</sub>	6.63 <sup>+1.16</sup> <sub>-1.15</sub>	27.30 <sup>+7.2</sup> <sub>-7.0</sub>	19.72 <sup>+5.9</sup> <sub>-5.4</sub>

# Halo modeling

*Bulge: 30% light*      ( $R_e/R_d = .2$ )

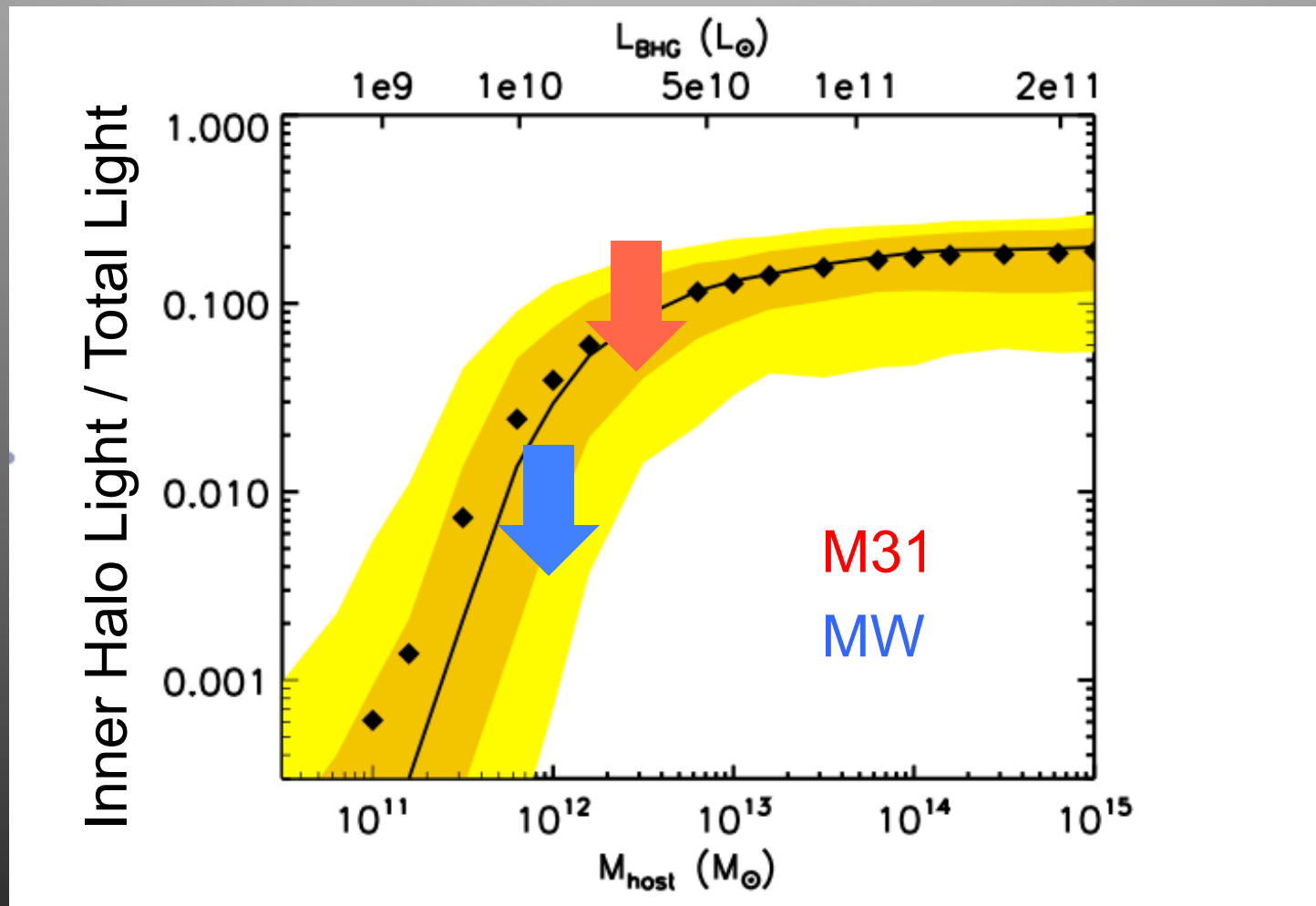
*Disk: 70% light*       $B/D = .43$

***Add halo (27 mag arcsec<sup>-2</sup>):***

*Nucleus(0.05%)*      *Bulge (34%)*

*Disk (47%)*      *Halo (19%)*

# Fractional Halo Light



Purcell, Bullock, Zentner (2008)

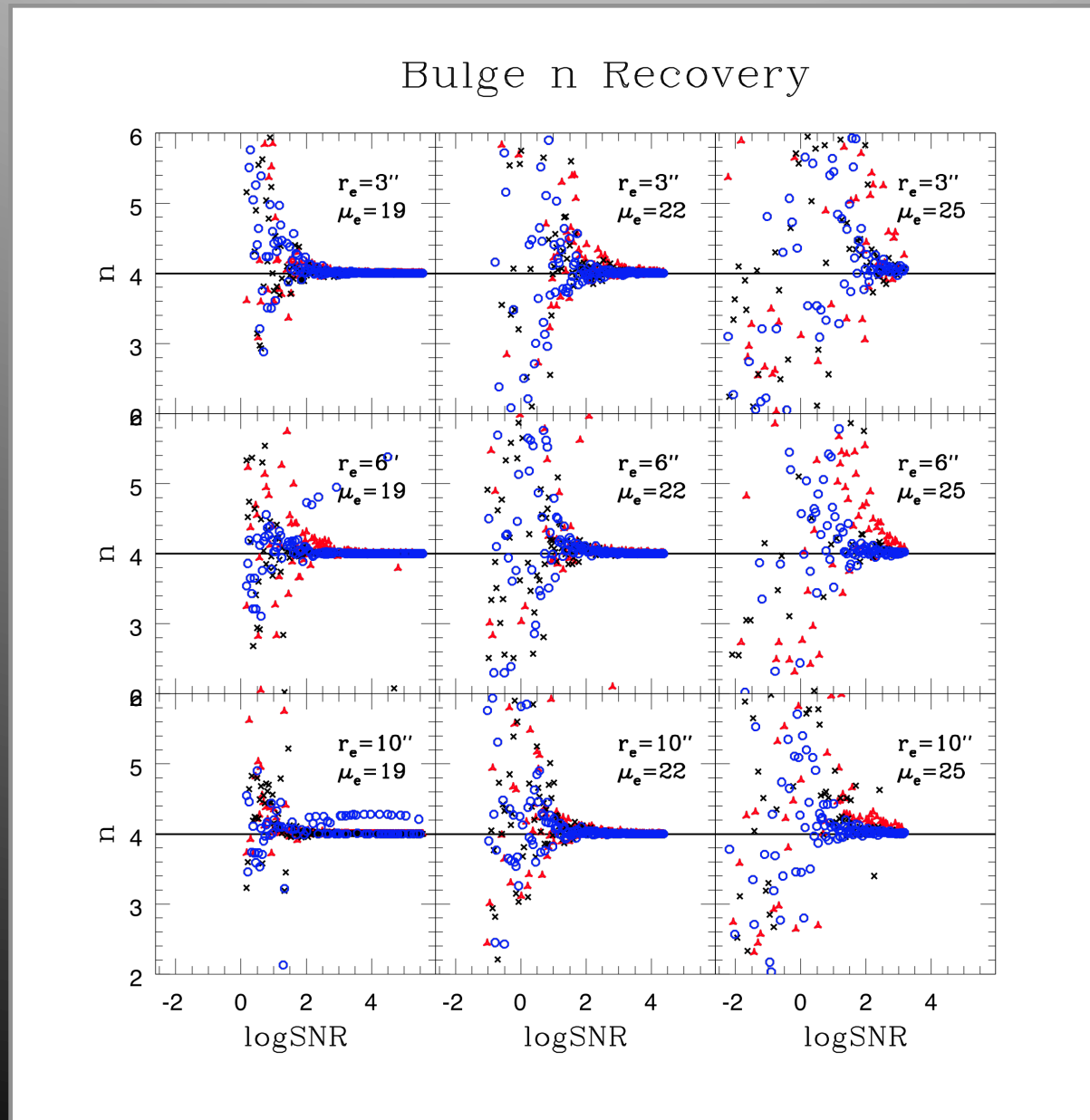
Courteau & van den Bergh (1999)

# 1D / 2D simul<sup>n</sup>s (after MacArthur et al 2003)

- create a suite of 2-component model galaxies (full range of structural parameters,  $\Delta PA$ , SNR, seeing, ...)
- Reduce images via 1D and 2D decomposition minimization techniques
- test 2D (GALFIT)'s recovery

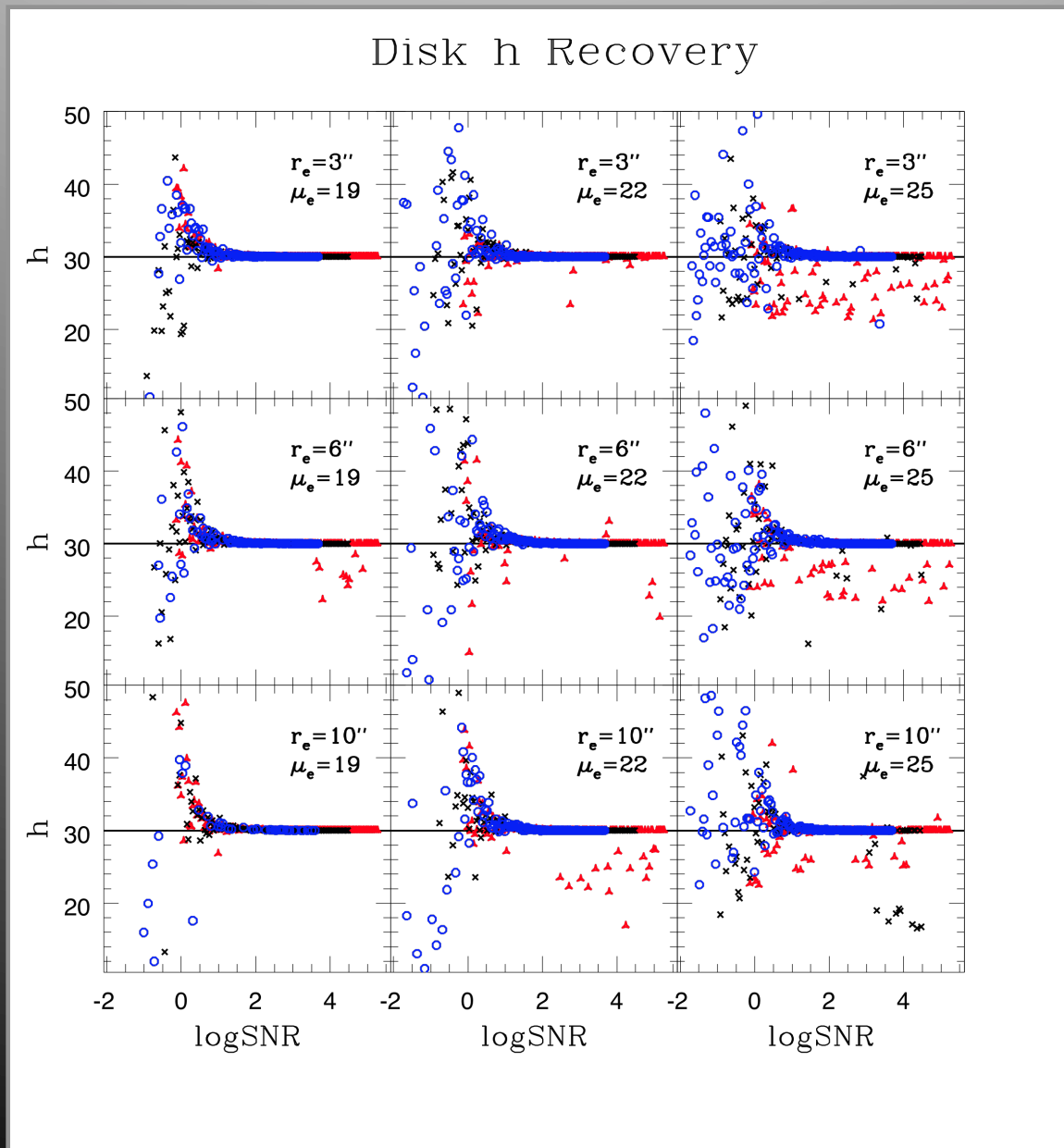
Define:  $SNR = 10^{\{0.4 * (\mu_{sky} - \mu_{eff})\}}$

# 2D model recovery – $n$



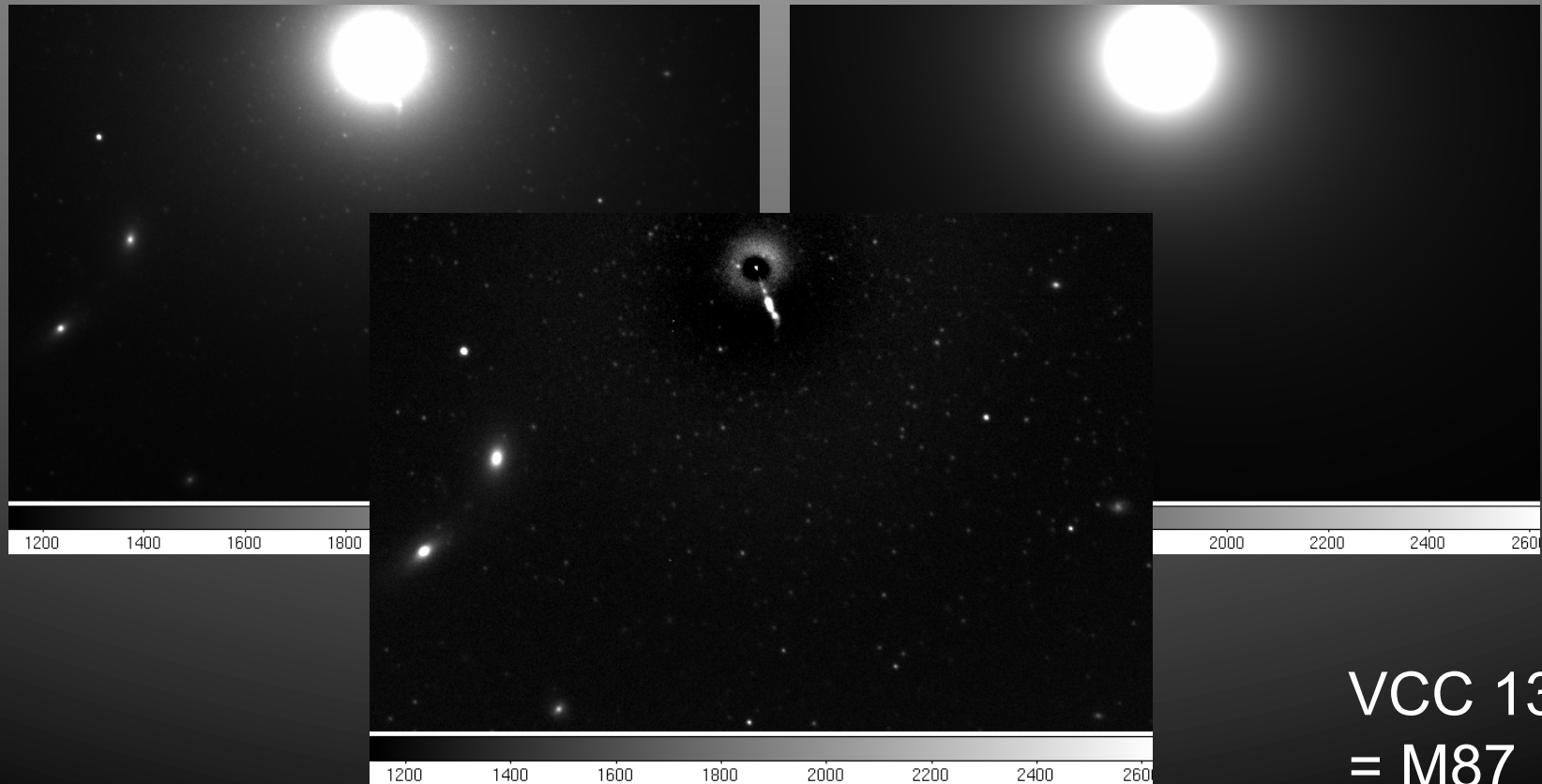


# 2D model recovery – $R_d$



# GALFIT analysis of 786 VCC/SDSS galaxies

- 1D isophotal fits (Courteau et al 1996)  
+ 1D B/D decomps (McDonald et al 2009)
- 2D GALFIT (Peng et al 2002)

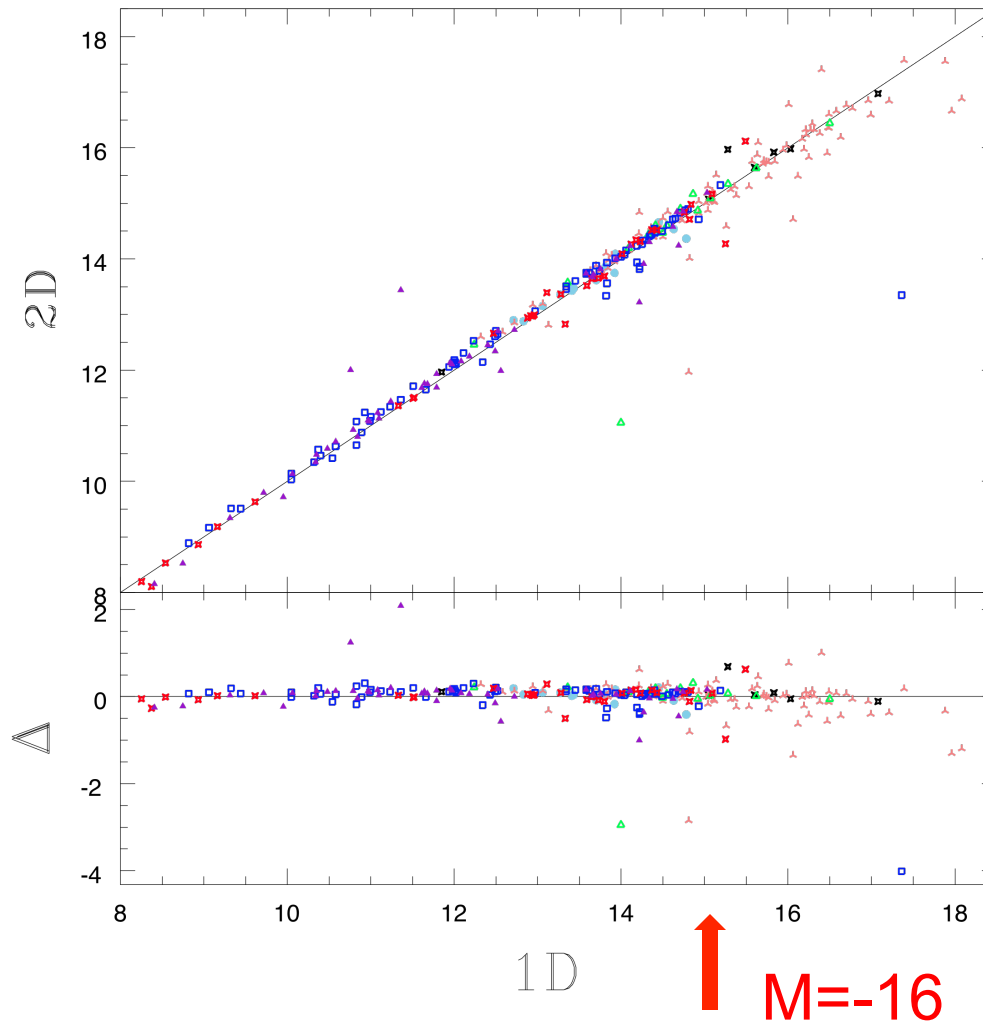


VCC 1316  
= M87

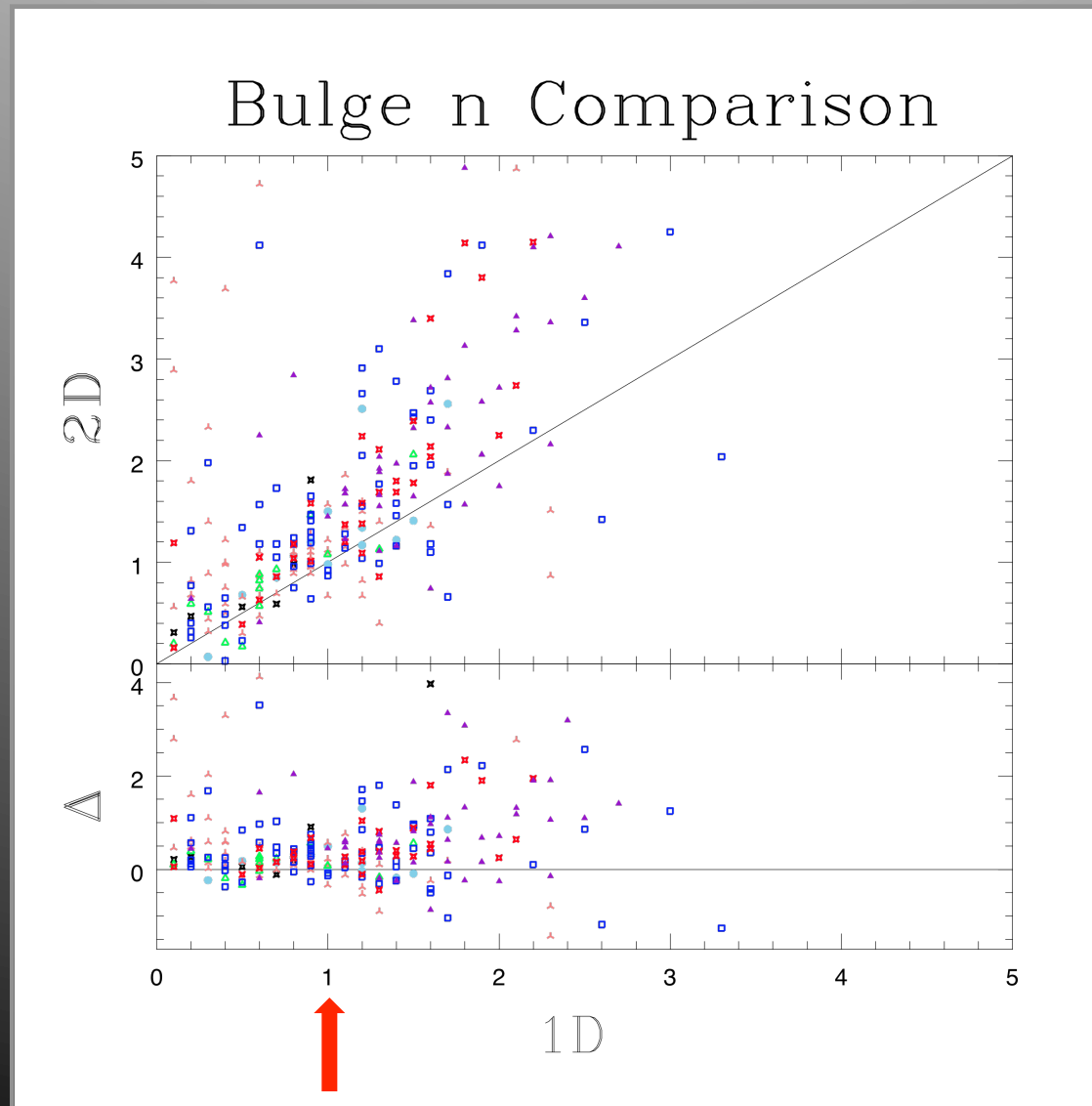
# 1D vs 2D for 786 VCC/SDSS galaxies (number reduces to ~300 gals for good GALFIT fits)

Spirals  
Lenticulars  
Ellipticals  
dE's  
dSp  
Irr  
unknown

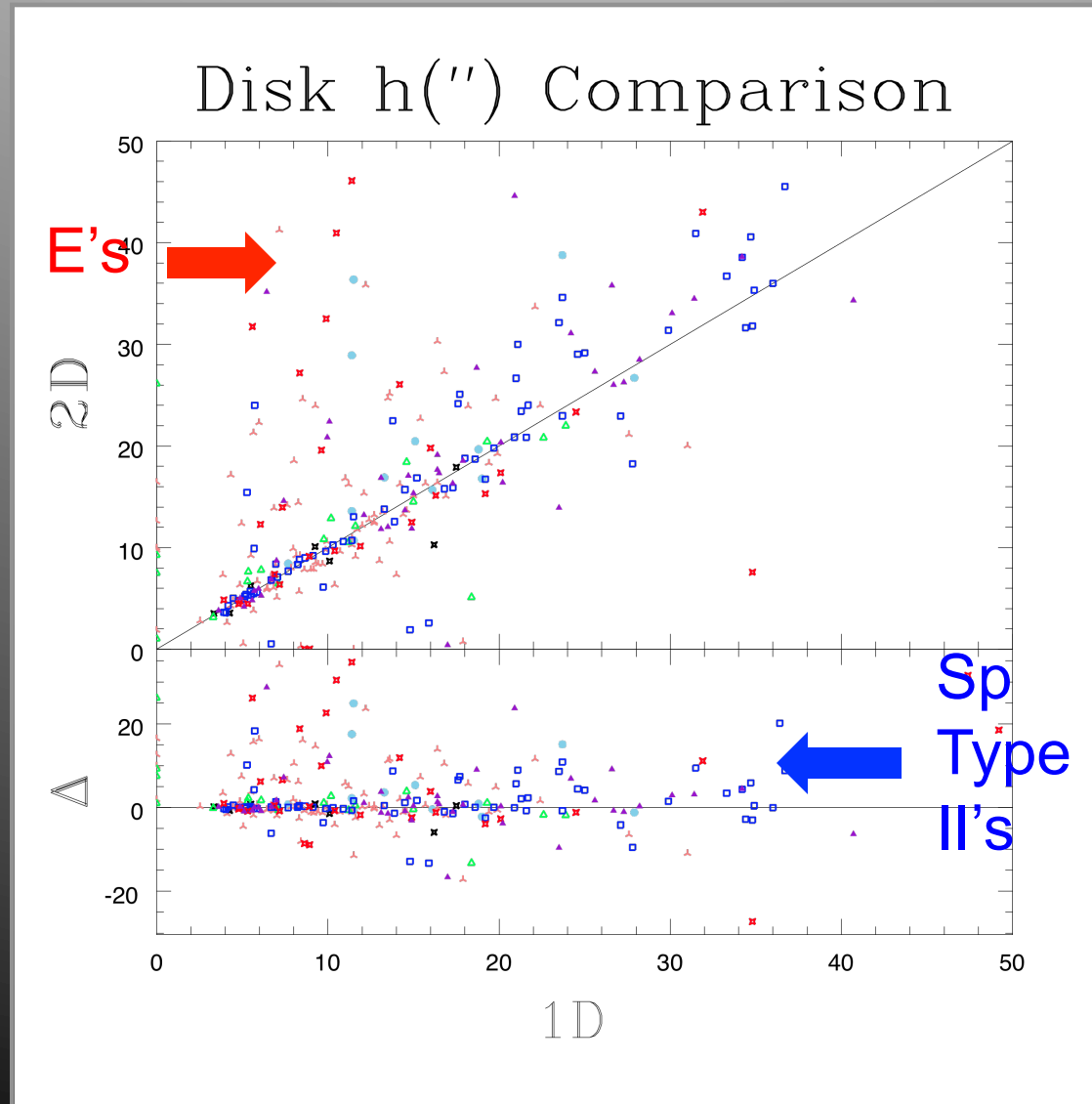
## Total magnitude (comparison)



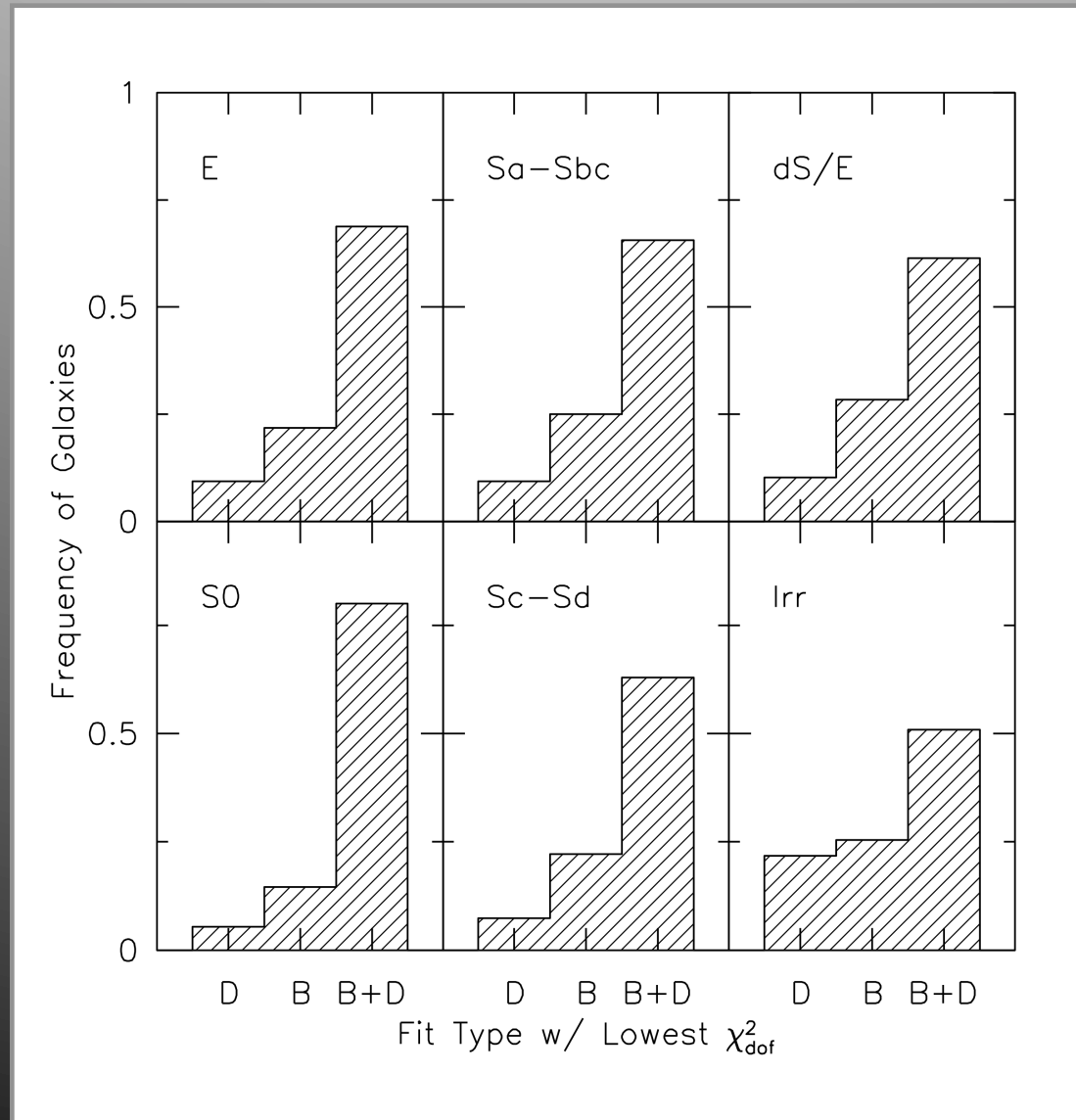
# 1D vs 2D for 786 VCC/SDSS galaxies (number reduces to ~300 gals for good GALFIT fits)



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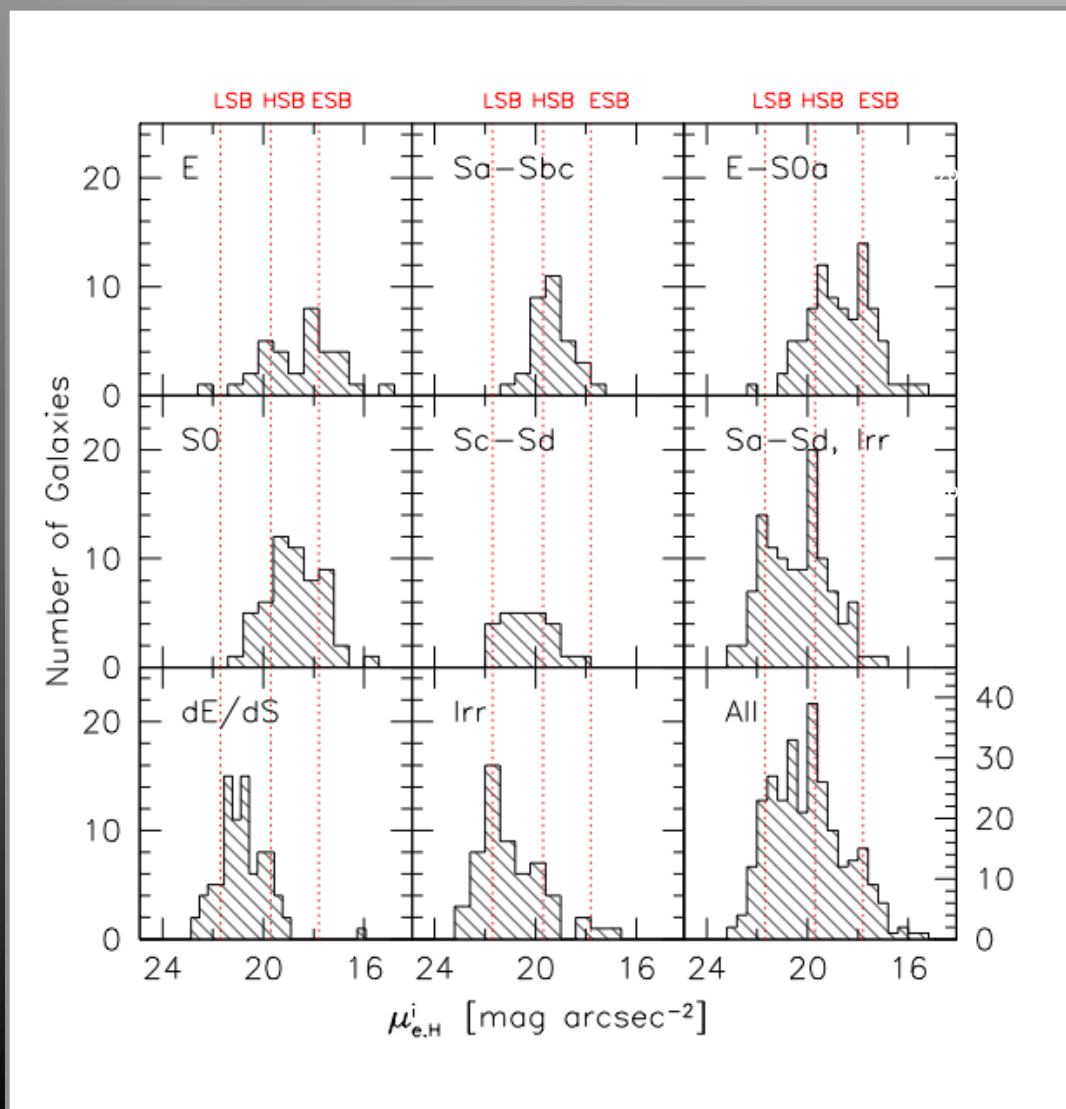


# 1D NLLS Decomps for 300 VCC galaxies



McDonald et al 2009; 2010

# Non-parametric analysis BIMODAL $\mu_e$ distribution!



- ❖ Model independent  $\mu_e$  for 286 VCC galaxies
  - HSB-ESB=1.61
  - LSB-HSB=1.61
- ❖ Most unexpected!
- ❖ McDonald, Courteau & Tully 2009, MNRAS

# Conclusions

- **Structural decomps of spiral galaxies accurate to ~20%**
- **2D decomps ideal for treatment of variable PA**
- **MCMC ideal for proper treatment of errors (GALFIT uses NLLS)**
  - **need 2D GALFIT/MCMC**



# Conclusions ...

- Halo significant below SDSS SB limit ( $27 \text{ mag arcsec}^{-2}$ ) → NGVS!
- Sky determination critical
- Decomps for  $n > 3$  (cross-talk with sky) and with multi-components (strong covariances) **AT RISK!**