

Neutrino Quantum Kinetic Equations - II

Vincenzo Cirigliano, LANL

George Fuller, UCSD

Alexey Vlasenko, UCSD

Based on 1309.2628, 1406.5558, 1406.6724, and references therein

Outline

Lectures

- Motivation: neutrinos and the cosmos
 - (I) ● Neutrinos in hot and dense media
 - Structure of QKEs from quantum field theory
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- Anatomy of the QKEs
 - Coherent evolution: flavor *and* spin
 - Inelastic collisions
 - (II) ● Comparison to other approaches & future challenges
- Talk by A. Vlasenko**
- Neutrino-antineutrino transformation in astrophysical environments

Structure of the QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$\begin{aligned} iDF &= [H, F] + iC \\ i\bar{D}\bar{F} &= [\bar{H}, \bar{F}] + i\bar{C} \end{aligned}$$

Derivative along v
world line:
drift & force term

“Vlasov”

Coherent evolution:
vacuum mass &
forward scattering
(refractive potential)

“MSW”

Inelastic collisions

“Boltzmann”

- F, H, C : $2n_f \times 2n_f$ matrices, all components coupled in general
- D, H, C are functionals of F, \bar{F} : non-linear system

Interlude on kinematics

- For ultra-relativistic V 's of 3-momentum \mathbf{p} , express all Lorentz tensors in terms of following basis:

$$n^\mu(p) = (1, \hat{p})$$

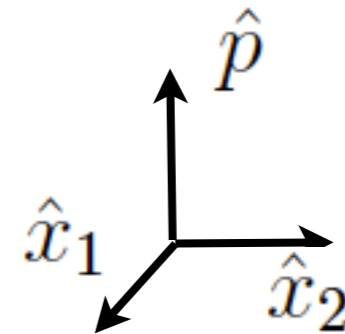
light-like

$$\bar{n}^\mu(p) = (1, -\hat{p})$$

light-like

$$x_{1,2}(p)$$

transverse



$$n \cdot n = \bar{n} \cdot \bar{n} = 0$$

$$n \cdot \bar{n} = 2$$

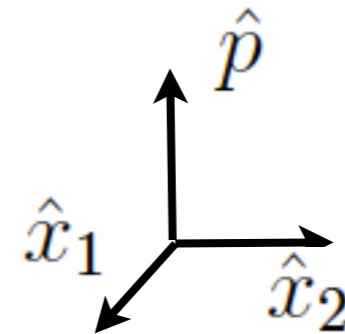
$$n \cdot x_i = \bar{n} \cdot x_i = 0$$

$$x_i \cdot x_j = -\delta_{ij}$$

Interlude on kinematics

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$\bar{n}^\mu(p) = (1, -\hat{p})$	light-like
$x_{1,2}(p)$	transverse



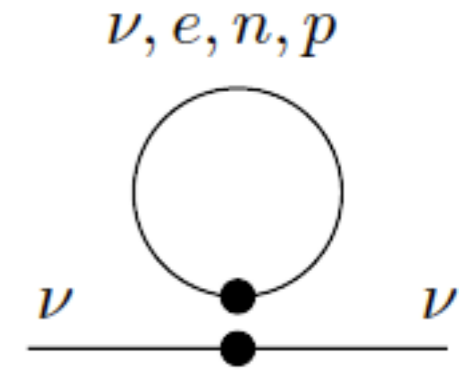
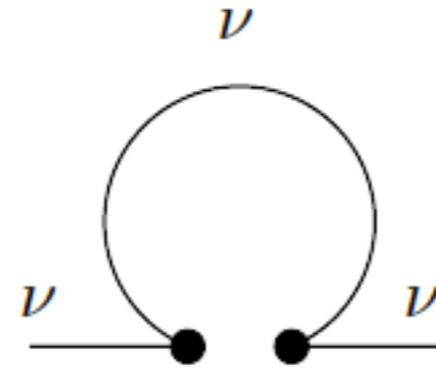
$$\begin{aligned}
 n \cdot n &= \bar{n} \cdot \bar{n} = 0 & n \cdot \bar{n} &= 2 \\
 n \cdot x_i &= \bar{n} \cdot x_i = 0 & x_i \cdot x_j &= -\delta_{ij}
 \end{aligned}$$

- Four-vector components along basis vectors:

$$\begin{aligned}
 V^\mu &\rightarrow V^\kappa \equiv n \cdot V & V^i &\equiv x_i \cdot V \\
 & & \partial^\kappa &\equiv n \cdot \partial & \partial^i &\equiv x_i \cdot \partial
 \end{aligned}$$

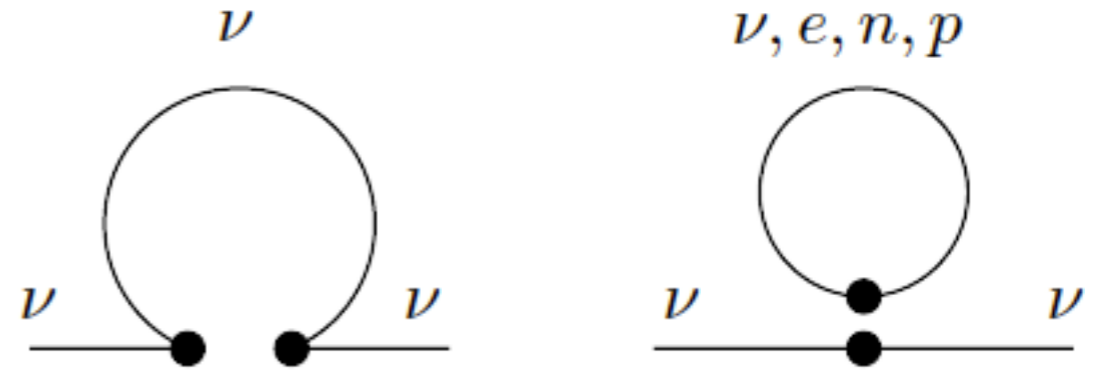
Interlude on potentials

- Neutrino self-energy diagrams \rightarrow in-medium *4-vector* potential (time- and space-like components in non-isotropic medium)



Interlude on potentials

- Neutrino self-energy diagrams \rightarrow in-medium 4-vector potential (time- and space-like components in non-isotropic medium)



- Computed from neutrino interactions in the Standard Model. Ex: neutrino-matter interaction at low-energy can be put in the form

$$\mathcal{L}_{\nu\psi} = -G_\psi \bar{\nu} \gamma_\mu P_L Y_\psi \nu \bar{\psi} \Gamma_\psi^\mu \psi$$

$$G_\psi = n_\psi \times G_F \sim g^2 / M_W^2$$

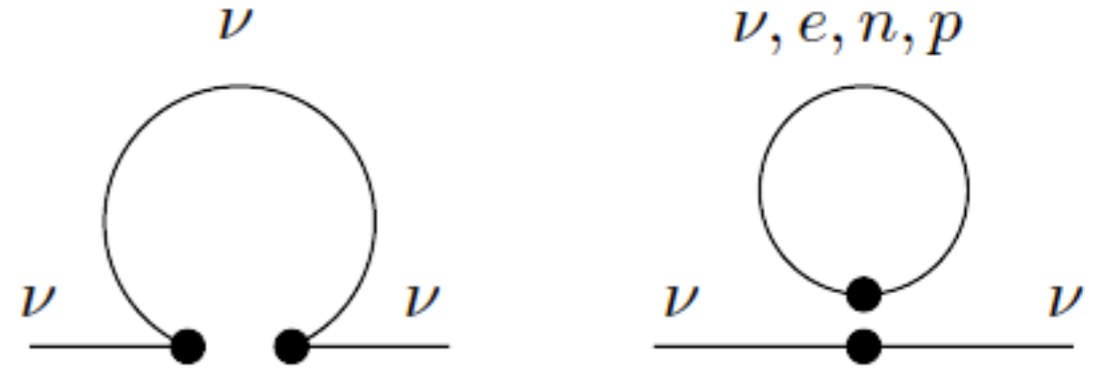
$$P_L = (1 - \gamma_5) / 2.$$

$$\Gamma_\psi^\mu = (g_V)_\psi \gamma^\mu - (g_A)_\psi \gamma^\mu \gamma^5.$$

$$\nu = \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} \quad Y_\psi = \begin{pmatrix} Y_{e\psi} & 0 & 0 \\ 0 & Y_{\mu\psi} & 0 \\ 0 & 0 & Y_{\tau\psi} \end{pmatrix}$$

Interlude on potentials

- Neutrino self-energy diagrams \rightarrow in-medium 4-vector potential (time- and space-like components in non-isotropic medium)



- $2n_f \times 2n_f$ matrix structure:

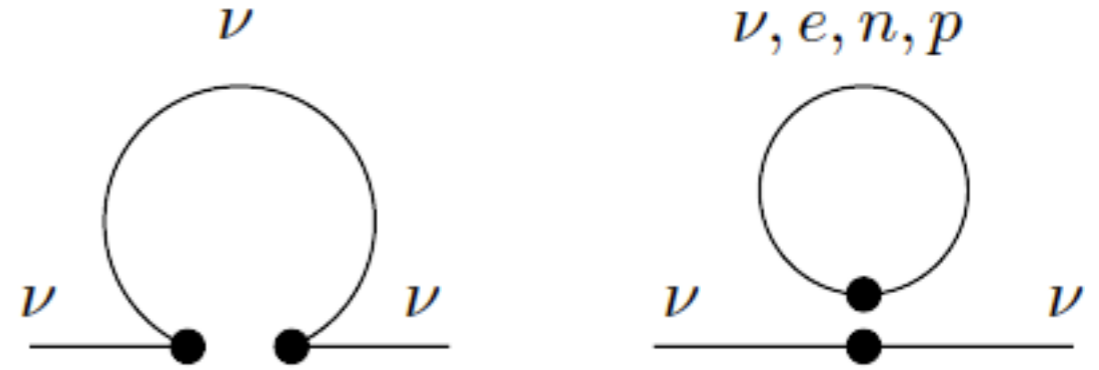
$$\Sigma^\mu(x) = \begin{pmatrix} \Sigma_R^\mu & 0 \\ 0 & \Sigma_L^\mu \end{pmatrix}$$

- Induced interaction

$$\mathcal{L}_{\text{int}} = -\bar{\nu}_L \not{A}_R \nu_L - \bar{\nu}_R \not{A}_L \nu_R + \text{h.c.}$$

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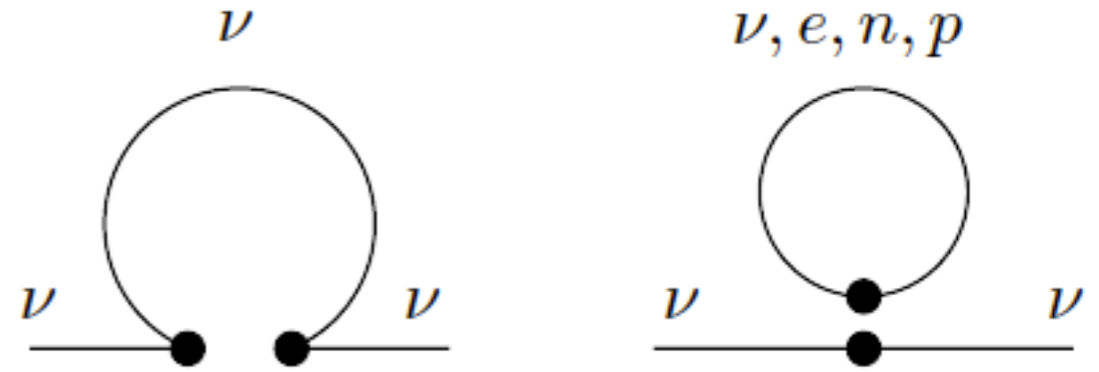
Potential for L-handed ν 's

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Potential for L-handed ν 's

Potential for R-handed ν 's:

Dirac: $\Sigma_L \propto G_F m^2 \sim O(\epsilon^3)$

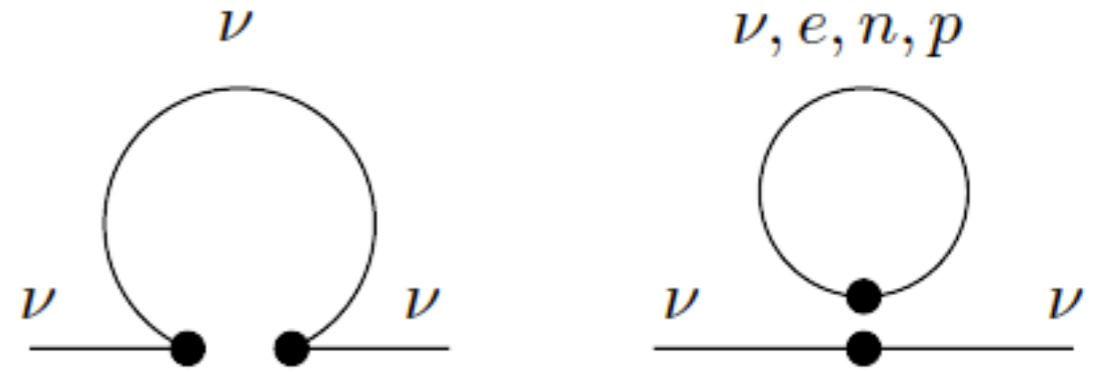
Majorana: $\Sigma_L = -\Sigma_R^T$

- Induced interaction

$$\mathcal{L}_{\text{int}} = -\bar{\nu}_L \not{\Sigma}_R \nu_L - \bar{\nu}_R \not{\Sigma}_L \nu_R + \text{h.c.}$$

Interlude on potentials

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Potential for L-handed ν 's

Potential for R-handed ν 's:

Dirac: $\Sigma_L \propto G_F m^2 \sim O(\epsilon^3)$

Majorana: $\Sigma_L = -\Sigma_R^T$

- For a test ν of momentum \mathbf{p} , get components

$$\Sigma^\kappa \equiv n(p) \cdot \Sigma$$

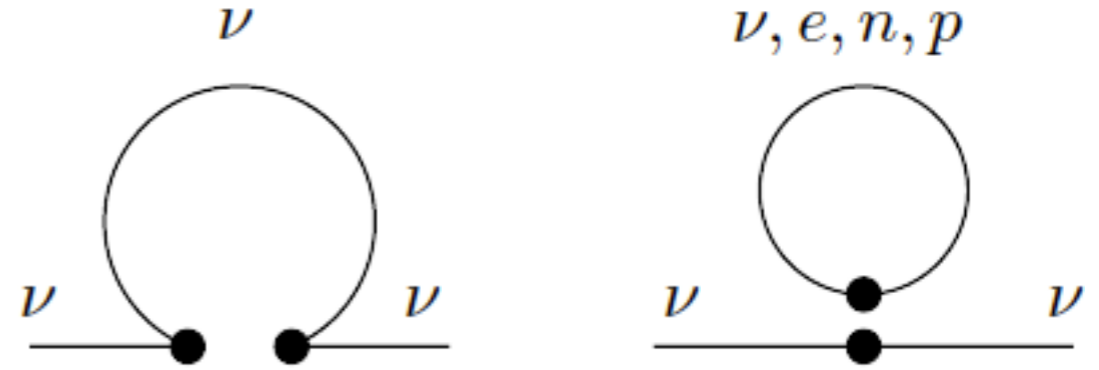
$$\Sigma^i \equiv x^i(p) \cdot \Sigma$$

approximately along ν trajectory

approximately transverse to ν trajectory

Interlude on potentials

- Neutrino self-energy diagrams \rightarrow in-medium 4-vector potential (time- and space-like components in non-isotropic medium)



- Explicit form of neutrino-induced Σ_R :

$$\Sigma_R^\mu \Big|_\nu = \sqrt{2}G_F \left(J_{(\nu)}^\mu + \mathbf{1} \left(\text{tr} J_{(\nu)}^\mu \right) \right)$$

$$J_{(\nu)}^\mu(x) = \int \frac{d^3q}{(2\pi)^3} n^\mu(q) \left(f_{LL}(\vec{q}, x) - \bar{f}_{RR}(\vec{q}, x) \right)$$

$$n^\mu(q) = (1, \hat{q})$$

Anatomy of the QKEs

Drift & force terms

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$iD F = [H, F] + iC$$

$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

Derivative along v
world line:
drift & force term

“Vlasov”

Drift & force terms

$$DF = \partial^\kappa F + \frac{1}{2|\vec{p}|} \{ \Sigma^i, \partial^i F \} - \frac{1}{2} \left\{ \frac{\partial \Sigma^\kappa}{\partial \vec{x}}, \frac{\partial F}{\partial \vec{p}} \right\}$$
$$\bar{D}\bar{F} = \partial^\kappa \bar{F} - \frac{1}{2|\vec{p}|} \{ \Sigma^i, \partial^i \bar{F} \} + \frac{1}{2} \left\{ \frac{\partial \Sigma^\kappa}{\partial \vec{x}}, \frac{\partial \bar{F}}{\partial \vec{p}} \right\}$$

- Simple interpretation if one notes that $v(+)$ and $\bar{v}(-)$ dispersion relations are:

$$\omega_\pm = |\vec{p}| \pm \Sigma^\kappa$$

$$\Sigma^\kappa \equiv n(p) \cdot \Sigma$$

Drift & force terms

$$DF = \partial^\kappa F + \frac{1}{2|\vec{p}|} \{ \Sigma^i, \partial^i F \} - \frac{1}{2} \left\{ \frac{\partial \Sigma^\kappa}{\partial \vec{x}}, \frac{\partial F}{\partial \vec{p}} \right\}$$

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- Simple interpretation if one notes that $v(+)$ and $\bar{v}(-)$ dispersion relations are:

$$\omega_\pm = |\vec{p}| \pm \Sigma^\kappa$$

$$\Sigma^\kappa \equiv n(p) \cdot \Sigma$$

- Then one finds:

$$D = \partial_t + \frac{1}{2} \{ \partial_{\vec{p}} \omega_+, \partial_{\vec{x}} \} - \frac{1}{2} \{ \partial_{\vec{x}} \omega_+, \partial_{\vec{p}} \}$$

$$\bar{D} = \partial_t + \frac{1}{2} \{ \partial_{\vec{p}} \omega_-, \partial_{\vec{x}} \} - \frac{1}{2} \{ \partial_{\vec{x}} \omega_-, \partial_{\vec{p}} \}$$

- Generalization of familiar

$$d_t = \partial_t + \dot{\vec{x}} \partial_{\vec{x}} + \dot{\vec{p}} \partial_{\vec{p}}$$

$$\dot{\vec{x}} = \partial_{\vec{p}} \omega \quad \dot{\vec{p}} = -\partial_{\vec{x}} \omega$$

Coherent evolution

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$iD F = [H, F] + iC$$

$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

Coherent evolution:
vacuum mass &
forward scattering
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$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

- Often written in the equivalent form of a Schrodinger-like equation for “ ν flavor wave-function”
- Mapping of the two approaches: off-diagonal entries in f_{LL} encode information about relative QM phases
- Not clear how to include inelastic collisions in wave-function approach

$$i\partial_t \psi = H \psi$$

$$\psi = \begin{pmatrix} \psi_e \\ \psi_\mu \\ \psi_\tau \end{pmatrix}$$

$$f_{LL}^{\alpha\beta} = \psi_\alpha \psi_\beta^*$$

Coherent evolution

- Controlled by $2n_f \times 2n_f$ Hamiltonian-like operators

$$H = \begin{pmatrix} H_R & H_m \\ H_m^\dagger & H_L \end{pmatrix} \quad \bar{H} = \begin{pmatrix} \bar{H}_R & H_m \\ H_m^\dagger & \bar{H}_L \end{pmatrix}$$

$$\bar{H}_R = \Sigma_R^\kappa \pm \frac{1}{2|\vec{p}|} \left(m^\dagger m - \epsilon^{ij} \partial^i \Sigma_R^j + 4\Sigma_R^+ \Sigma_R^- \right)$$

$$\bar{H}_L = \Sigma_L^\kappa \pm \frac{1}{2|\vec{p}|} \left(m m^\dagger + \epsilon^{ij} \partial^i \Sigma_L^j + 4\Sigma_L^- \Sigma_L^+ \right)$$

Coherent evolution

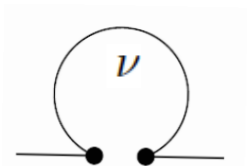
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Standard vacuum mass
term + medium refraction
(included in all analyses)



$$\longrightarrow \Sigma_R^\kappa(x) = \sqrt{2}G_F \int \frac{d^3q}{(2\pi)^3} n(p) \cdot n(q) (f_{LL}(\vec{q}, x) - \bar{f}_{RR}(\vec{q}, x))$$

$1 - \cos \theta_{\hat{p}\hat{q}}$

Coherent evolution

- Controlled by $2n_f \times 2n_f$ Hamiltonian-like operators

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$$\bar{H}_L = \Sigma_L^\kappa \pm \frac{1}{2|\vec{p}|} \left(m m^\dagger + \epsilon^{ij} \partial^i \Sigma_L^j + 4\Sigma_L^- \Sigma_L^+ \right)$$

Standard vacuum mass term + medium refraction (included in all analyses)

Additional $O(\epsilon^2)$ terms if potential has space-like components

$$\Sigma_{L,R}^\pm \equiv 1/2 (\Sigma_{L,R}^1 \pm i\Sigma_{L,R}^2)$$

Coherent evolution

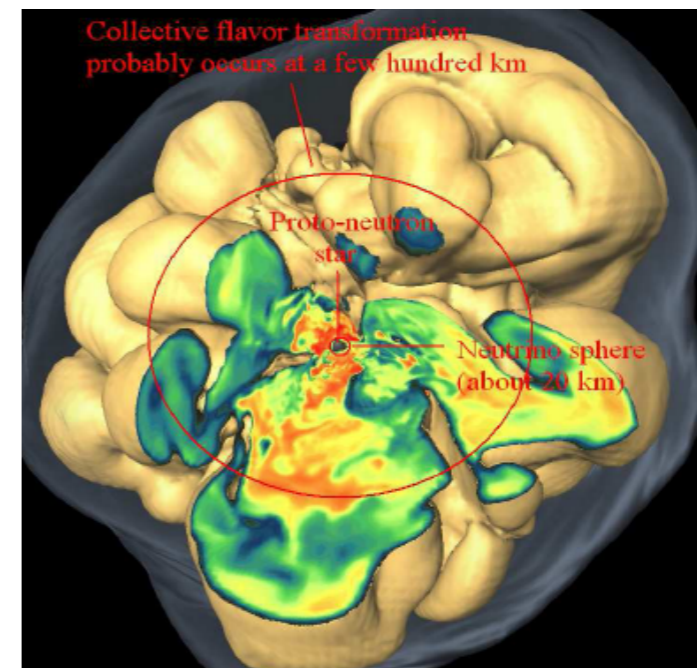
- Controlled by $2n_f \times 2n_f$ Hamiltonian-like operators

$$H = \begin{pmatrix} H_R & H_m \\ H_m^\dagger & H_L \end{pmatrix} \quad \bar{H} = \begin{pmatrix} \bar{H}_R & H_m \\ H_m^\dagger & \bar{H}_L \end{pmatrix}$$

$$H_m = -\frac{1}{|\vec{p}|} (\Sigma_R^+ m^\dagger - m^\dagger \Sigma_L^+)$$

- Qualitatively new $O(\varepsilon^2)$ effect: coherent conversion of LH \leftrightarrow RH ν 's

- Need anisotropic environment (transverse component of Σ)
- Need axial components, coupling to spin [I-flavor $H_m \sim m_\nu/p (\Sigma_R - \Sigma_L)^+$]
- Potentially big impact: Dirac (active-sterile) vs Majorana ($\nu-\bar{\nu}$)



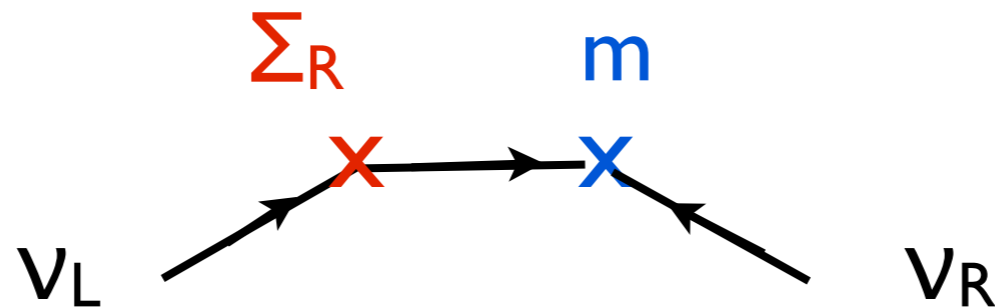
More on spin-mixing term

- Effect can be derived using effective hamiltonian approach

$$\langle i, \vec{p}', h' | j, \vec{p}, h \rangle \equiv -i(2\pi)^2 2|\vec{p}| \delta^{(4)}(p - p') \mathcal{H}_{h'h}^{ij}(p)$$

- Use medium-modified neutrino Lagrangian in perturbation theory

$$\mathcal{L}_{\text{int}} = -\bar{\nu}_L m \nu_R - \bar{\nu}_L \not{\Sigma}_R \nu_L - \bar{\nu}_R \not{\Sigma}_L \nu_R + \text{h.c.}$$



More on spin-mixing term

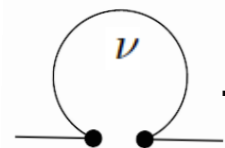
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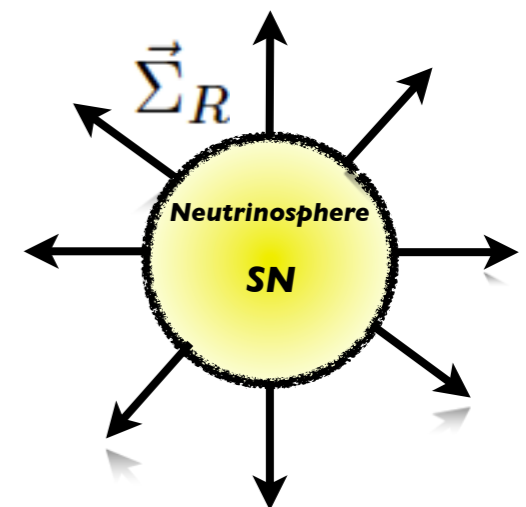
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* $\Sigma_{L,R}$: medium-induced vector potentials



$$\Sigma_R^\mu(x) = \sqrt{2}G_F \int \frac{d^3q}{(2\pi)^3} n^\mu(q) \left(f_{LL}(\vec{q}, x) - \bar{f}_{RR}(\vec{q}, x) \right)$$

* Even in simple “bulb” model for SN: $\vec{\Sigma}_R(x) \neq 0$.



More on spin-mixing term

- Effect can be derived using effective hamiltonian approach

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- 1-flavor result

$$\mathcal{H}_{LL}(p) = \Sigma_R^0 - \vec{\Sigma}_R \cdot \hat{p}$$

$$\mathcal{H}_{RR}(p) = \Sigma_L^0 - \vec{\Sigma}_L \cdot \hat{p}$$

$$\mathcal{H}_{LR}(p) = -\frac{m}{2|\vec{p}|} \vec{\Sigma}_A \cdot \vec{x}_+(p)$$

Axial potential

$$\Sigma_A^\mu \equiv \Sigma_L^\mu - \Sigma_R^\mu$$

↓

medium birefringence

+

mixing (transverse part)

More on spin-mixing term

- Effect can be derived using effective hamiltonian approach

$$\langle i, \vec{p}', h' | j, \vec{p}, h \rangle \equiv -i(2\pi)^2 2|\vec{p}| \delta^{(4)}(p - p') \mathcal{H}_{h'h}^{ij}(p)$$

- Similar mixing is induced by **magnetic moment** (Dirac for simplicity)

$$\Delta\mathcal{L} = (\mu_\nu/2) \bar{\nu}_R \sigma_{\mu\nu} F^{\mu\nu} \nu_L + \text{h.c.}$$

- 1-flavor result

$$\mathcal{H}_{LR}(p) = \mu_\nu \vec{B} \cdot \vec{x}_+(p)$$

← transverse component
of the magnetic field

See de Gouvea & Shalgar for impact on SN neutrino collective oscillations

Inelastic collisions

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

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Inelastic collisions

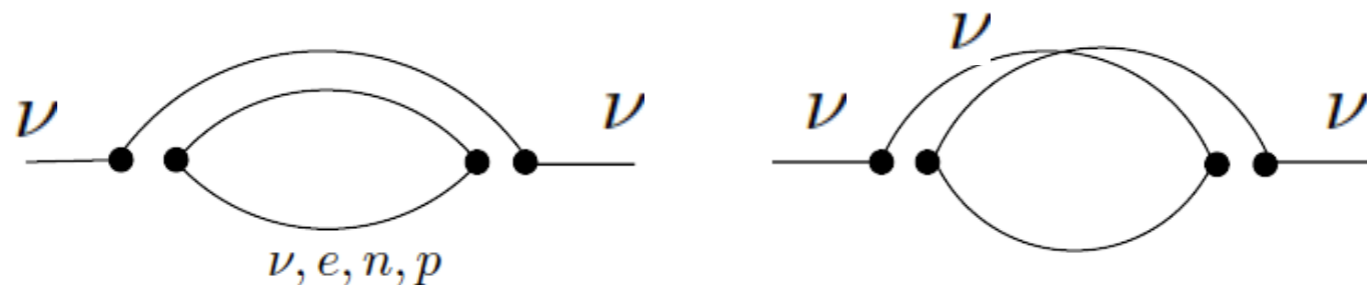
“Boltzmann”

Inelastic collisions

- Controlled by $2n_f \times 2n_f$ gain and loss potentials $\Pi^\pm[F, \bar{F}, f_{e,n,p,..}]$

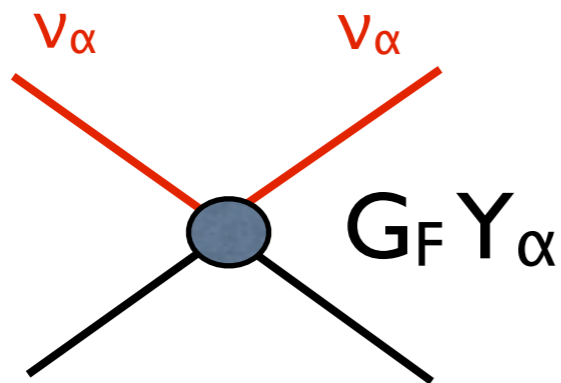
$$C = \frac{1}{2} \{ \Pi^+, F \} - \frac{1}{2} \{ \Pi^-, I - F \}$$

$$\bar{C} = \frac{1}{2} \{ \bar{\Pi}^+, \bar{F} \} - \frac{1}{2} \{ \bar{\Pi}^-, I - \bar{F} \}$$



- Π^\pm are **non-diagonal** in both flavor and spin (\rightarrow decoherence)

- Example: C_{LL} (upper $n_f \times n_f$ block) induced by neutrino scattering off medium particles (e,p,n,...) in isotropic environment

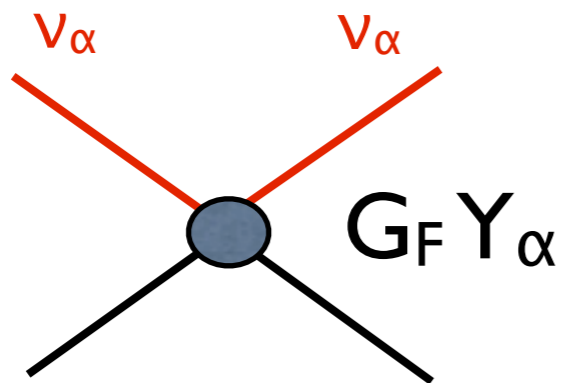


$$Y = \begin{pmatrix} Y_e & 0 & 0 \\ 0 & Y_\mu & 0 \\ 0 & 0 & Y_\tau \end{pmatrix}$$

Medium response function (knows about medium particle distributions and their interactions)

$$C_{LL}(\vec{p}) = - \int d^3p' R(\vec{p}, \vec{p}') \left\{ Y \left(1 - f_{LL}(\vec{p}') \right) Y, f_{LL}(\vec{p}) \right\} + \text{gain} :$$

- Example: C_{LL} (upper $n_f \times n_f$ block) induced by neutrino scattering off medium particles (e,p,n,...) in isotropic environment



$$Y = \begin{pmatrix} Y_e & 0 & 0 \\ 0 & Y_\mu & 0 \\ 0 & 0 & Y_\tau \end{pmatrix}$$

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$$\neq \begin{pmatrix} C_e & 0 & 0 \\ 0 & C_\mu & 0 \\ 0 & 0 & C_\tau \end{pmatrix}$$

Comparison with other QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$iDF = [H, F] + iC$$

NPB 406, 423 (1993)

- Restricting to f_{LL} and isotropic media, equivalent to Sigl-Raffelt

Comparison with other QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

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NPB 406, 423 (1993)

- Restricting to f_{LL} and isotropic media, equivalent to Sigl-Raffelt
- Similar in form to Strack-Burrows and Zhang-Burrows

$$\frac{\partial \mathcal{F}}{\partial t} + \vec{v} \cdot \frac{\partial \mathcal{F}}{\partial \vec{r}} + \dot{\vec{p}} \cdot \frac{\partial \mathcal{F}}{\partial \vec{p}} = -i[H, \mathcal{F}] + C$$

1310.2164
hep-ph/0504035

But $H, C, \dot{\vec{p}}$ are quite different

Comparison with other QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$iDF = [H, F] + iC$$

NPB 406, 423 (1993)

- Restricting to f_{LL} and isotropic media, equivalent to Sigl-Raffelt
- Similar in form to Strack-Burrows and Zhang-Burrows

$$\frac{\partial \mathcal{F}}{\partial t} + \vec{v} \cdot \frac{\partial \mathcal{F}}{\partial \vec{r}} + \dot{\vec{p}} \cdot \frac{\partial \mathcal{F}}{\partial \vec{p}} = -i[H, \mathcal{F}] + C$$

1310.2164
hep-ph/0504035

But $H, C, \dot{\vec{p}}$ are quite different

- Quite different from Volpe et al., who include “abnormal densities” (correlations of v and \bar{v} of opposite momentum) and discuss their coherent evolution coupled to “normal densities”.
We do not include this, based on $L_{\text{gradients}} \gg L_{\text{deBroglie}}$

1302.2347

Summary & future challenges

- Neutrino QKEs can be formulated from QFT + power counting in ratio of length scales ($L_{\text{osc}}, L_{\text{mfp}}, L_{\text{gradients}} \gg L_{\text{deBroglie}}$)
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- Challenges:
 - Explicit form of the collision term (in progress)
 - Computational implementation

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$$\begin{pmatrix} f & \cancel{\phi} \\ \cancel{\phi}^\dagger & \bar{f}^T \end{pmatrix}$$

Early Universe: $F(\mathbf{x}, \mathbf{p}) \rightarrow F(t, |\mathbf{p}|) \rightarrow F_{|\mathbf{p}|}(t)$

isotropy:
no L-R coherence

binning

$2 * (n_f)^2 * n_{|\mathbf{p}|}$ coupled ODEs, initial value problem

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Supernovae with spherical symmetry:

$$F(\mathbf{x}, \mathbf{p}) \xrightarrow{\text{geometry}} F(r, |\mathbf{p}|, \theta) \xrightarrow[\text{all } \theta \text{ contribute}]{\text{binning}} F_{|\mathbf{p}|, \theta}(r)$$

$4 * (n_f)^2 * n_{|\mathbf{p}|} * n_\theta$ coupled ODEs, boundary value problem

$$\begin{pmatrix} f & \phi \\ \phi^\dagger & \bar{f}^T \end{pmatrix}$$

