

NEUTRINO AND NUCLEAR ASTROPHYSICS

The 2014 International Summer School on AstroComputing, UCSD, July 21 - August 1 2014

Neutrino Quantum Kinetic Equations - I

Vincenzo Cirigliano, LANL

George Fuller, UCSD

Alexey Vlasenko, UCSD

Based on 1309.2628, 1406.5558, 1406.6724, and references therein

Outline

Lectures

- Motivation: neutrinos and the cosmos
 - (I) ● Neutrinos in hot and dense media
 - Structure of QKEs from quantum field theory

 - Anatomy of the QKEs
 - Coherent evolution: flavor *and* spin
 - Inelastic collisions
 - (II) ● Comparison to other approaches & future challenges
- Talk by A. Vlasenko**
- Neutrino-antineutrino transformation in astrophysical environments

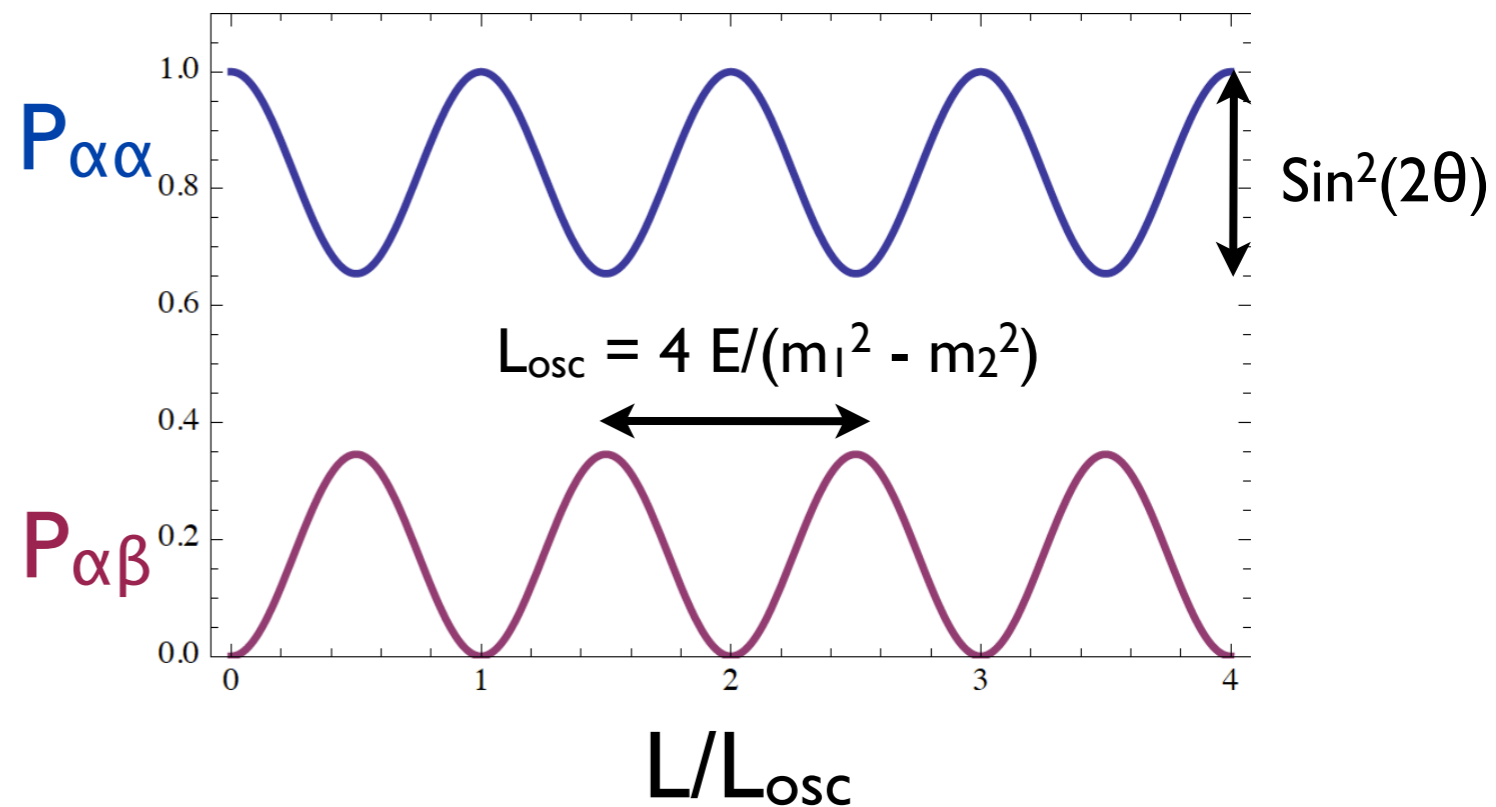
Neutrinos

- **Elusive particles:** lightest fermions, feel only the “weak” force

- Interaction (“flavor”) states $\nu_{e,\mu,\tau}$ do not coincide with mass states $\nu_{1,2,3}$

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

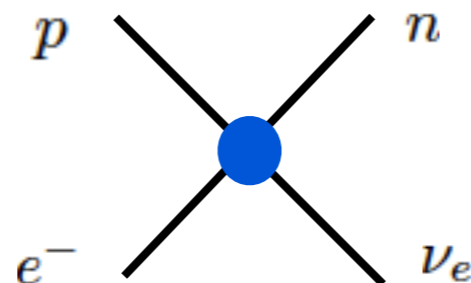
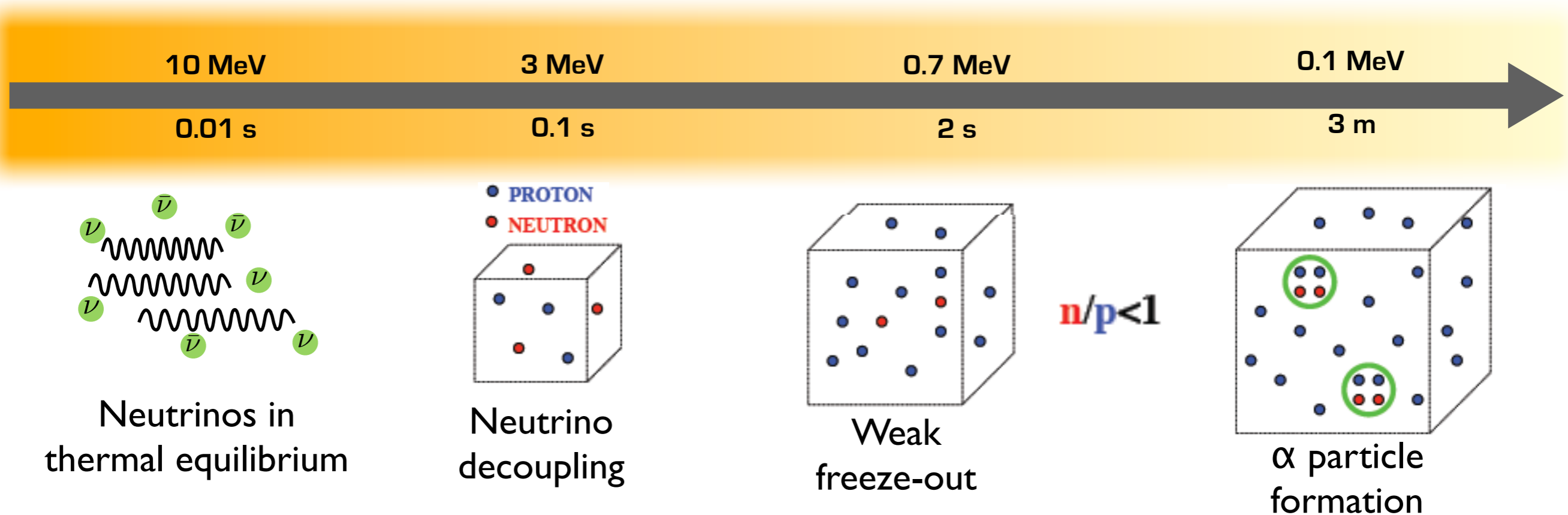
- A neutrino produced in a given flavor state can “**oscillate**” into another flavor state through QM interference effect!



Despite elusive nature, ν 's play a key role in cosmology / astrophysics

Neutrinos and the Cosmos (I)

- I. **What is the spectrum and flavor content of ν 's when they decouple in the Early Universe?** Far reaching implications for energy density, and n/p ratio \rightarrow **Big Bang Nucleosynthesis**



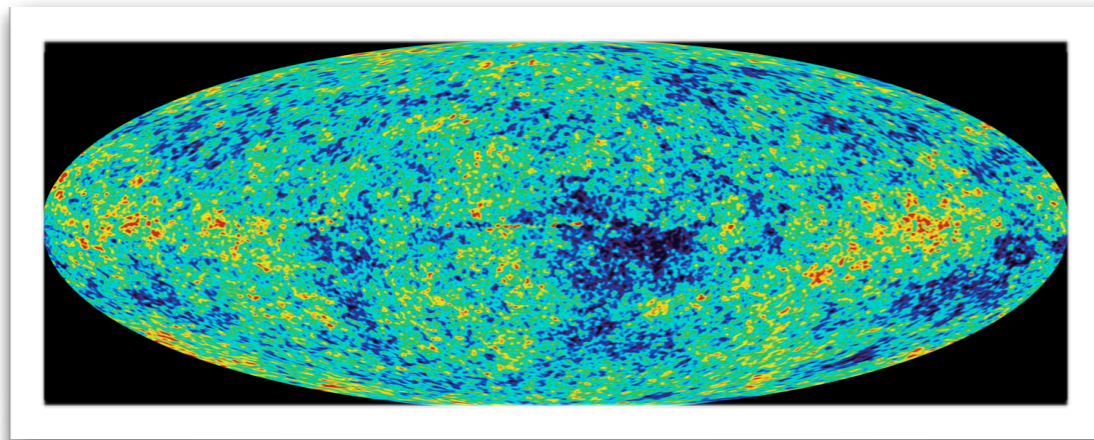
$$e^- + p \leftrightarrow \nu_e + n, \quad \bar{\nu}_e + p \leftrightarrow e^+ + n,$$

$$n \leftrightarrow p + e^- + \bar{\nu}_e$$

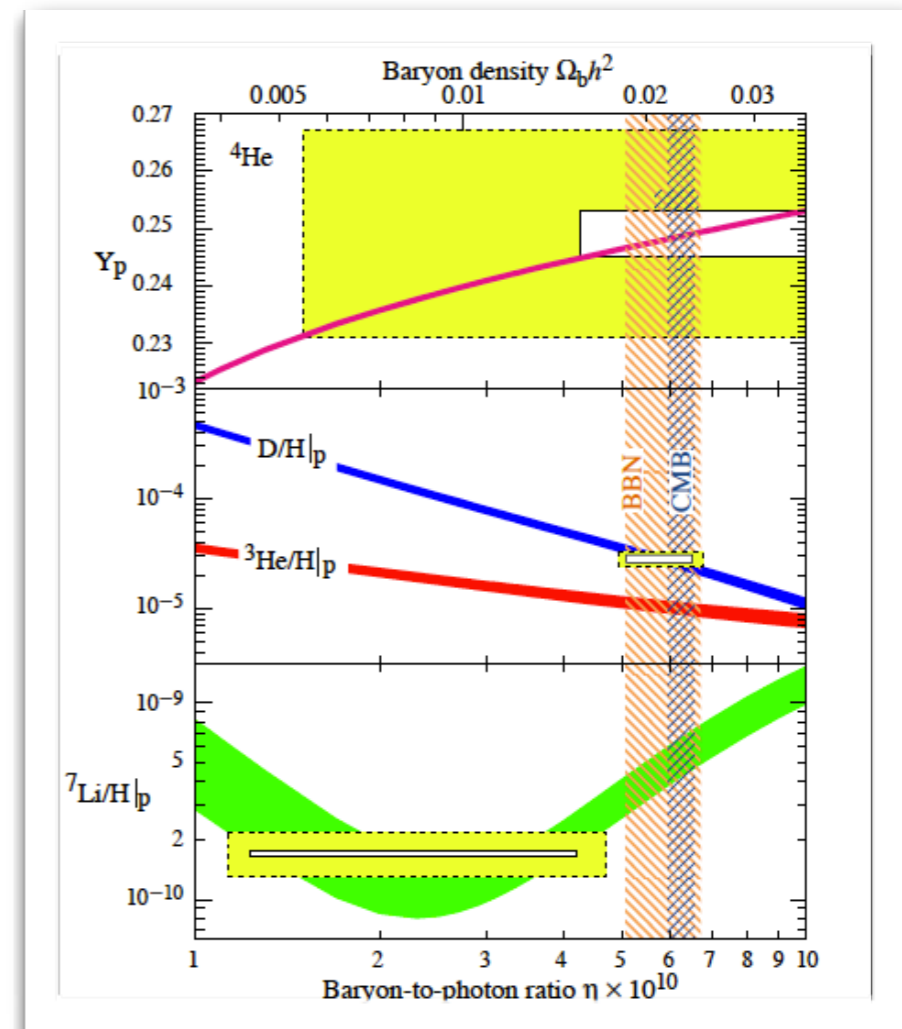
Reaction rates depend strongly on E_ν

Neutrinos and the Cosmos (I)

1. **What is the spectrum and flavor content of ν 's when they decouple in the Early Universe?** Far reaching implications for energy density, and n/p ratio \rightarrow **Big Bang Nucleosynthesis**

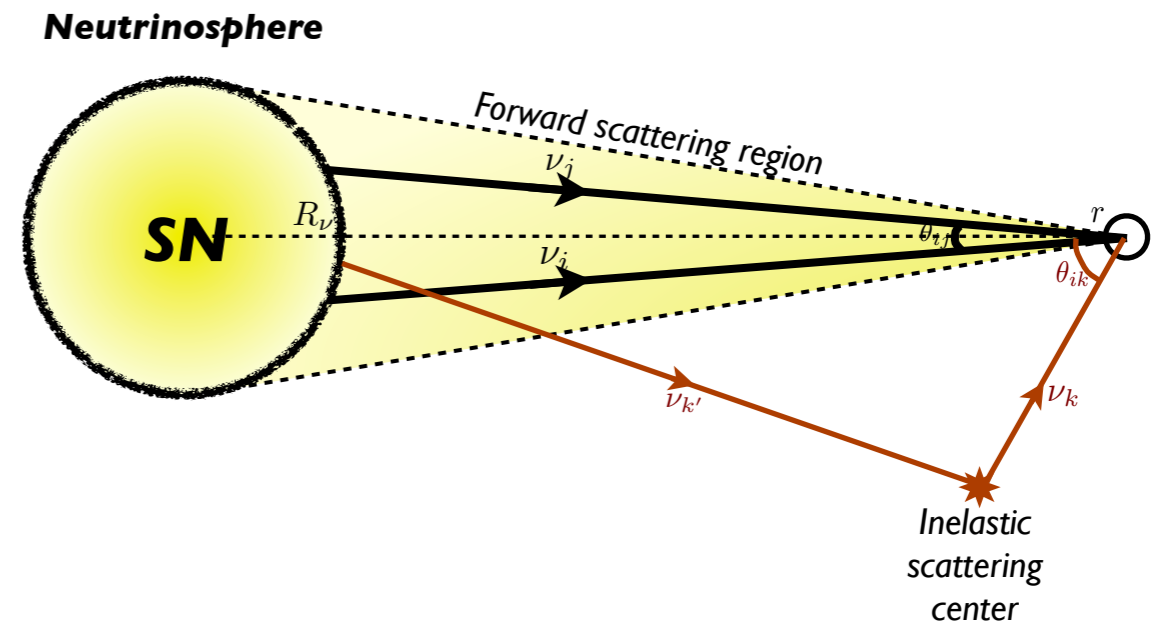


- Precise **observations** (η_B , N_{eff} , D , ${}^4\text{He}$) + **robust theory** can turn BBN into a deep probe of physics beyond the Standard Model in the lepton sector (sterile ν 's, non-zero L)

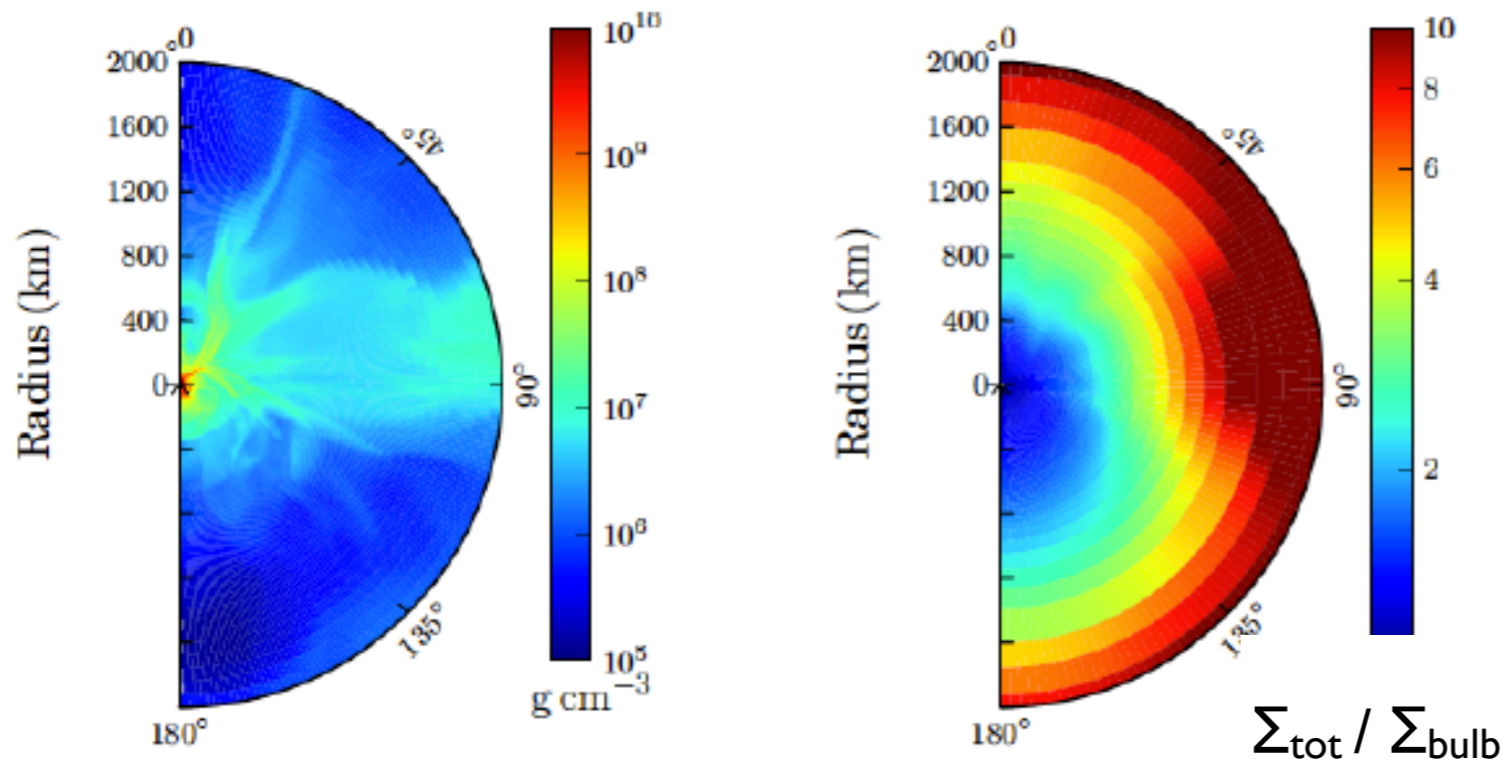


Neutrinos and the Cosmos (2)

2. What is the impact of inelastic collisions on ν propagation in the SN envelope? Implications for SN ν signal, nucleosynthesis in the neutrino-heated ejecta



Cherry-Carlson-Friedland-Fuller-Vlasenko 2012



First studies indicate that $< 1\%$ of ν scatter, but there is a large effect on the neutrino potential Σ (angular dependence)

$$\Sigma_{\nu\nu} \sim 1 - \cos \theta_{\hat{p}\hat{q}}$$

The need for QKEs

To fully address the issues described above, must set up the *analytic* and *computational* tools needed to describe neutrino kinetics in the EU and SN environments, simultaneously keeping track of the key quantum mechanical effect of **coherent flavor oscillations AND de-cohering inelastic collisions** with the medium

Neutrinos in hot / dense medium

- At a given time, ensemble of neutrinos described by incoherent mixture of states $|k\rangle$ with weight p_k ($\sum p_k = 1$)
- Physics controlled by density matrix

$$\rho = \sum_k p_k |k\rangle \langle k| \quad i \frac{d\rho}{dt} = [H, \rho]$$

Example: in thermal equilibrium

$$\rho_{\text{eq}} = \frac{e^{-H/(kT)}}{\text{Tr} (e^{-H/(kT)})}$$

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$$\rho = \sum_k p_k |k\rangle \langle k| \quad i \frac{d\rho}{dt} = [H, \rho]$$

- Ensemble average of any operator:

$$\langle \hat{O} \rangle = \sum_k p_k \langle k | \hat{O} | k \rangle = \text{Tr}(\rho \hat{O})$$

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- In EU and SN we need densities and fluxes of ν_α , $\alpha=e,\mu,\tau,X \Rightarrow$ generalized number operator

$$\hat{O} \leftrightarrow n_{ij} = a_j^\dagger a_i$$

creation / annihilation operators:
i,j label one-particle states

$$|\nu(i)\rangle = a_i^\dagger |0\rangle$$

- 1-particle states associated with massive spin-1/2 field

$$|\nu(\vec{p}, i, h)\rangle = a_{i,h}^\dagger(\vec{p})|0\rangle \quad |\bar{\nu}(\vec{p}, i, h)\rangle = b_{i,h}^\dagger(\vec{p})|0\rangle$$

creation operator for particle / antiparticle labeled by
3-momentum \mathbf{p} , mass m_i , helicity $h=L,R$

- Dirac \rightarrow 4 states: L- and R-handed neutrino *and* antineutrino
- Majorana \rightarrow 2 states: L- and R-handed neutrino ($\Psi = \Psi^c \Rightarrow a_i = b_i$)

- Key dynamical objects are the “matrices of densities”

$i = 1, 2, 3, \dots$

$h, h' = L, R$

neutrinos

$$\langle a_{j,h'}^\dagger(\vec{p}') a_{i,h}(\vec{p}) \rangle = (2\pi)^3 2n_{ij}(\vec{p}) \delta^{(3)}(\vec{p} - \vec{p}') f_{hh'}^{ij}(\vec{p})$$

creation / annihilation operators

normalization (conventional)

$$n_{ij} = 2\omega_i\omega_j / (\omega_i + \omega_j)$$

$$\{a_{i,h}(\vec{p}), a_{j,h'}^\dagger(\vec{p}')\} = (2\pi)^3 2\omega_i(\vec{p}) \delta_{hh'} \delta_{ij} \delta^{(3)}(\vec{p} - \vec{p}')$$

$$\omega_i(\vec{p}) = \sqrt{\vec{p}^2 + m_i^2}$$

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anti-
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$$\langle b_{i,h'}^\dagger(\vec{p}') b_{j,h}(\vec{p}) \rangle = (2\pi)^3 2n_{ij}(\vec{p}) \delta^{(3)}(\vec{p} - \vec{p}') \bar{f}_{hh'}^{ij}(\vec{p})$$

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- Physical content:

$f_{hh}^{ii}(\vec{p})$ Represents occupation number of neutrinos of mass m_i , helicity h , momentum p

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$f_{hh}^{ii}(\vec{p})$ Represents occupation number of neutrinos of mass m_i , helicity h , momentum p

$f_{hh}^{ij}(\vec{p})$ Signals quantum coherence between states of **same helicity and different mass**

Non-zero if there are states in the ensemble that are coherent superpositions of states of same helicity and different mass, e.g., L-handed neutrino flavor states

$$\begin{pmatrix} |\nu_\alpha\rangle \\ |\nu_\beta\rangle \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \end{pmatrix}$$

$$\Rightarrow f_{LL}^{12} \propto \sin \theta \cos \theta$$

- Key dynamical objects are the “matrices of densities”

$$i = 1, 2, 3, \dots \quad h, h' = L, R$$

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$$f_{hh'}^{ii}(\vec{p})$$

Signals quantum coherence between states of **same mass and different helicity**

$$f_{hh'}^{ij}(\vec{p})$$

Signals quantum coherence between states of **different mass and different helicity**

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- $2n_f \times 2n_f$ matrix structure: Dirac case, need F and \bar{F}

$$F(\vec{p}, x) = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F}(\vec{p}, x) = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

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$n_f \times n_f$ blocks describing
matrix of density for active
states (L-handed neutrinos
and R-handed antineutrinos)

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$n_f \times n_f$ blocks describing
L-R (active-sterile)
coherence

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- $2n_f \times 2n_f$ matrix structure: Majorana case

$$a_i(\vec{p}, h) = b_i(\vec{p}, h) \quad \longrightarrow \quad f \equiv f_{LL}, \quad \bar{f} \equiv \bar{f}_{RR} = f_{RR}^T \quad \phi = f_{LR}$$

- Key dynamical objects are the “matrices of densities”

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$n_f \times n_f$ blocks describing
matrix of density for
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$n_f \times n_f$ blocks describing matrix of density for neutrinos and antineutrinos

$n_f \times n_f$ block describing L-R (neutrino-antineutrino) coherence

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- QKEs are nothing but the evolution equations for the f 's
- We work in the flavor basis, related to the above by:

Matrix that puts neutrino propagator in diagonal form

$$\nu_\alpha = U_{\alpha i} \nu_i$$

$$f_{\alpha\beta} = U_{\alpha i} f_{ij} U_{\beta j}^*$$

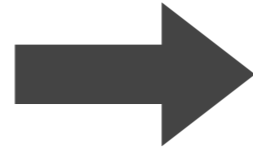
flavor basis

mass basis

QKEs from Quantum Field Theory

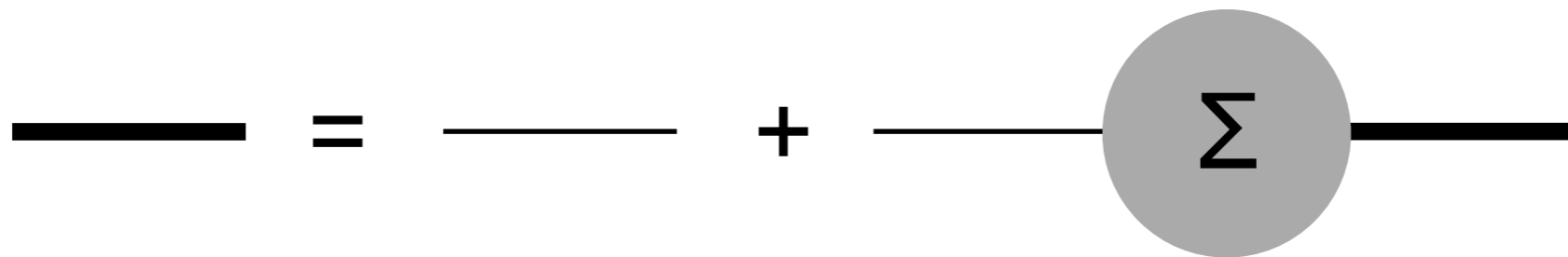
Equations of motion
for Green Functions

$$\langle \nu_{\alpha}^i(x) \bar{\nu}_{\beta}^j(y) \rangle$$



Kinetic equations for
“matrix of densities” $f(x,p)$

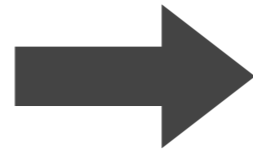
$$f_{\alpha\beta}^{\lambda\lambda'}(x,p) \sim \langle a_{\beta}^{\dagger}(p,\lambda') a_{\alpha}(p,\lambda) \rangle$$



QKEs from Quantum Field Theory

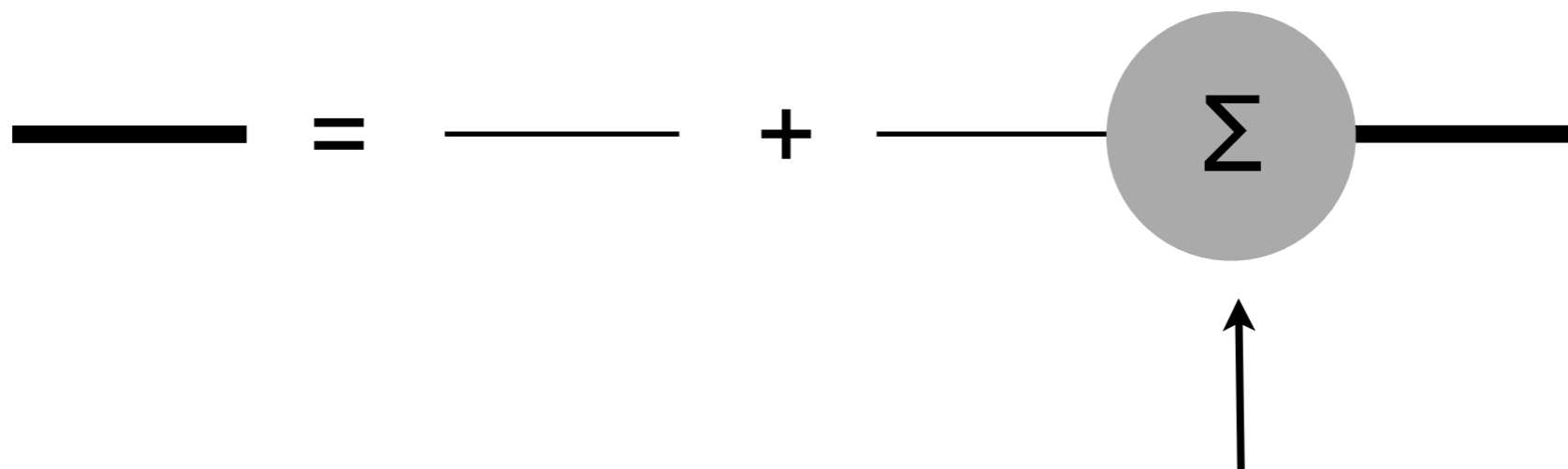
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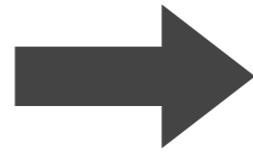
$$\Sigma_{\text{forward}} \sim G_F n$$

$$\Sigma_{\text{inelastic}} \sim (G_F)^2 T^5 ; (G_F)^2 n T^2$$

QKEs from Quantum Field Theory

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Kinetic equations for
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$$f_{\alpha\beta}^{\lambda\lambda'}(x,p) \sim \langle a_{\beta}^{\dagger}(p,\lambda') a_{\alpha}(p,\lambda) \rangle$$

- Exploit hierarchy of scales. Work to 2nd order in small ratios ($E \sim T$):

$$m_{\nu}/E \sim \Delta m_{\nu}/E \sim \Sigma_{\text{forward}}/E \sim \partial_{\chi}/E \sim O(\varepsilon)$$

$$\Sigma_{\text{inelastic}}/E \sim O(\varepsilon^2)$$

Small
neutrino
(Δ)masses

Comparable** potential induced
by forward scattering on matter
and other neutrinos

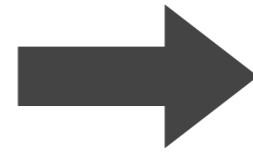
Slowly
varying
background

Weak
interaction
rates

QKEs from Quantum Field Theory

Equations of motion
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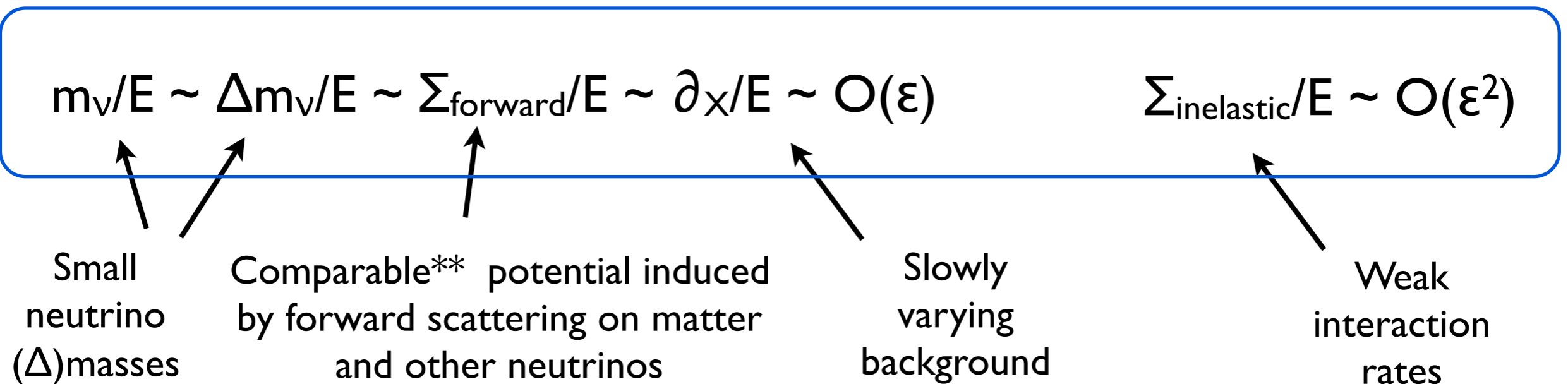
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Kinetic equations for
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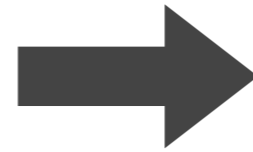


- The physics: $L_{\text{osc}} \sim E/(\Delta m_{\nu})^2$, L_{mfp} , $L_{\text{gradients}} \gg L_{\text{deBroglie}}$

QKEs from Quantum Field Theory

Equations of motion
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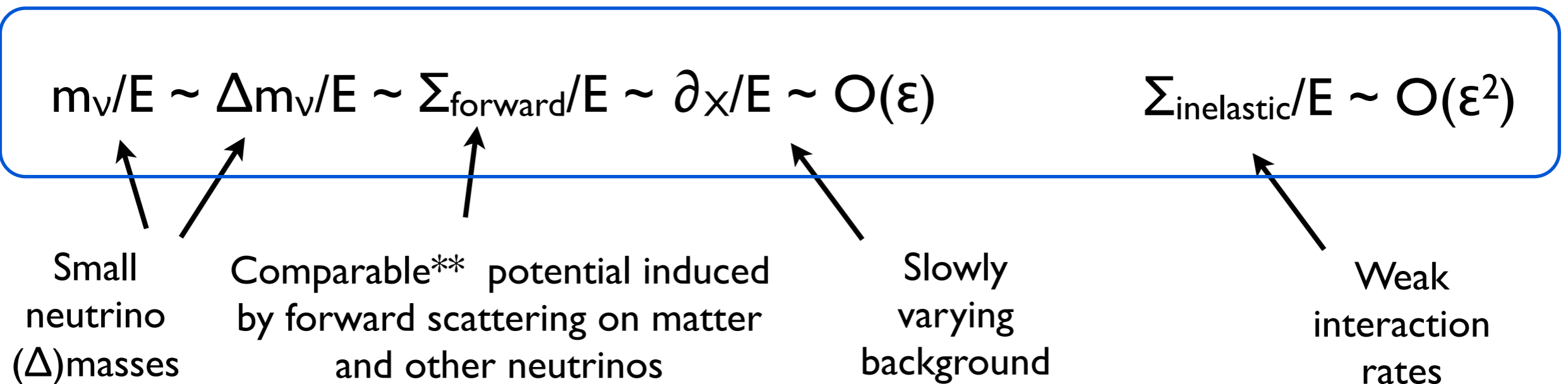
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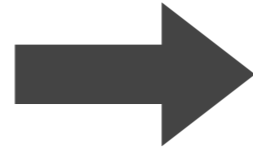


- Initial density matrix of the system [recall $\langle O \rangle = \text{Tr}(\rho O)$] \rightarrow initial (or boundary) conditions for the QKEs

QKEs from Quantum Field Theory

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Kinetic equations for
“matrix of densities” $f(x,p)$

$$f_{\alpha\beta}^{\lambda\lambda'}(x,p) \sim \langle a_{\beta}^{\dagger}(p,\lambda') a_{\alpha}(p,\lambda) \rangle$$

- Advantages of this approach (used already in other contexts, such as baryogenesis in the Early Universe):
 - First principles method, forced us to think about L-R coherence
 - No guesses or fudging: diagrammatic computations in non-eq QFT determine all terms of the QKEs
 - Systematic approximations (based on power counting in ϵ 's)

Structure of the QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$iD F = [H, F] + iC$$

$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

Structure of the QKEs

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$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

Derivative along v
world line:
drift & force term

“Vlasov”

Structure of the QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$iD F = [H, F] + iC$$

$$i\bar{D} \bar{F} = [\bar{H}, \bar{F}] + i\bar{C}$$

Derivative along v
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Coherent evolution:
vacuum mass &
forward scattering
(refractive potential)

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- F, H, C : $2n_f \times 2n_f$ matrices, all components coupled in general
- D, H, C are functionals of F, \bar{F} : non-linear system

Structure of the QKEs

$$F = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$
$$\bar{F} = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

$$\begin{aligned} iDF &= [H, F] + iC \\ i\bar{D}\bar{F} &= [\bar{H}, \bar{F}] + i\bar{C} \end{aligned}$$

Current state-of-the art:

- Early Universe: approximate treatment of inelastic collisions, inadequate in decoupling regime
- Supernovae:
 - no simultaneous treatment of forward AND inelastic collisions (separation of low- and high-density regimes)
 - no inclusion of spin degrees of freedom ($n_f \times n_f$ problem)

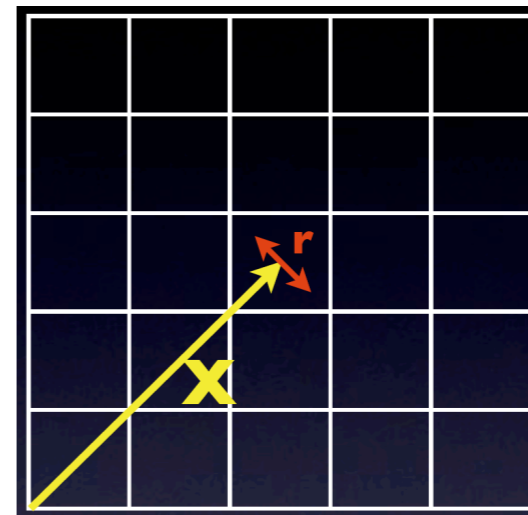
Backup

Green's function approach

- Dynamics contained in the two-point function

$$G^{ij}(x, r) \equiv \frac{1}{2} \langle [\nu^i(x + r/2), \bar{\nu}^j(x - r/2)] \rangle$$

Wigner transform



$$G^{ij}(x, p) \equiv \int dr e^{ip \cdot x} G^{ij}(x, r)$$

- Take spinor projections (vector, tensor)

$$F_{L,R} = \frac{1}{4} \text{Tr} \left(\gamma_\mu P_{L,R} G(p, x) \right) \bar{n}^\mu$$

$$\Phi^{(\dagger)} = \mp \frac{i}{16} \text{Tr} \left(\sigma_{\mu\nu} P_{L/R} G(p, x) \right) (\bar{n}^\mu x_\pm^\nu - \bar{n}^\nu x_\pm^\mu)$$

$$P_{L,R} \equiv (1 \mp \gamma_5)/2, \quad x_\pm \equiv x_1 \pm ix_2,$$

- Collet into $2n_f \times 2n_f$ matrix $\hat{F} = \begin{pmatrix} F_L & \Phi \\ \Phi^\dagger & F_R \end{pmatrix}$

- Take frequency projections

$$-2 \int_0^\infty \frac{dp^0}{2\pi} \hat{F}(p, x) = F(\vec{p}, x) = \begin{pmatrix} f_{LL} & f_{LR} \\ f_{RL} & f_{RR} \end{pmatrix}$$

$$-2 \int_{-\infty}^0 \frac{dp^0}{2\pi} \hat{F}(p, x) = \bar{F}(-\vec{p}, x) = \begin{pmatrix} \bar{f}_{RR} & \bar{f}_{RL} \\ \bar{f}_{LR} & \bar{f}_{LL} \end{pmatrix}$$

↖
In free theory coincide
with definition in
terms of creation and
annihilation operators