

Giant Molecular Cloud Evolution

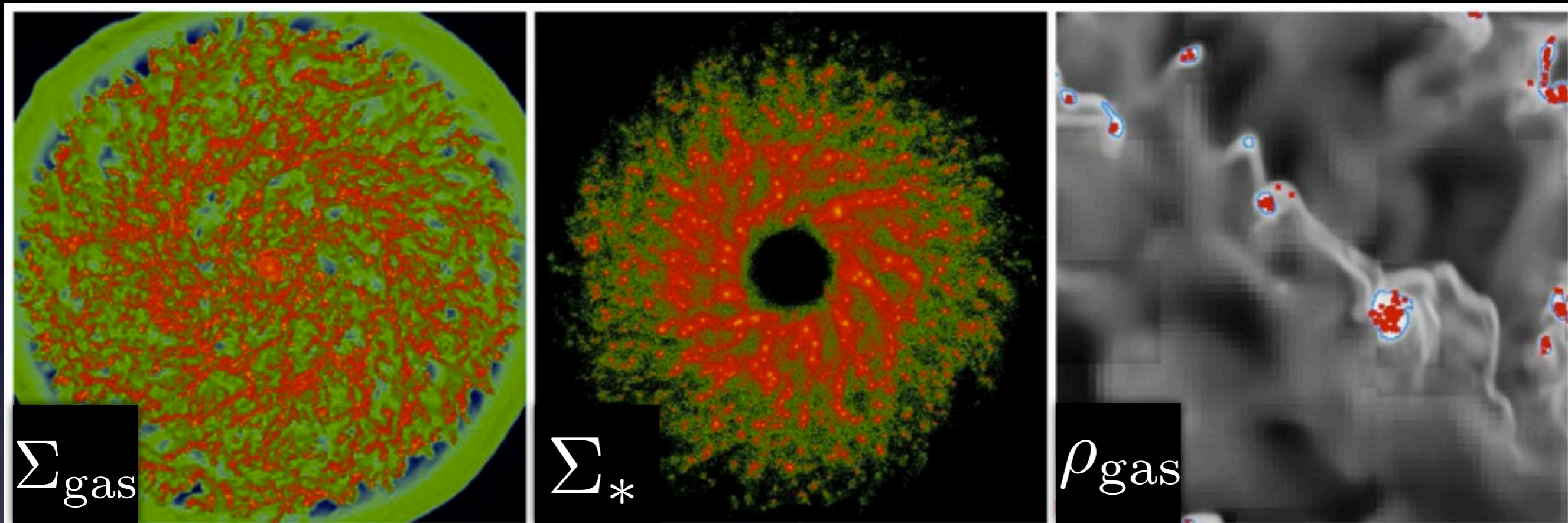
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How is **star formation** modeled in galaxy evolution simulations?



Cold, Dense Gas $\xrightarrow{\epsilon_{\text{ff}}}$ Stars

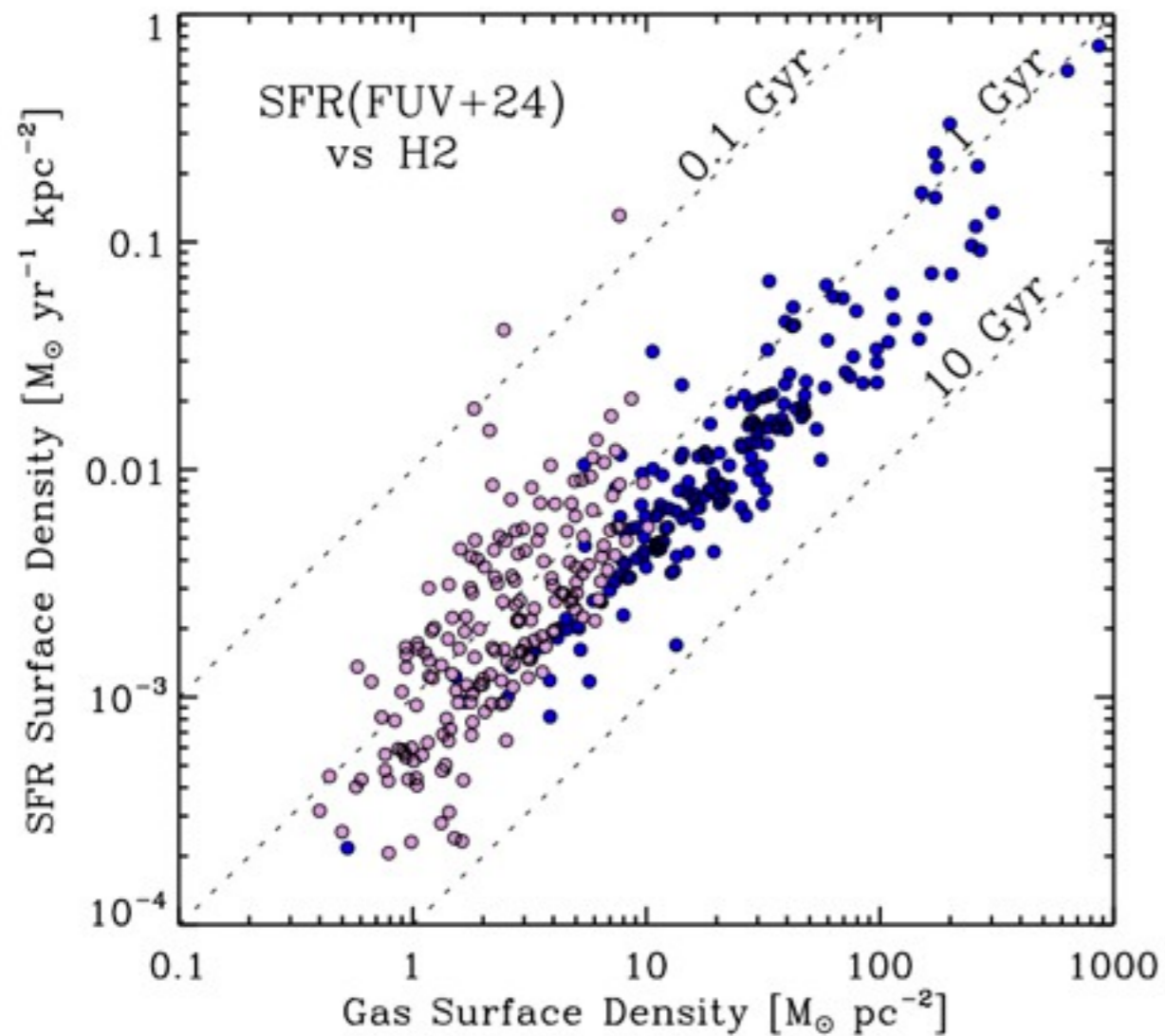
We can do better!

And maybe learn something about
molecular clouds along the way...

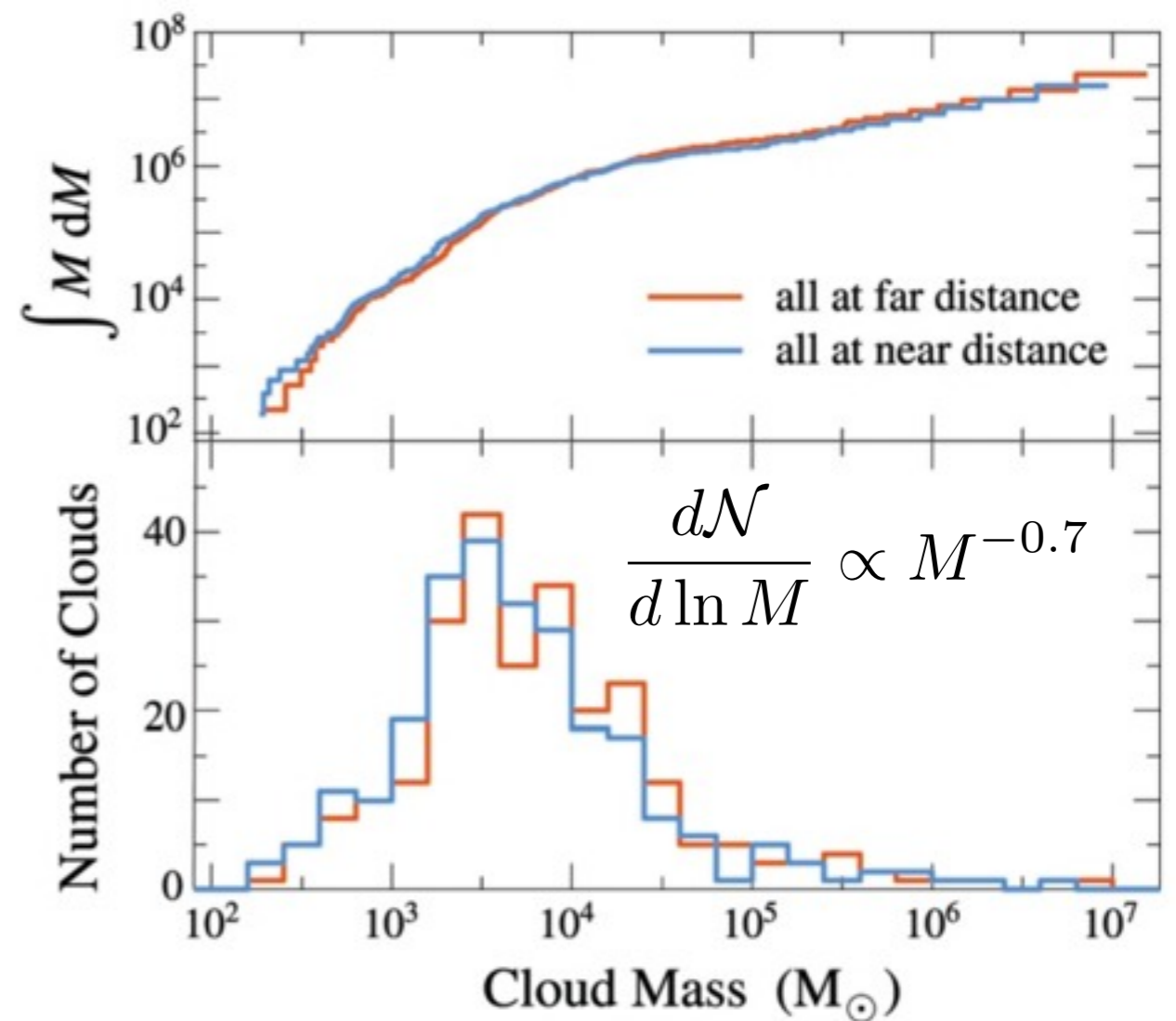
Outline

- Background on Giant Molecular Clouds
- `GMCEvo1`: A semianalytic model for molecular cloud evolution
- Comparison to Observations of GMCs
- A novel subgrid prescription for star formation in galaxy evolution simulations

Giant molecular clouds are the primary sites of star formation in the Milky Way

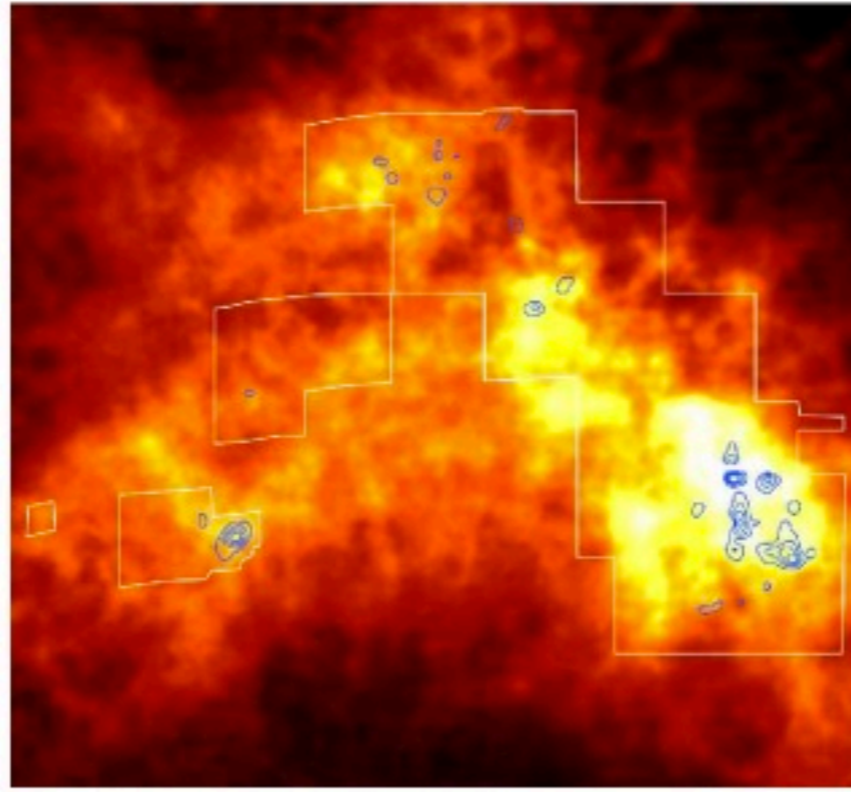
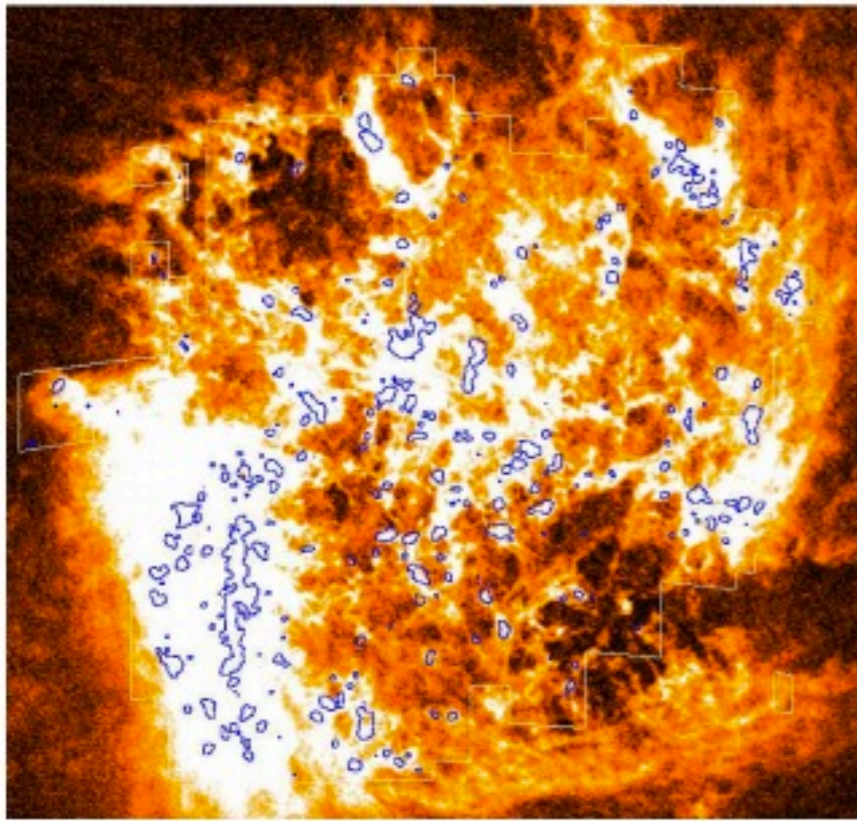


Schruba+ 2011



Stark & Lee 2006

GMCs are ubiquitously observed in quiescently star forming galaxies

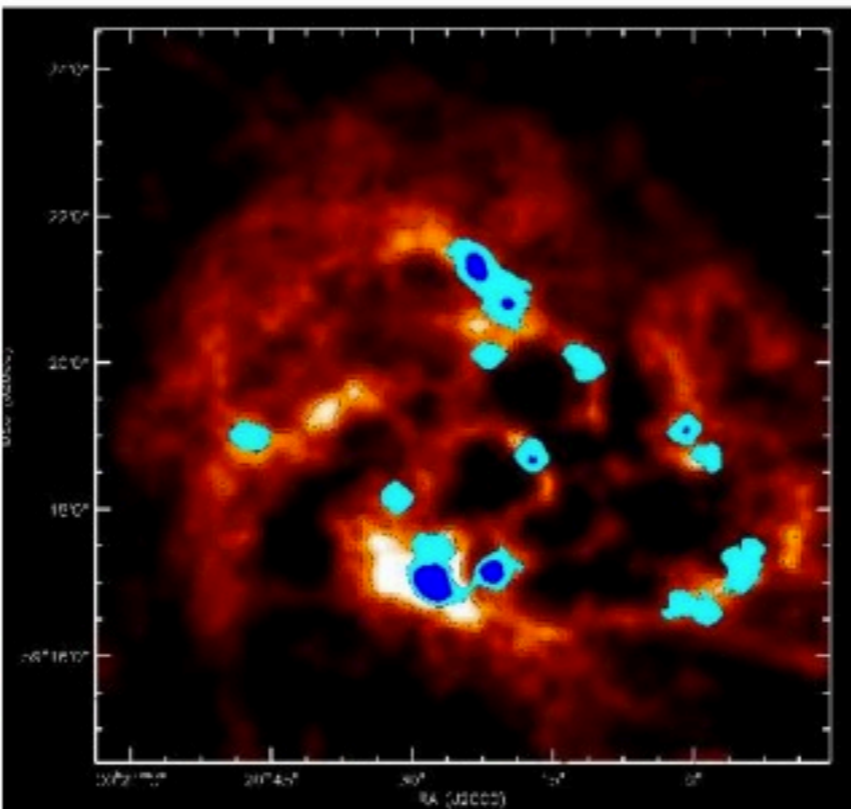
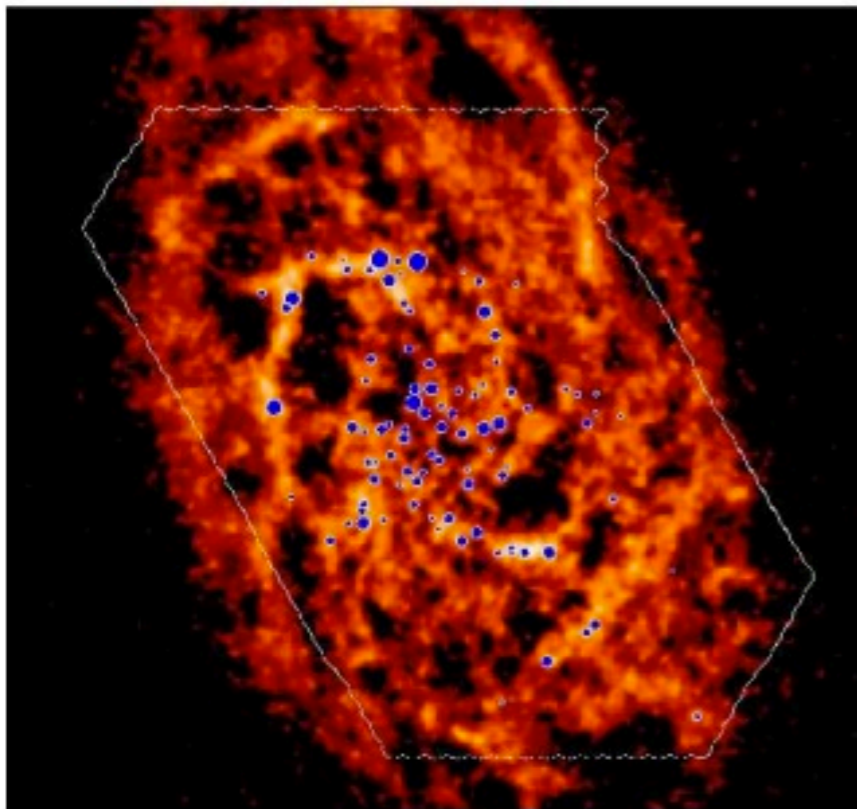


Can measure
global properties:

Mass: M_{cl}

Radius: R_{cl}

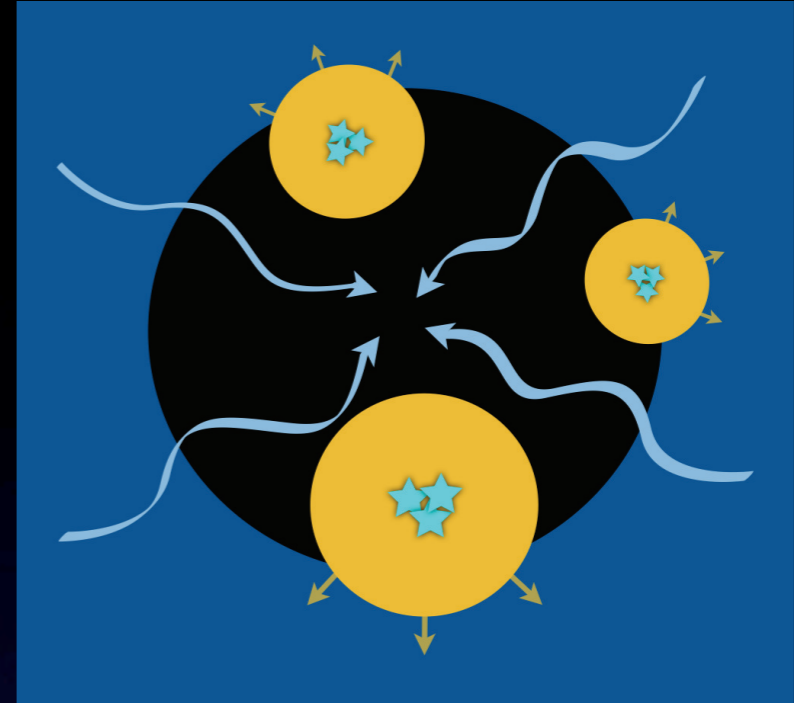
Velocity
Dispersion: σ_v



Blitz+ 2007

GMCEvol

A semianalytic model for
giant molecular cloud evolution



$$\frac{1}{2}\ddot{I}_{\text{cl}} = 2(\mathcal{T} - \mathcal{T}_0) + \mathcal{B} + \mathcal{W} - \frac{1}{2}\frac{d}{dt}\int_{S_{\text{vir}}}(\rho\mathbf{v}r^2)\cdot d\mathbf{S} + \int_{V_{\text{vir}}}\mathbf{r}\cdot\mathbf{F}_b dV$$

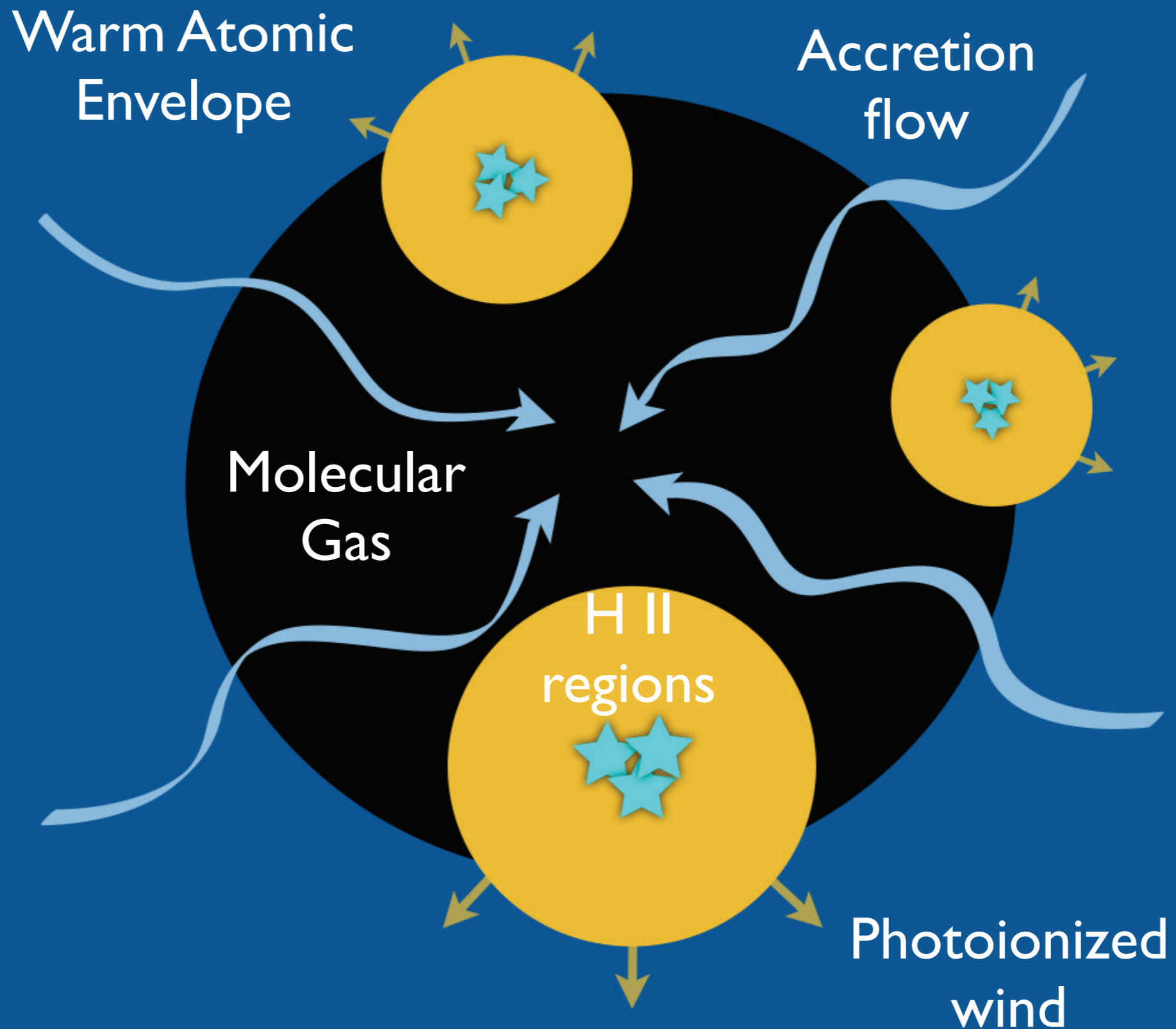
$$\frac{d\mathcal{E}_{\text{cl}}}{dt} + \int_{S_{\text{vir}}}\left[\rho\mathbf{v}\left(\frac{1}{2}v^2 + e + \phi + \frac{P}{\rho}\right) + \mathbf{S}_P\right]\cdot d\mathbf{S} = \int_{V_{\text{vir}}}\mathbf{v}\cdot\mathbf{F}_b dV - \Lambda$$

$$\dot{M}_{\text{cl}} = \dot{M}_{\text{ej}} + \dot{M}_{\text{acc}}$$

$$\mathbf{F}_b = \mathbf{F}_* + \mathbf{F}_w + \mathbf{F}_{\text{acc}}$$

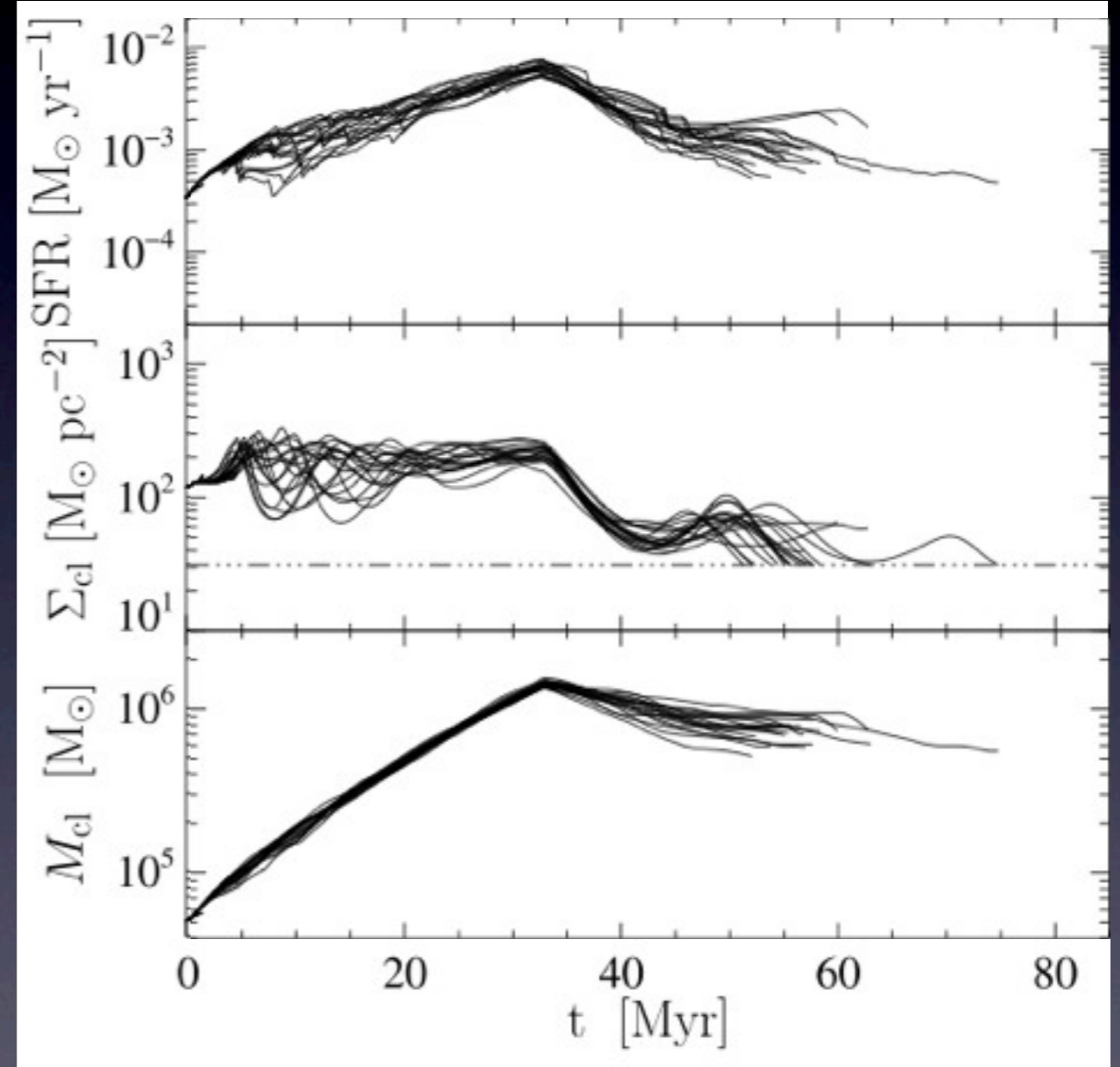
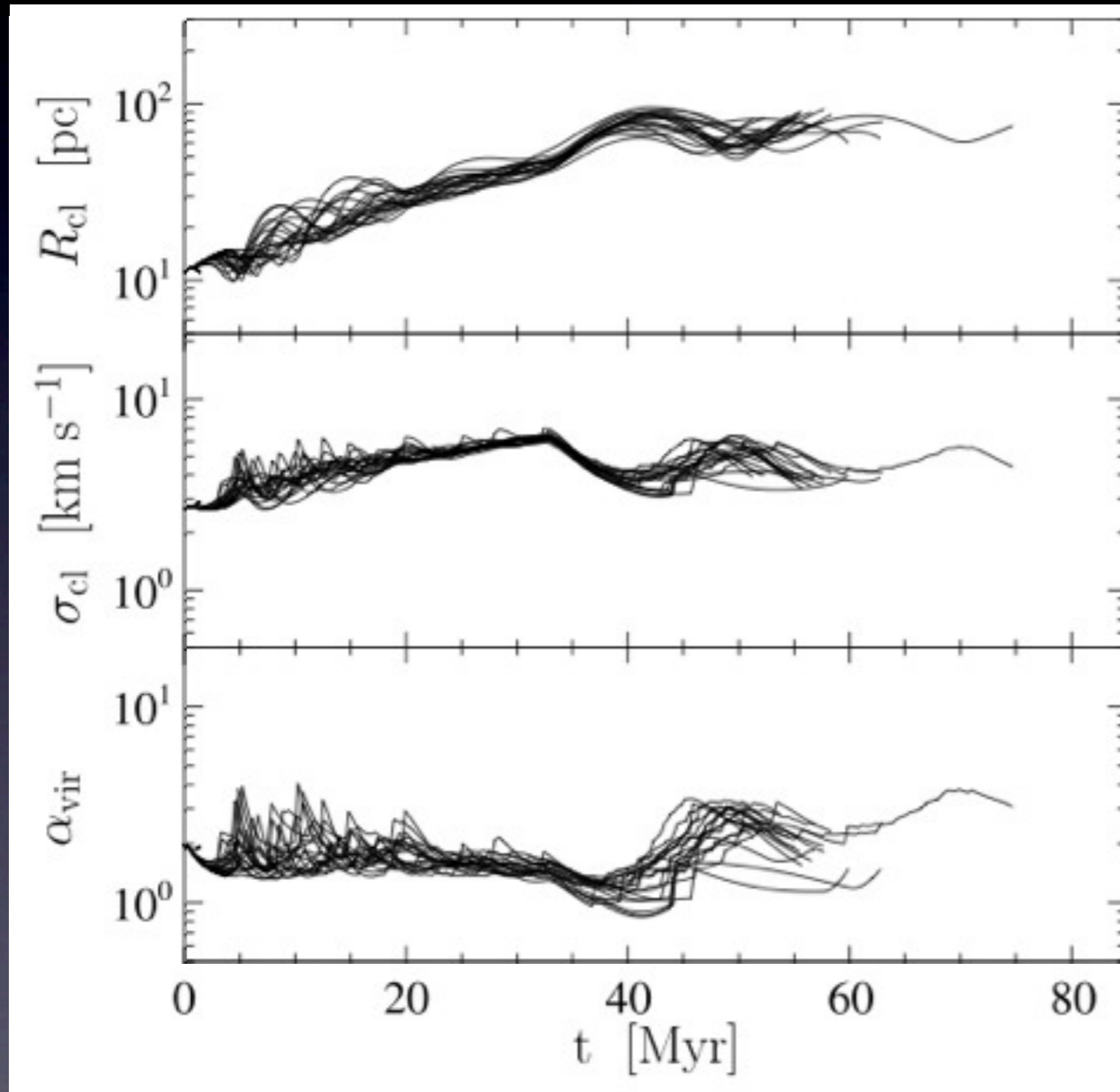
Follow evolution of: M_{gas} , M_{star} , R_{cl} , dR_{cl}/dt , σ_{cl}

Components of the GMC model



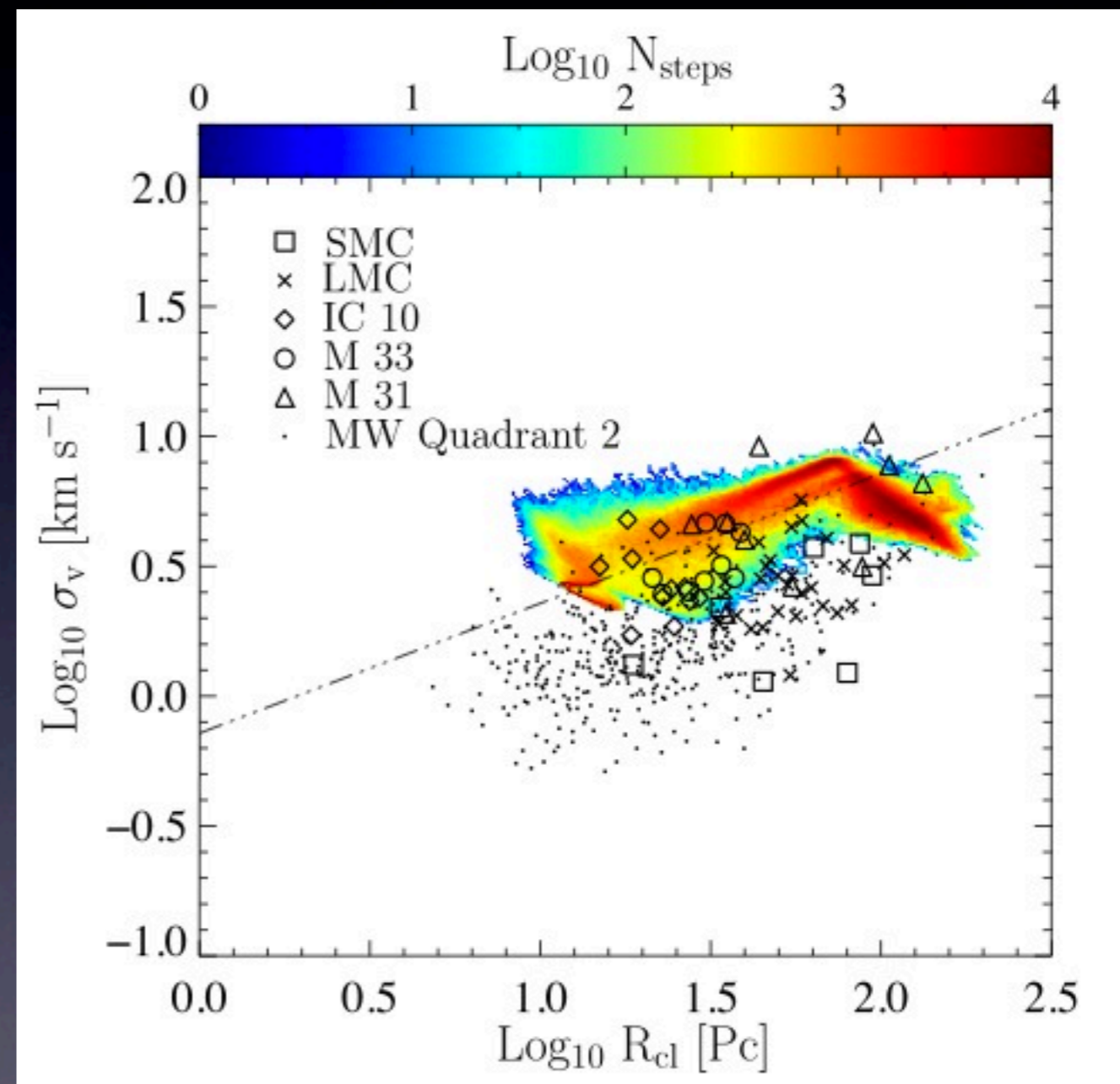
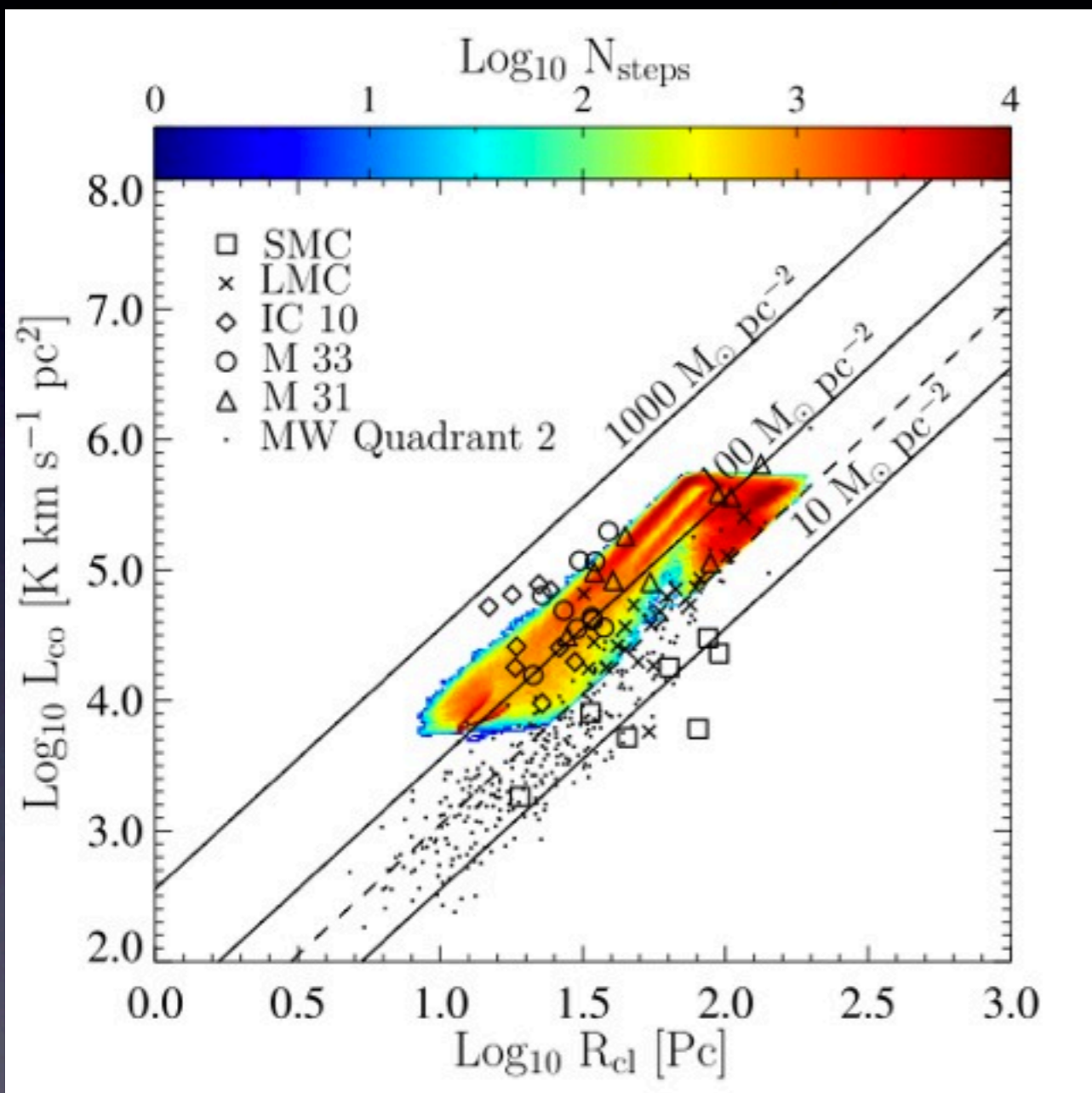
Cloud growth, evolution, and death

$$\Sigma_{\text{res}} = 16 M_{\odot} \text{ pc}^{-2}$$

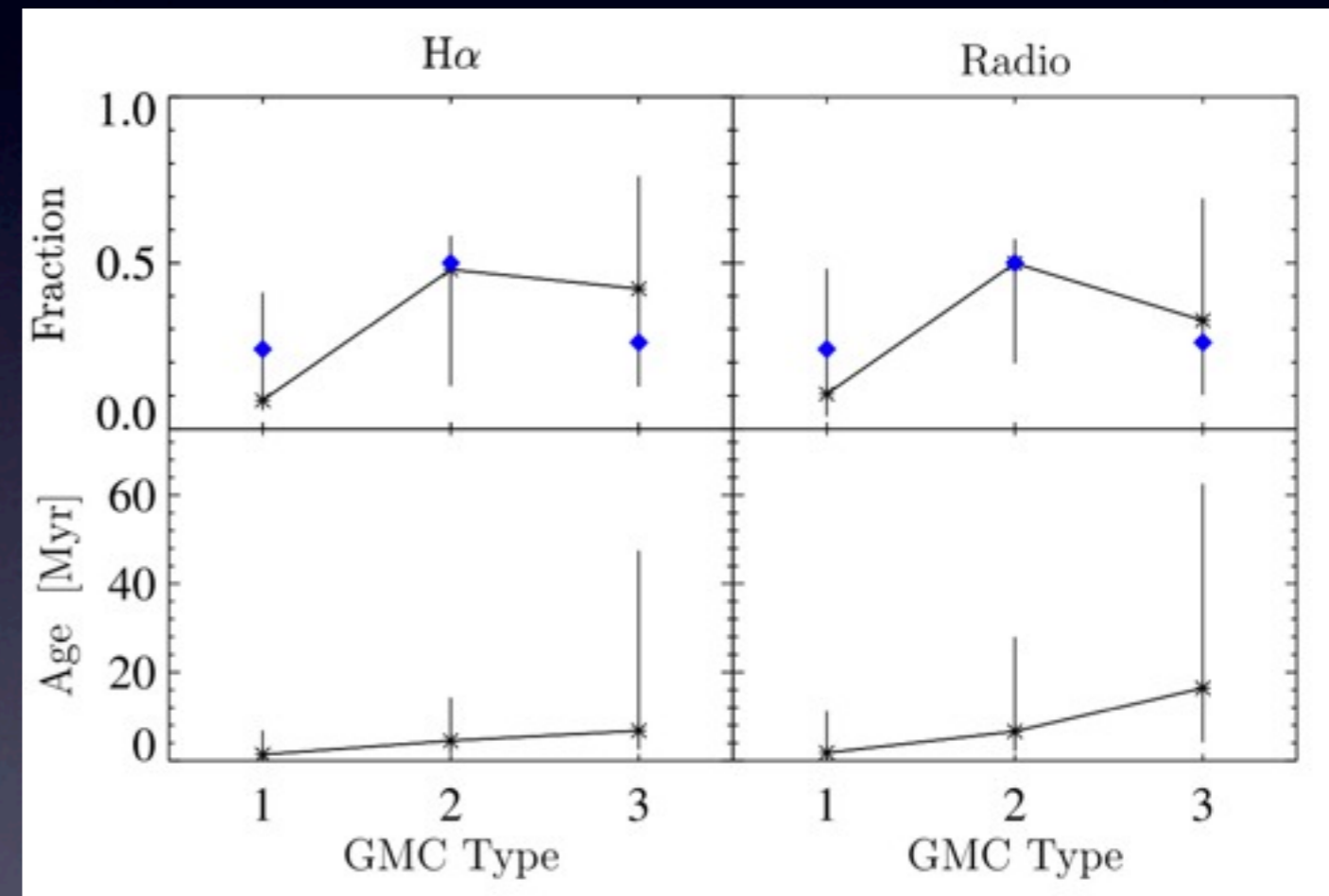
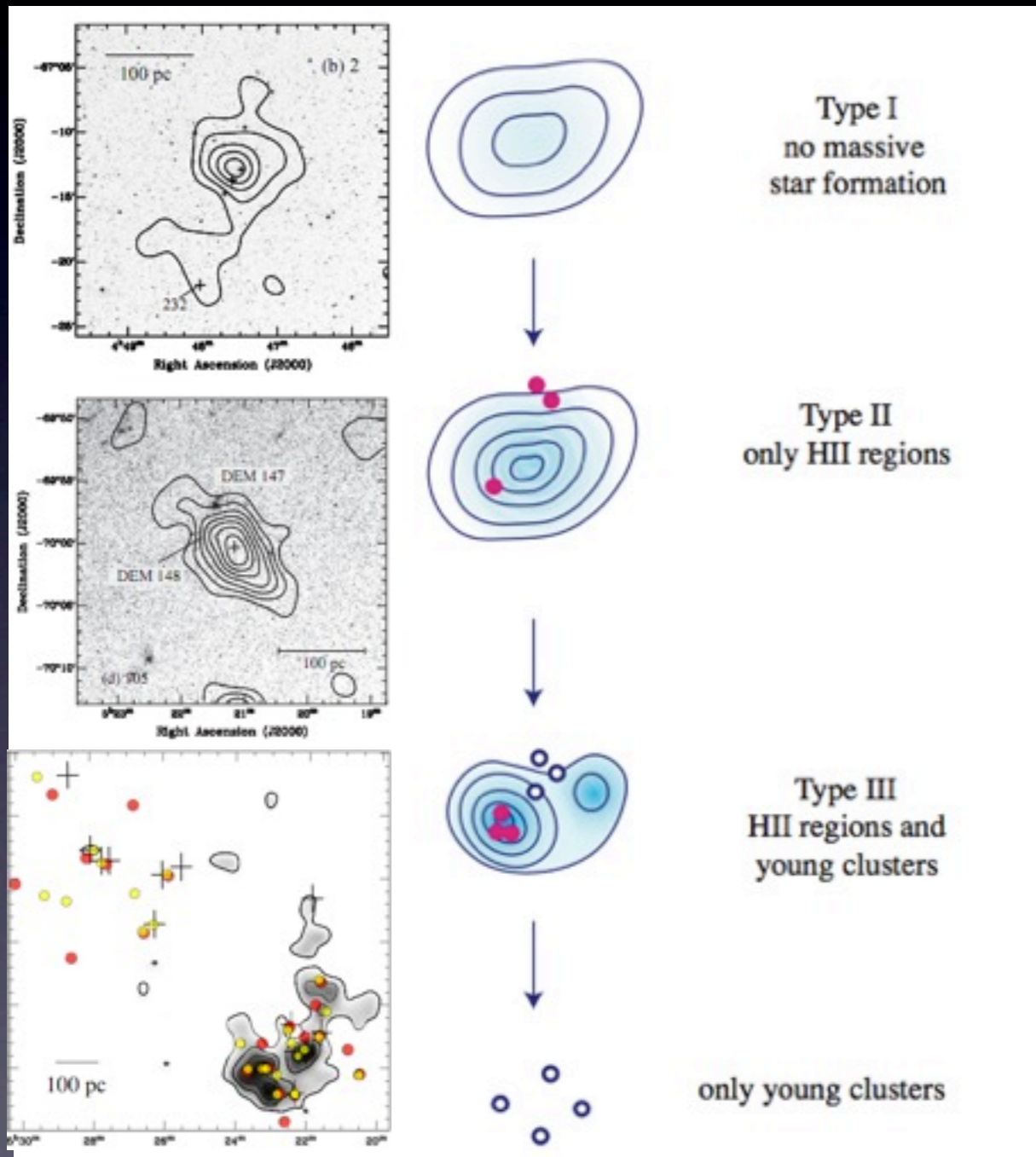


Models run in a fraction of a second

Comparison to Observations: Larson Scaling Laws

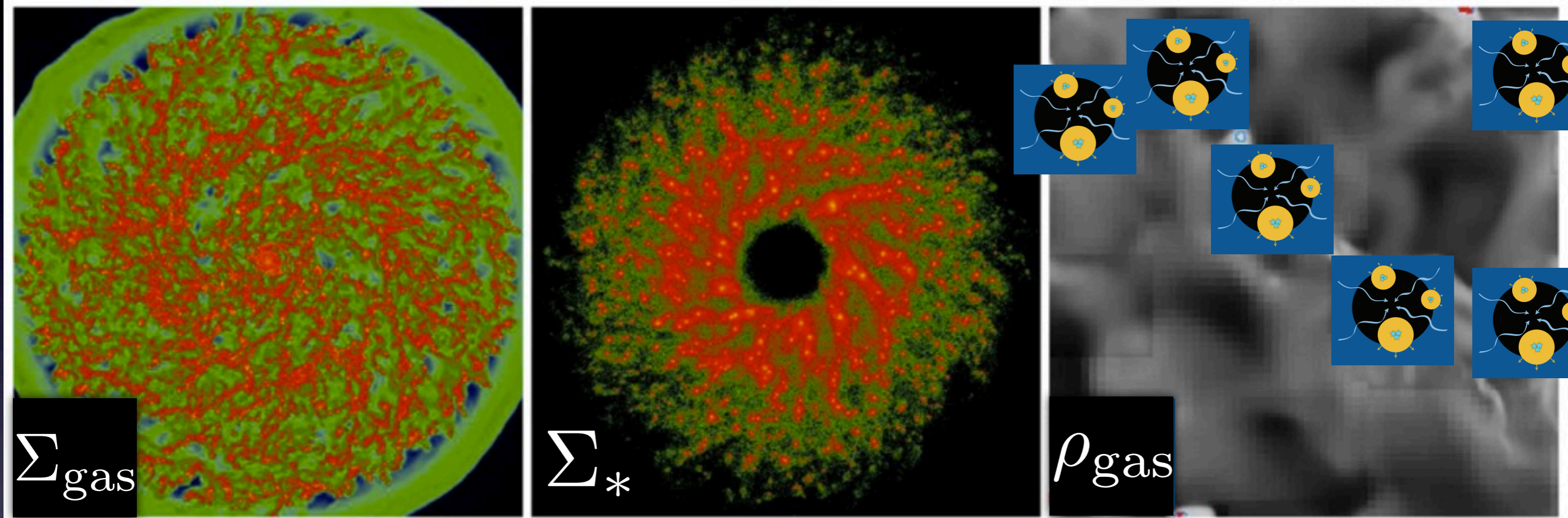


Comparison to Observations: GMC Classification



Kawamura+ 2009

What next?



Self-consistently model the co-evolution of a population of GMCs and a quiescently star-forming galaxy

For more details, see our paper:

THE GLOBAL EVOLUTION OF GIANT MOLECULAR CLOUDS II:
THE ROLE OF ACCRETION

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ApJ in press

<http://arxiv.org/abs/1105.6097>