

Cooling halo and
fragmenting disk

Running Ramses

3D case: Cooling
halo and fragmenting
disk

Toying with the
polytropic pressure
floor

Conclusion

My first steps with Ramses: Cooling Halo and fragmenting disk

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¹ I'm french (BZH), french mother but english father

First impression

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- Easy to compile, just read the Makefile
- Code: Well written and organised in directories.
- Quite strait forward to travel within the code. (usually by grepping in `*/*.f90`)
- Patching is strait forward and easy... (could patch a bug by myself)
- ...if you 're sure to copy the right file into your patch directory ²

² directories mhd/ and hydro/ containing files in with are quite but not totality unlike each other

Test: Sedov 2d case

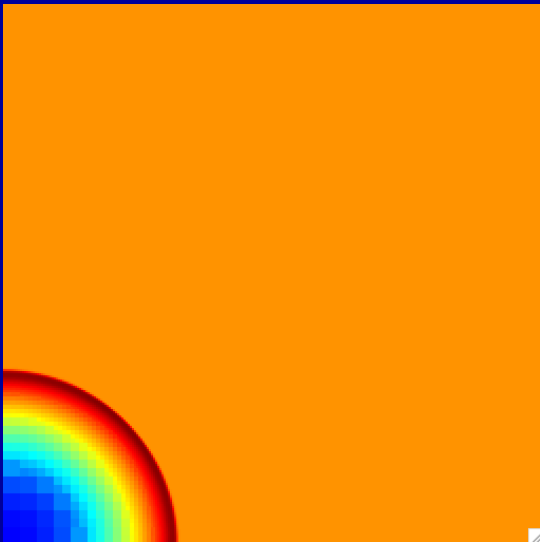
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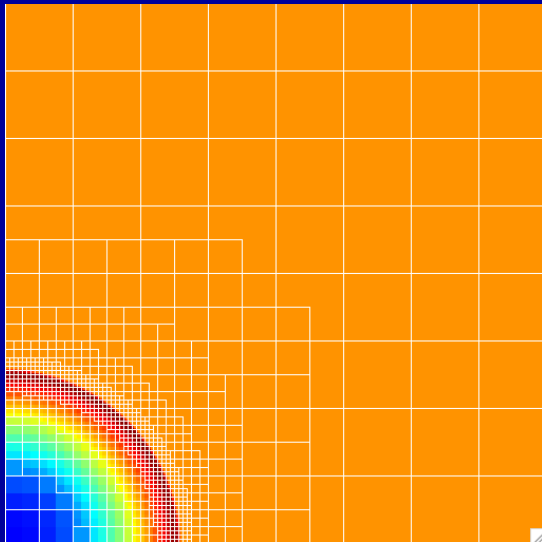
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- isolated halo ($v_{200} = 35 \text{ km/s}$; $c = 10.$; $\text{spin} = 0.1$;
 $f_{\text{gas}} = 0.15$; $B = 10^{-5}$)
- cooling parameters [$T_{2\text{star}}$; n_{star}] = [10^4 , 0, 1] and
default [z_{ave} , g_{star}] = [0., 1.6]
- coarse grid level 7, max refinement 11
- Box size ~ 300 kpc, resolution ~ 0.15 kpc.
- output at ~ 5.74 Gyr

First result

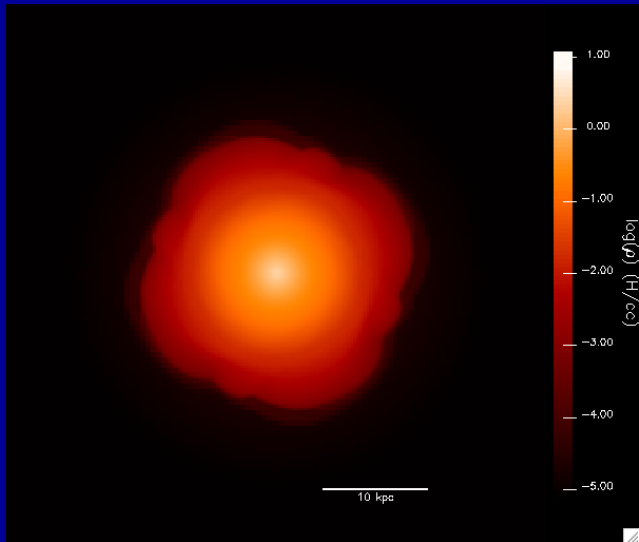
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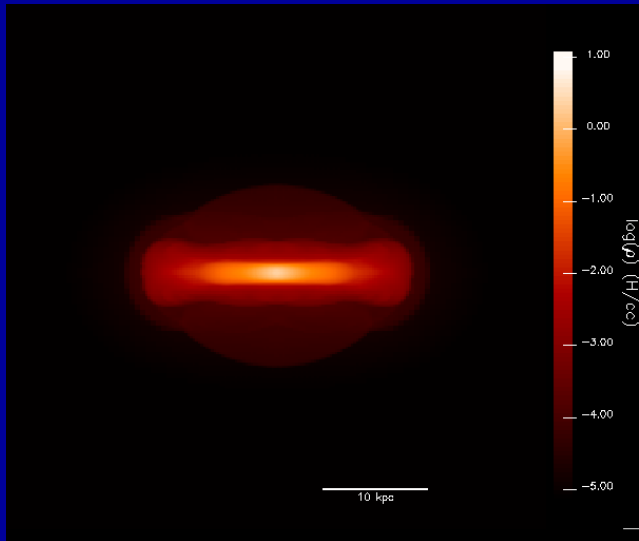
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Adding solar metallicity $z_{\text{ave}} = 1$ and setting $g_{\text{star}} = 2$

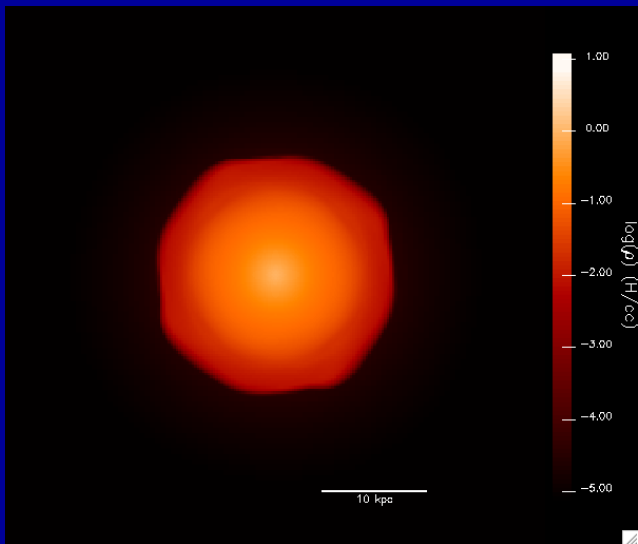
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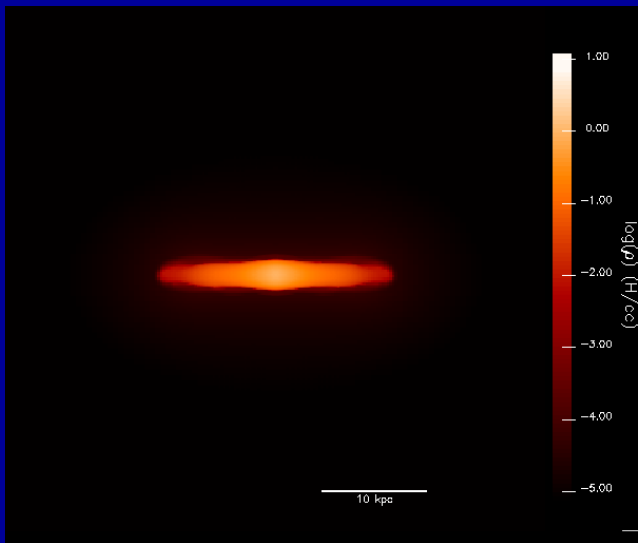
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Net cooling rate for those runs

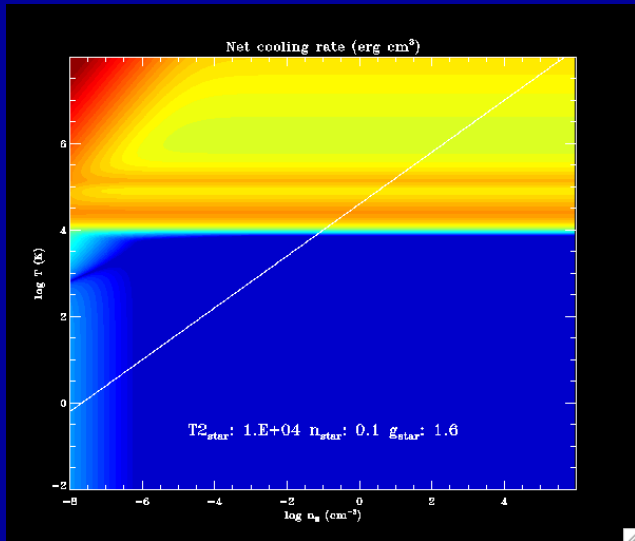
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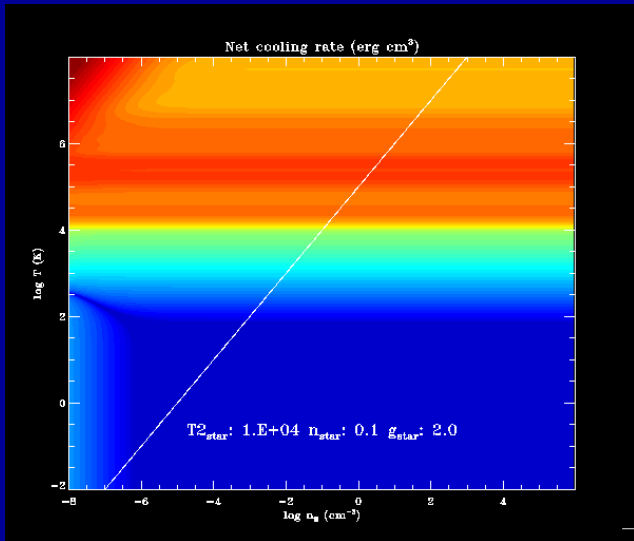
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The polytropic pressure floor

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- Polytropic pressure floor

$$T_{\min} = T_{2\text{star}} * (n_{\text{H}}/n_{\text{star}})^{g_{\text{star}}-1}$$

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$$T_{\min} = T_{2\text{star}} * (n_{\text{H}}/n_{\text{star}})^{g_{\text{star}}-1}$$

- Minimum Jeans Length for gravitational collapse

$$\Lambda_J = c * \tau_{\text{ff}}$$

- with $c = \sqrt{P/\rho} = \sqrt{k_B T/m_H}$ and

$$\tau_{\text{ff}} = \sqrt{\pi/G\rho} = \sqrt{\pi/Gm_H * n_H}$$

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- Well..

$$\Lambda_J = \sqrt{k_B \pi / G m_H^2} * \sqrt{T_{\min} / n_H}$$

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$$\text{with } g_{\text{star}} = 2 \text{ we get } \Lambda_J = \sqrt{k_B \pi / G m_H^2} * \sqrt{T_{2\text{star}} / n_{\text{star}}}$$

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- To prevent artificial fragmentation : $\Lambda_J > 4\Delta x$

Decreasing the Pressure floor

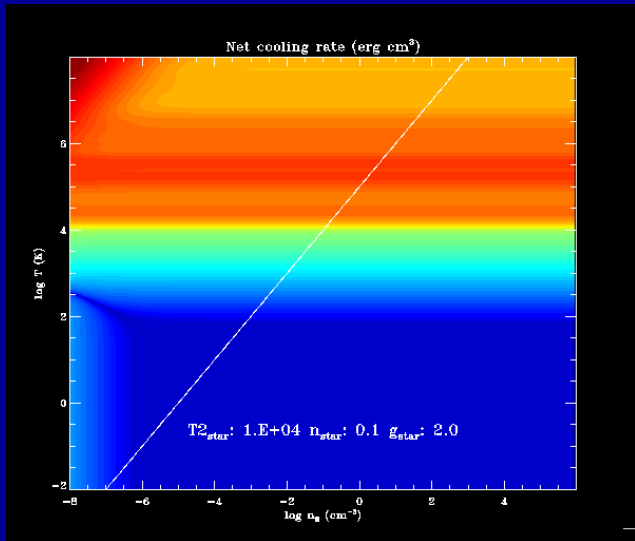
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Decreasing the Pressure floor

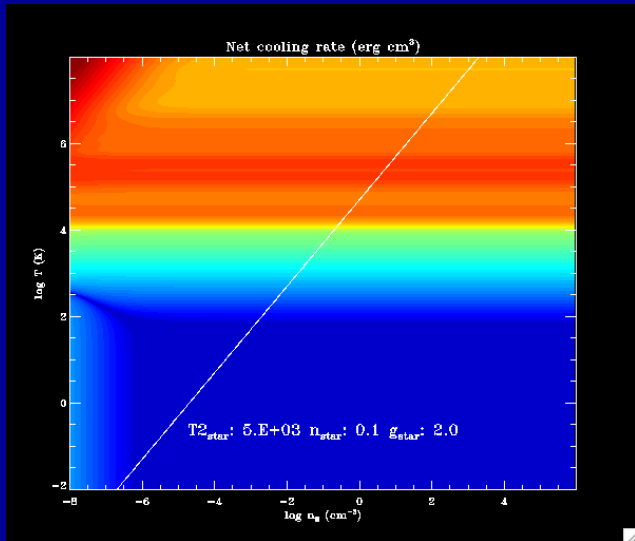
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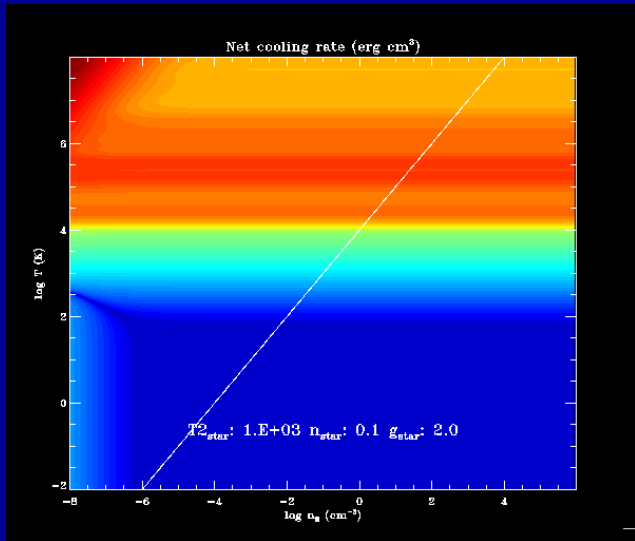
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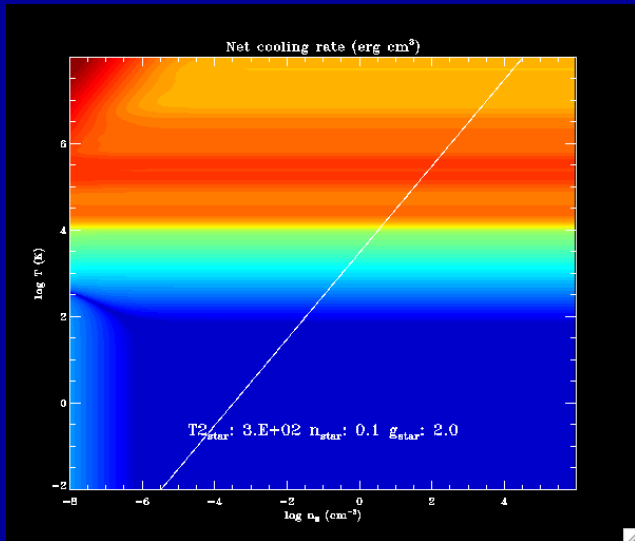
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$$T_{2\text{star}} = 10000, n_{\text{star}} = 0.1, \Lambda_J = 33.98\Delta x$$

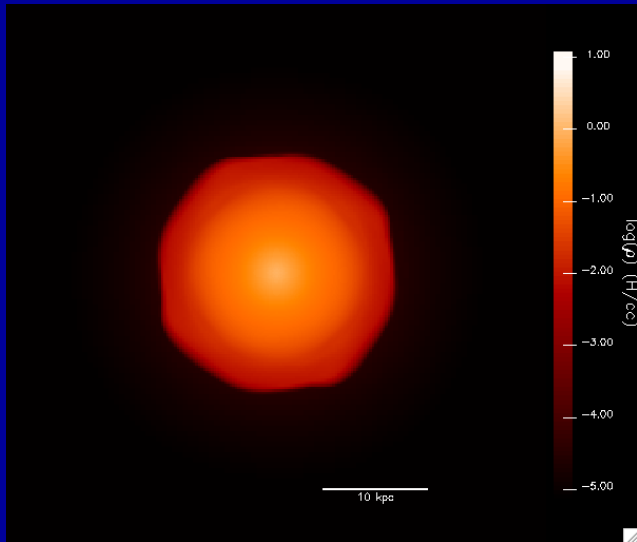
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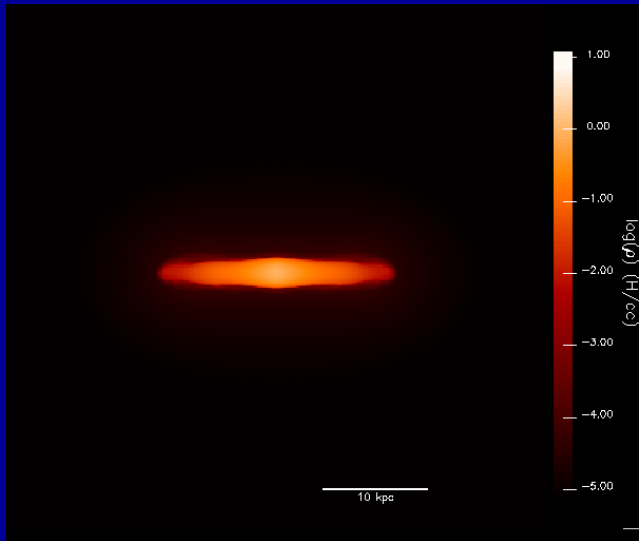
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$$T_{2\text{star}} = 5000, n_{\text{star}} = 0.1, \Lambda_J = 24.03\Delta x$$

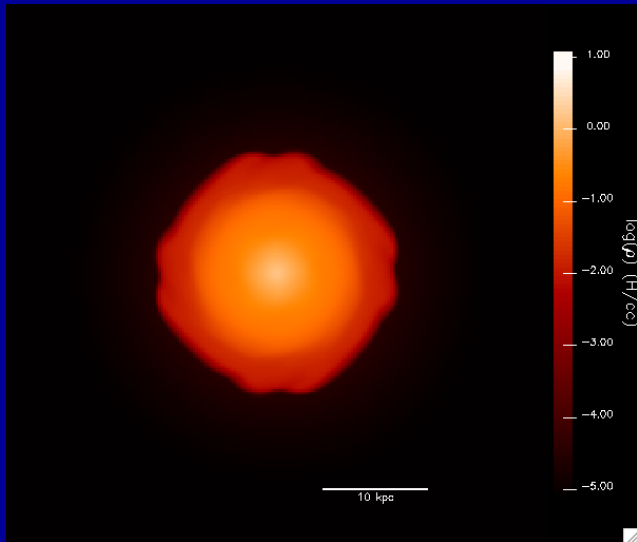
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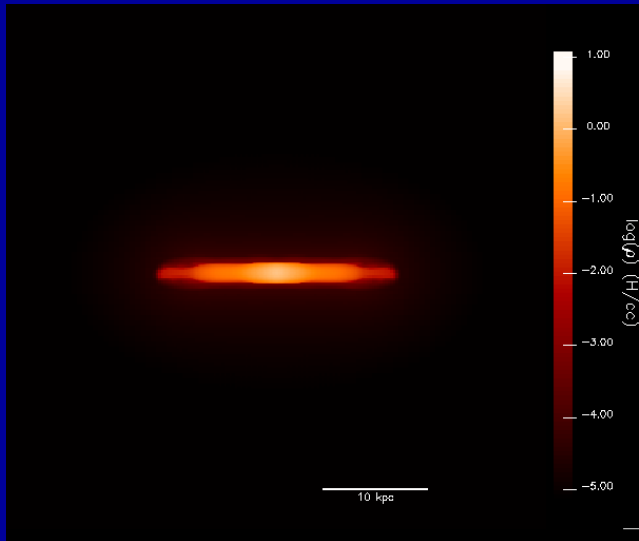
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$$T_{2\text{star}} = 1000, n_{\text{star}} = 0.1, \Lambda_J = 10.75\Delta x$$

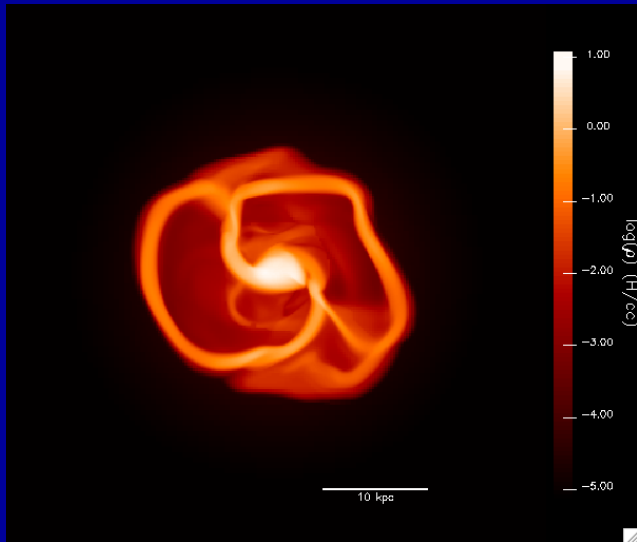
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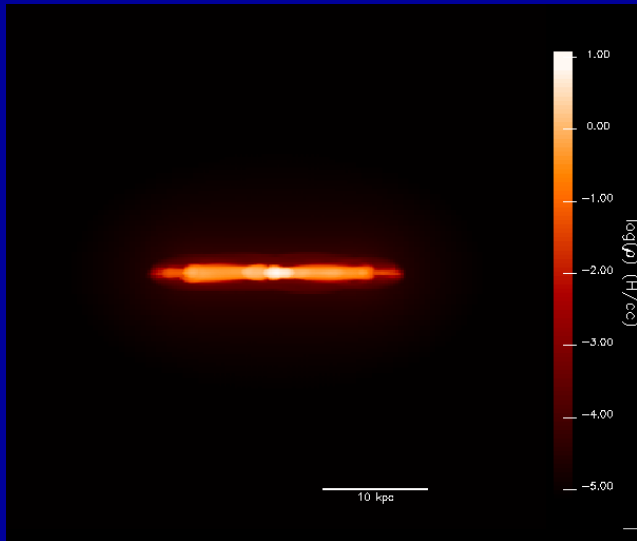
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$$T_{2\text{star}} = 300, n_{\text{star}} = 0.1, \Lambda_J = 5.89\Delta x$$

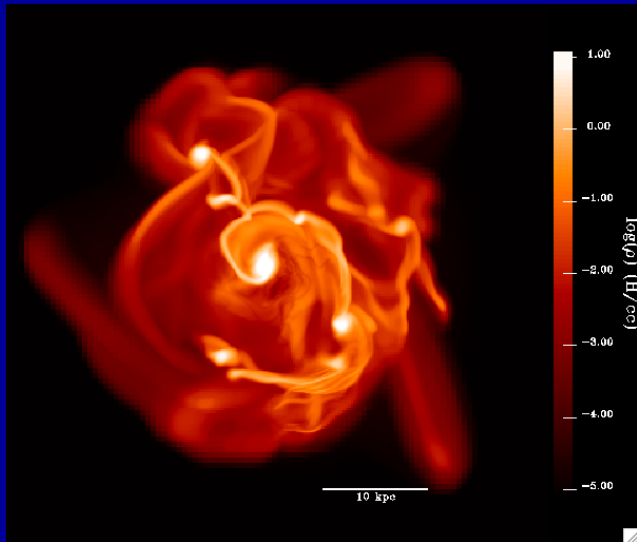
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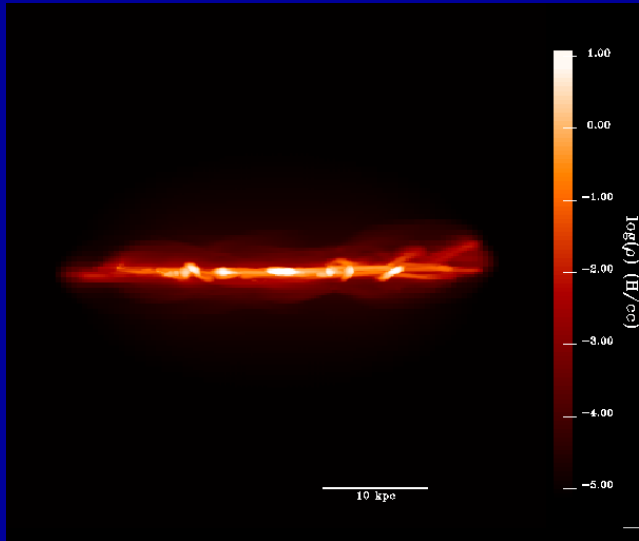
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- Learned how to send a job using PBS script.
- got much much familiar with Ramses
- got an idea, of the work involved in getting accurate physics from a simulation
- created my own visualisation routines from the one provided. (learned new good tricks with idl)
- was happy to share the little experience I had with others running the same project.